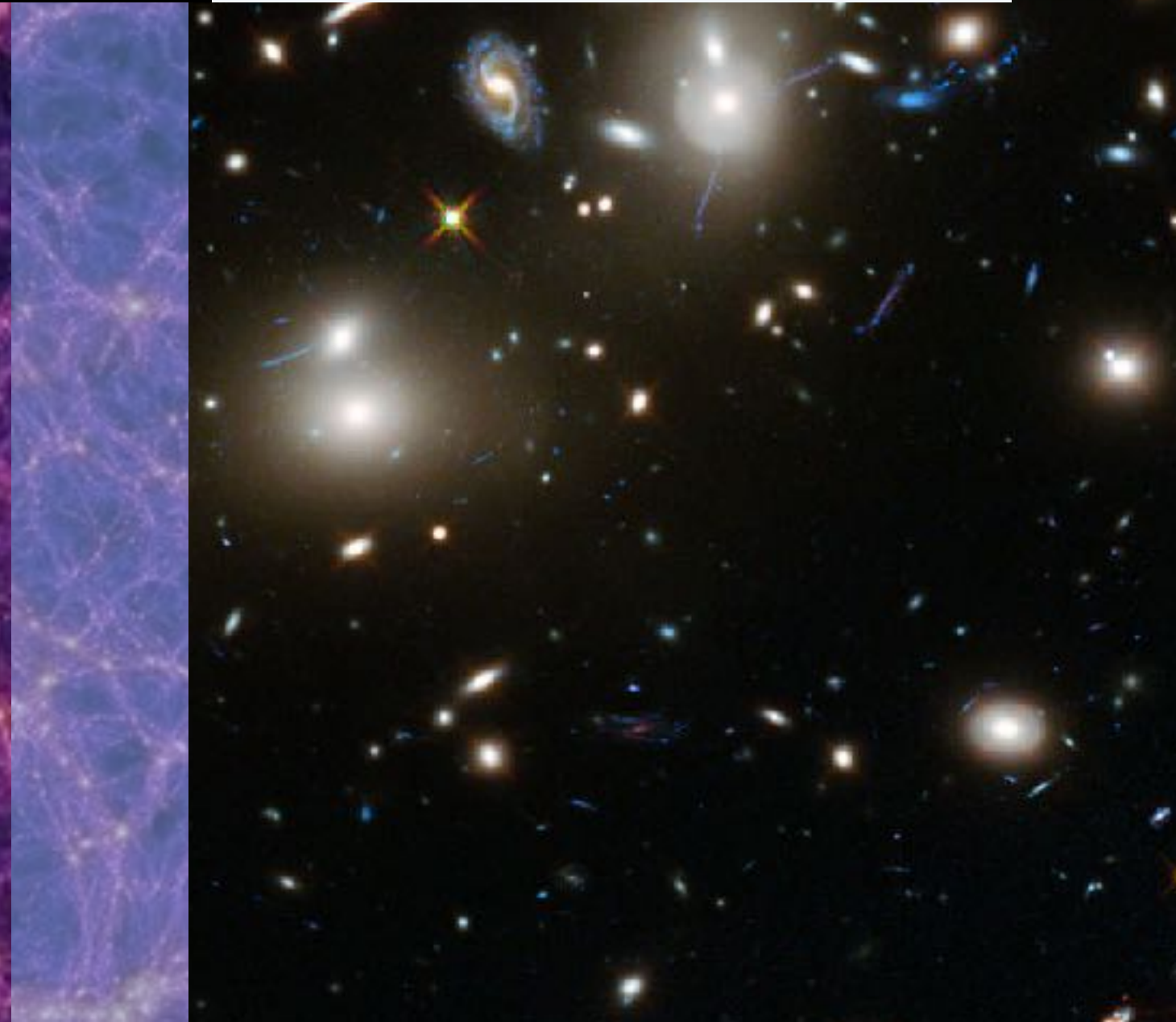


From Effective Field Theory to **lensing**

ICG & University of Portsmouth

Lucia F. de la Bella — 18/05/2021





Lucia F. de la Bella

- I'm Spanish
- I have a twin sister
- Craft hobbies
- Love nature & hiking
- Gospel Choir



B.Sc
Quantum
Cosmology

2008-2013

M.Sc
EFToDE

2013-2014

PhD
EFToLSS
RSDs
Halo bias

2014-2018

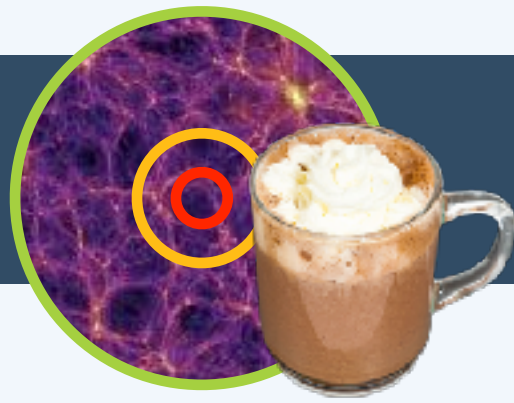
Teaching

2018-2019

2019-2021

PDRA
Weak Lensing
Unequal-time
correlators





Effective Field Theory of Large Scale Structure

de la Bella et al. 2017
de la Bella et al. 2018

Linear regime SPT works.
Mild non-linear regime SPT breaks down.
Non-linear regime unknown description.

$$P_{EFT} = P_{SPT} - c^2 k^2 P_{11}$$

counterterm

Traditionally

- One-loop power spectrum in standard perturbation theory
- Einstein-de Sitter approximation for growth functions
- Divergences from small-scale physics

New

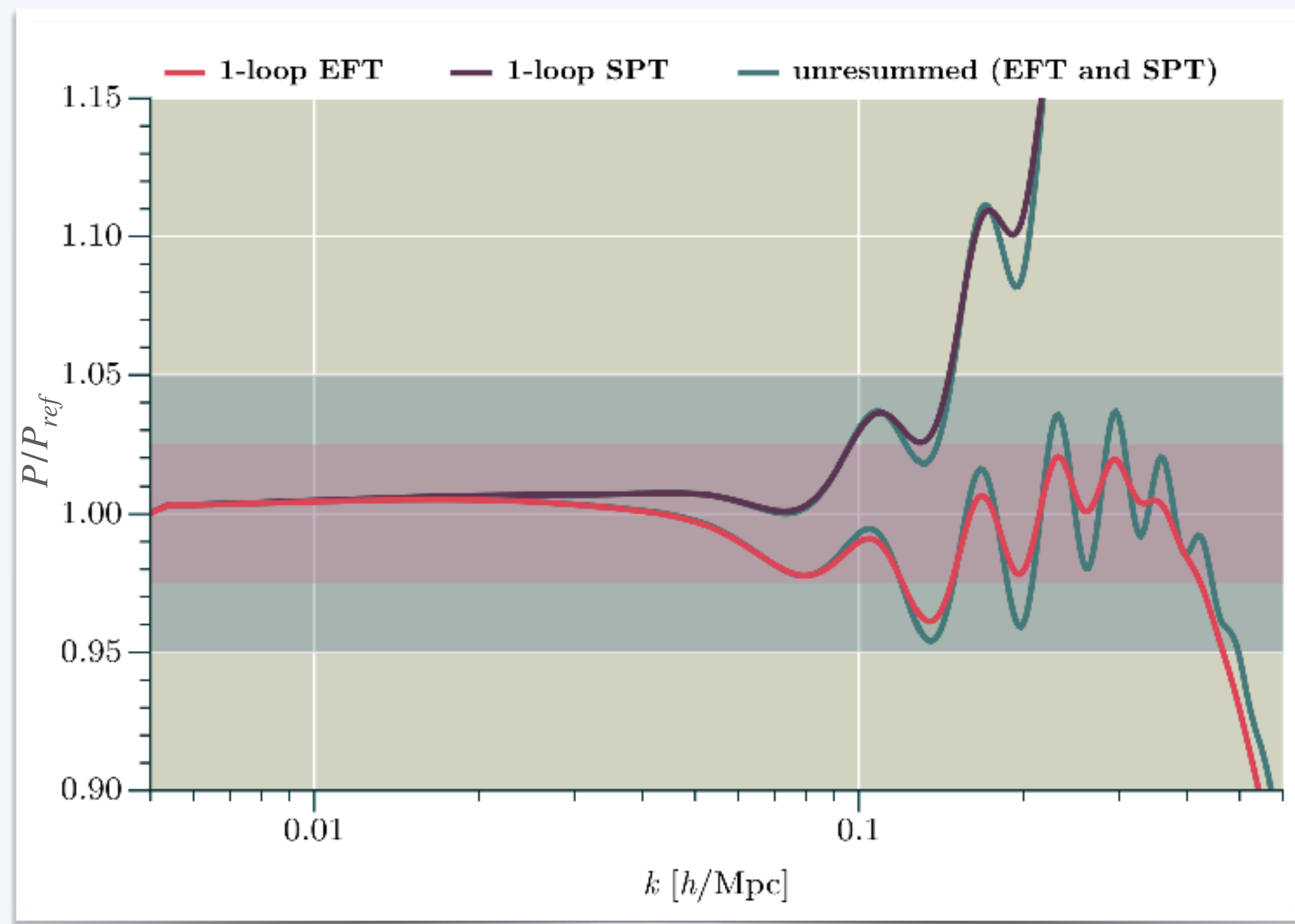
- One-loop power spectrum in redshift space
- Most general bias model
- Full time dependence non-linear growth functions
- Novel analytical methods:
 - Split tensor and scalar loop integrals
 - Fabrikant's procedure for Bessel functions
- Counterterms:
 - Uniform notation and language
- Applied the Vlah et al. IR-resummation scheme in redshift space.

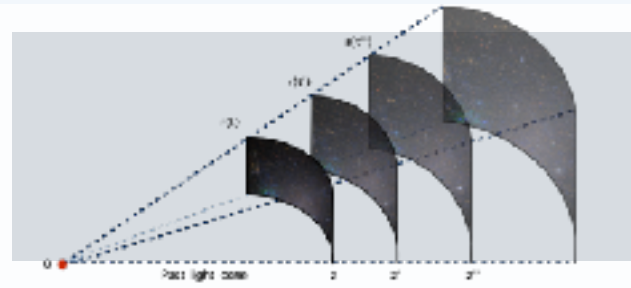
Application

- Redshift-space distortions
- Halo bias and WIZCOLA sims

Impact

- Accuracy on smaller scales!





Unequal-time correlators

de la Bella et al. (2020)
Tessore, de la Bella (in prep)

Correlations between fields

- Same time slice: *equal-time correlators*
- Different time slices: *unequal-time correlators*

$$w(\theta) = \iint_0^\infty dx_1 dx_2 f_1(x_1) f_2(x_2) \xi(r_{12}; t_1, t_2)$$

$$C(\ell) = \int_0^\infty \frac{dk}{k^2} \iint_0^\infty dx_1 dx_2 f_1(x_1) f_2(x_2) j_\ell(kx_1) j_\ell(kx_2) P(k; t_1, t_2)$$

Traditionally

- Limber's approximation and equal-time power spectrum
- Only good for linear theory and small-angle separations

New

- **Exact** angular power spectrum at all angular scales
- Flexibility to include equal and **unequal-time** correlators
- Unequal-time correlator:
 - One-loop SPT and EFToLSS
 - First fitting formula for counterterms
 - New midpoint approximation
 - Open-source software: **unequalpy**
- Exact angular correlations:
 - FFTLog methods
 - Avoid highly-oscillatory Bessel functions.
 - Open-source software: **corfu**

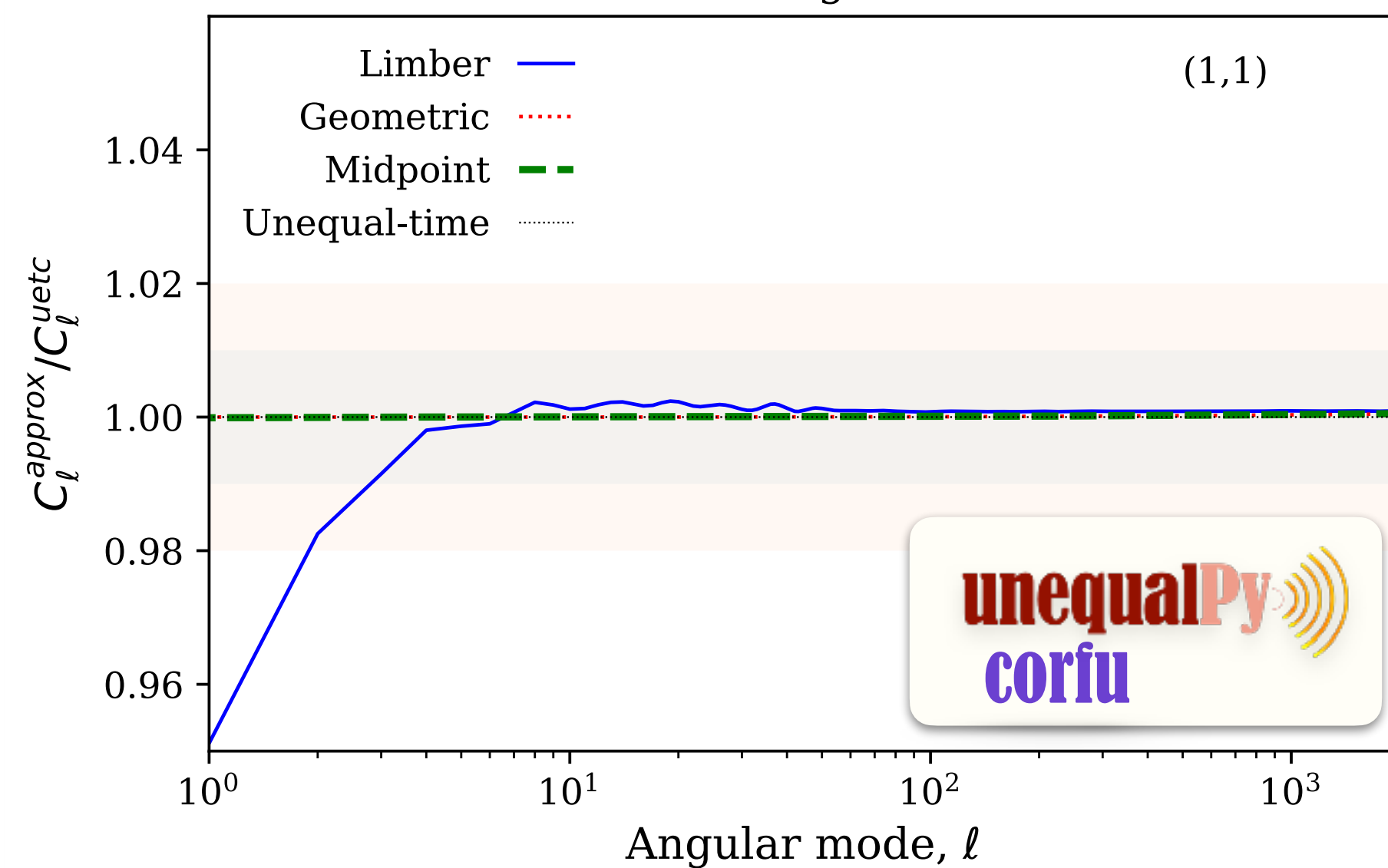
Application

- Cosmic convergence
- Galaxy clustering
- Galaxy-galaxy lensing
- DES Y1 & LSST Y10

Impact

- Drop Limber's approximation
- Relevant even for DES Y1

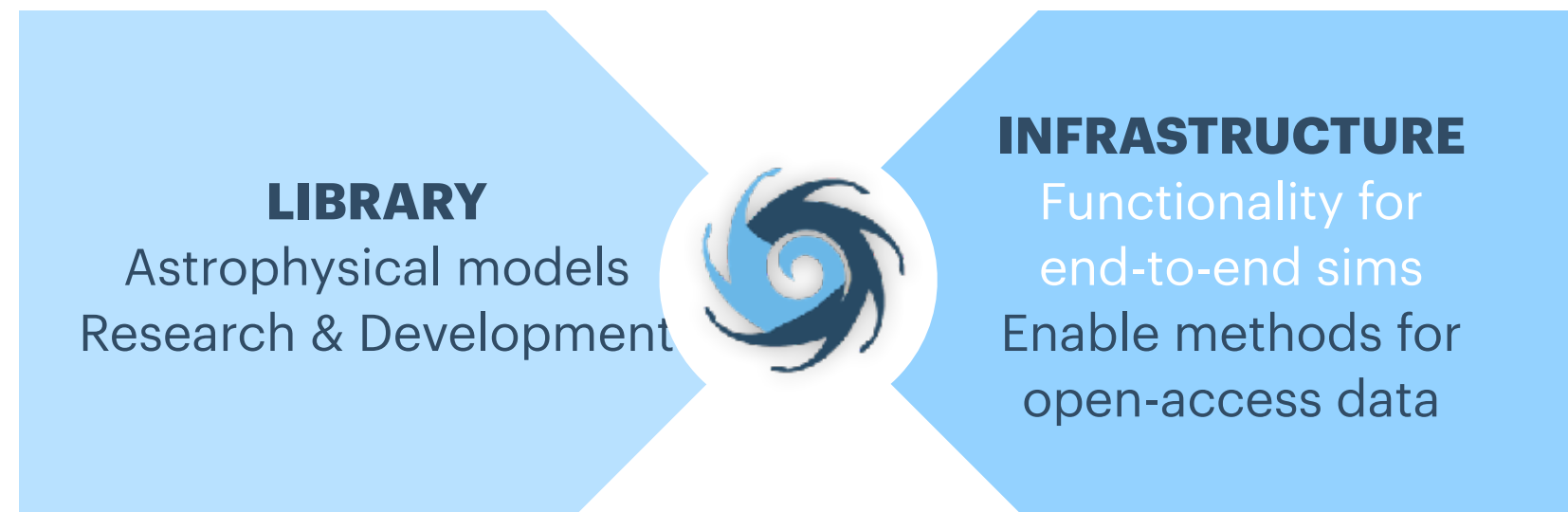
Convergence



COMMUNITY PACKAGE



- **Open-source** off-project
- High-quality **Python** package



- **GitHub** organisation
- Unit tests & high-quality documentation
- Code review & **Infrastructure** team

```

pip install skypy or
conda install -c conda-forge skypy or
git clone https://github.com/skypyproject/skypy.git
import skypy
    
```

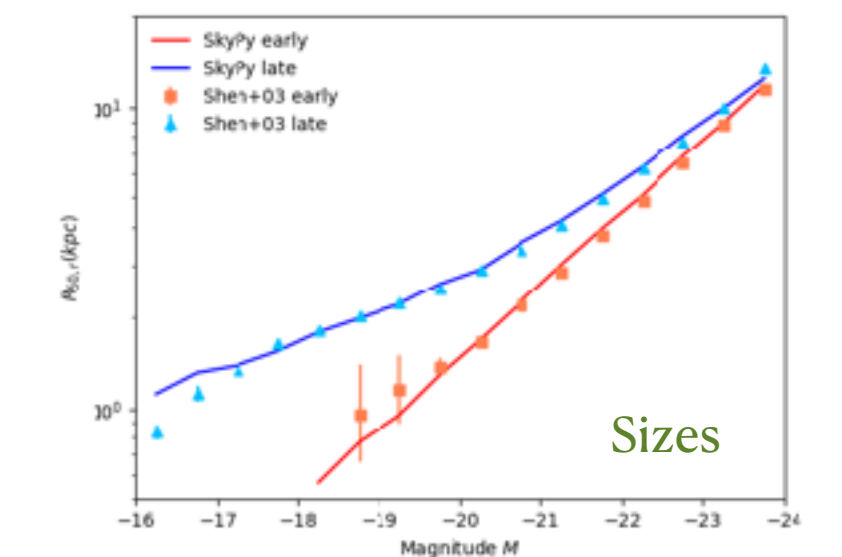
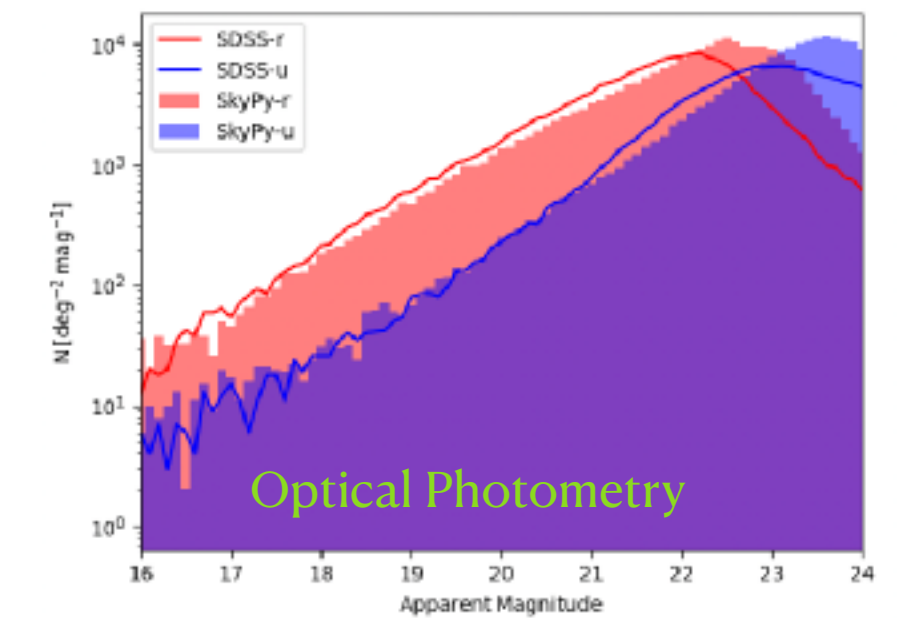
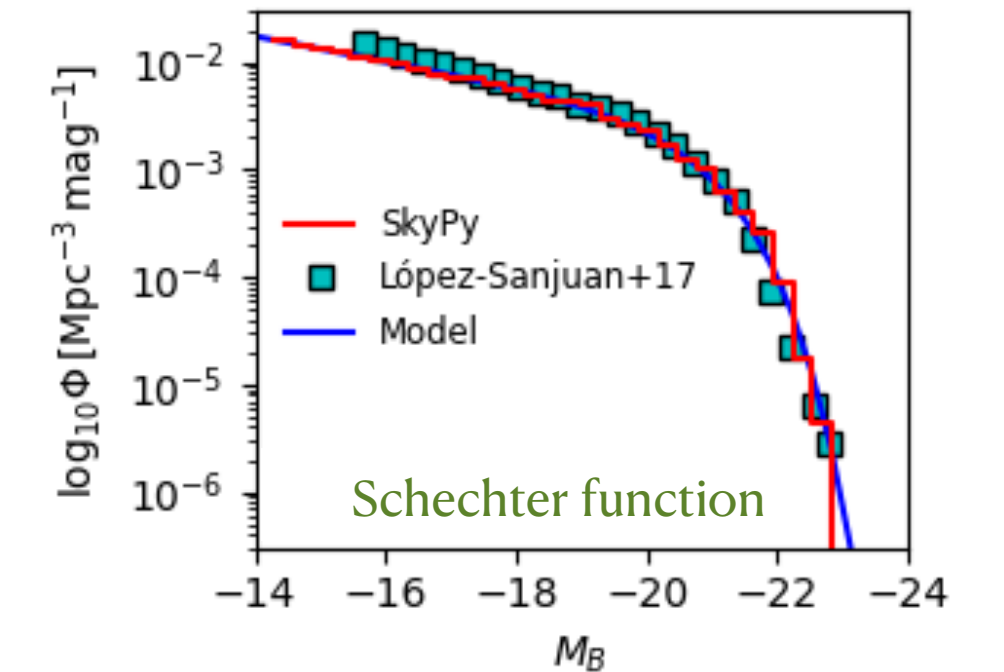
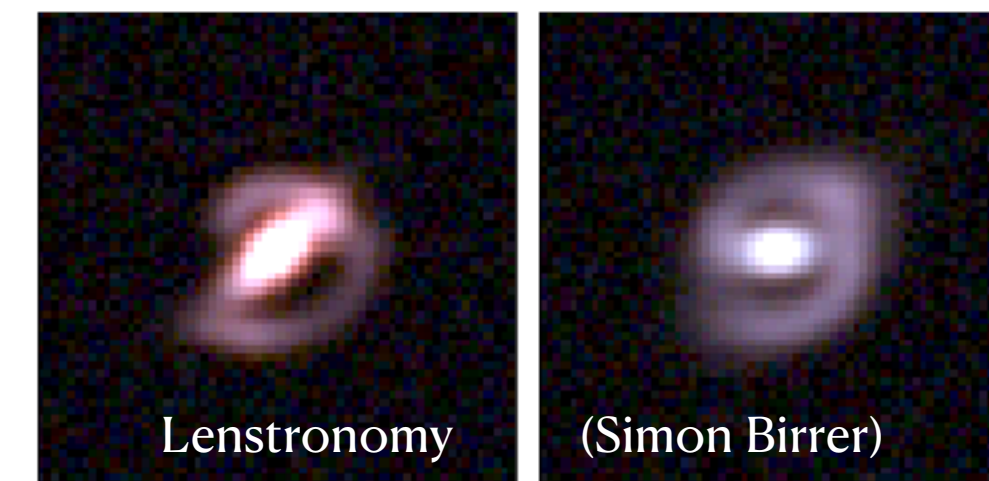
<https://github.com/skypyproject/skypy.git>

<https://skypyproject.org>

SIMULATION PIPELINES

- YAML-based config files
- The **SkyPy Driver** runs end-to-end **pipelines**
- **Total flexibility!**

- SkyPy Pipeline
- **KEY:** you can write your own **pipelines!**



<https://skypy.readthedocs.io/en/latest/examples/index.html>

Work in progress...



Idea

- **Screening** phenomena
- Merge **EFToLSS** and **EFToDE**
- Make use of counterterms

Traditionally

- Linear power spectrum for dark energy and modified gravity theories

New

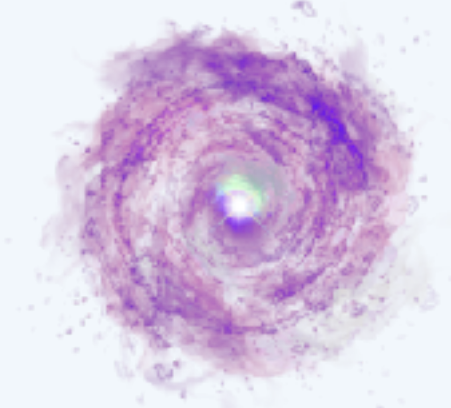
- First **library** for non-linear power spectrum within EFToDE
 - Parametrisation of screening
 - Determination of screening scale
 - Open-source software: **eftPy**

Application

- Covariance Galileons
- Simulations

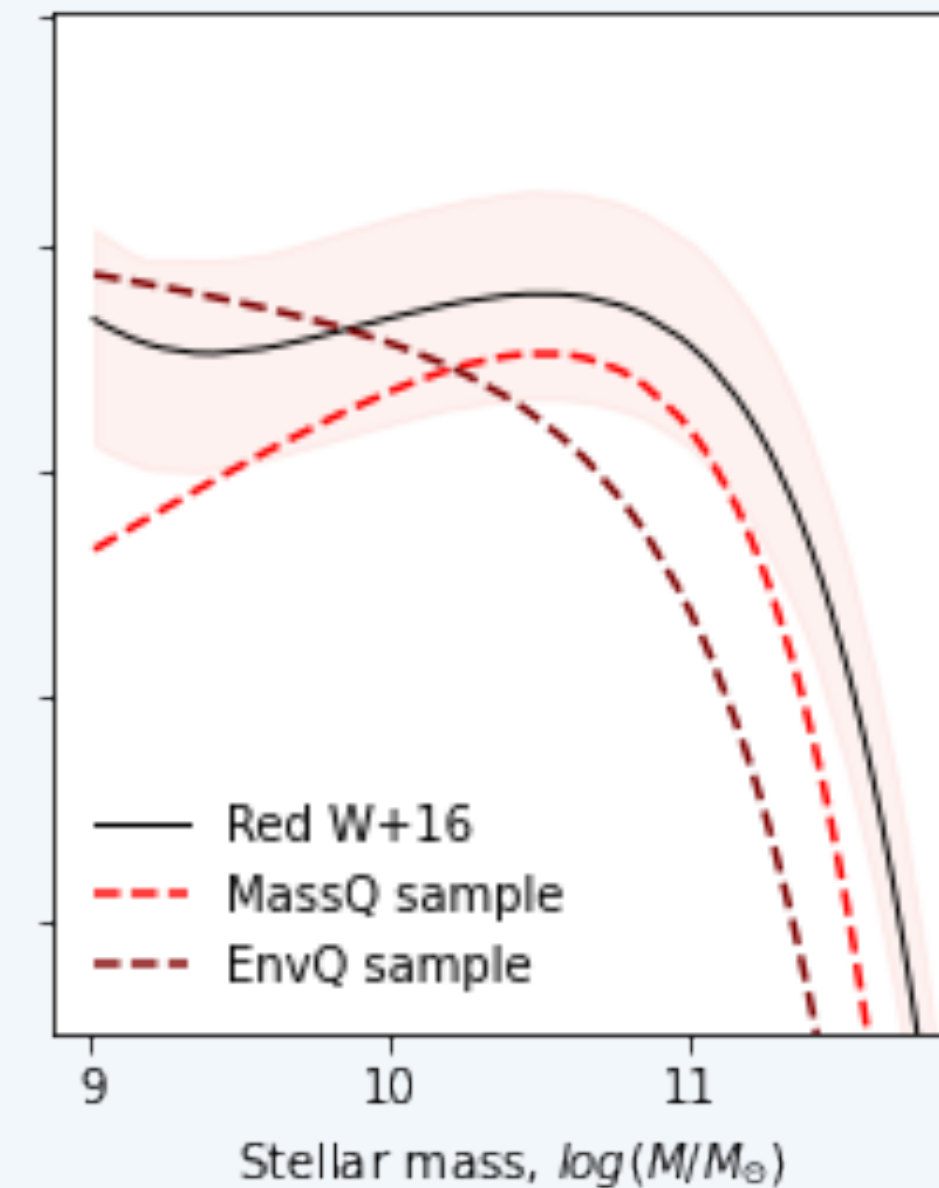
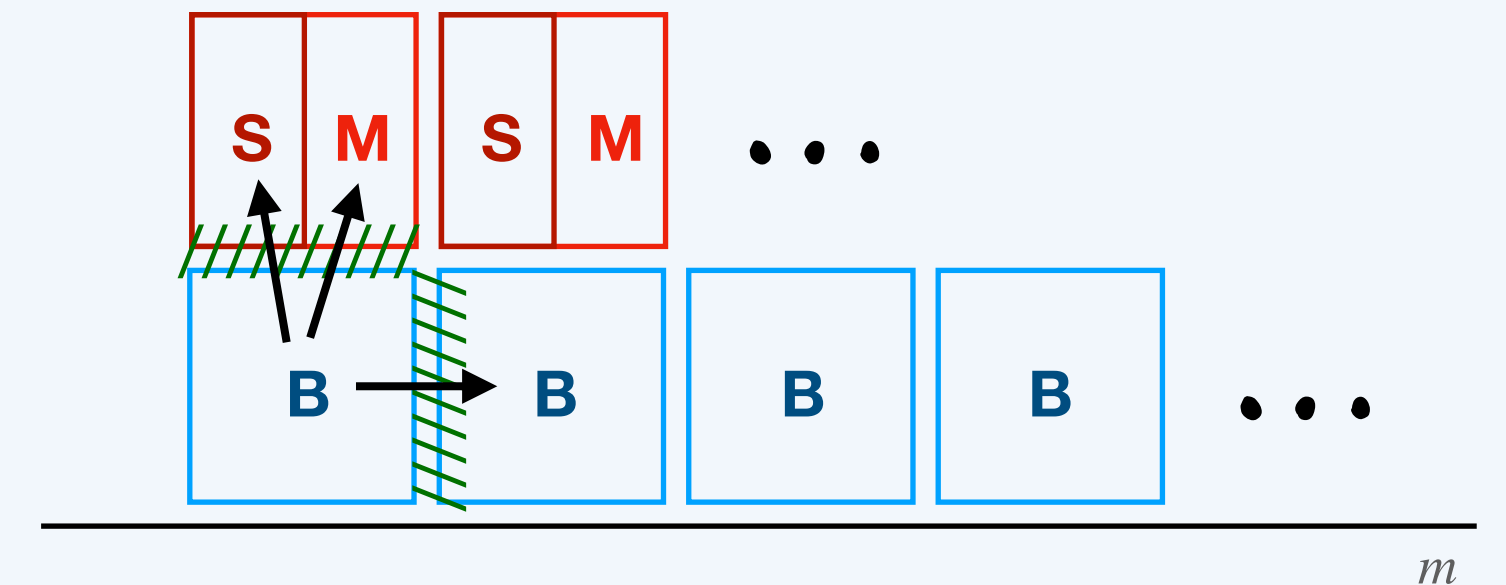
Impact

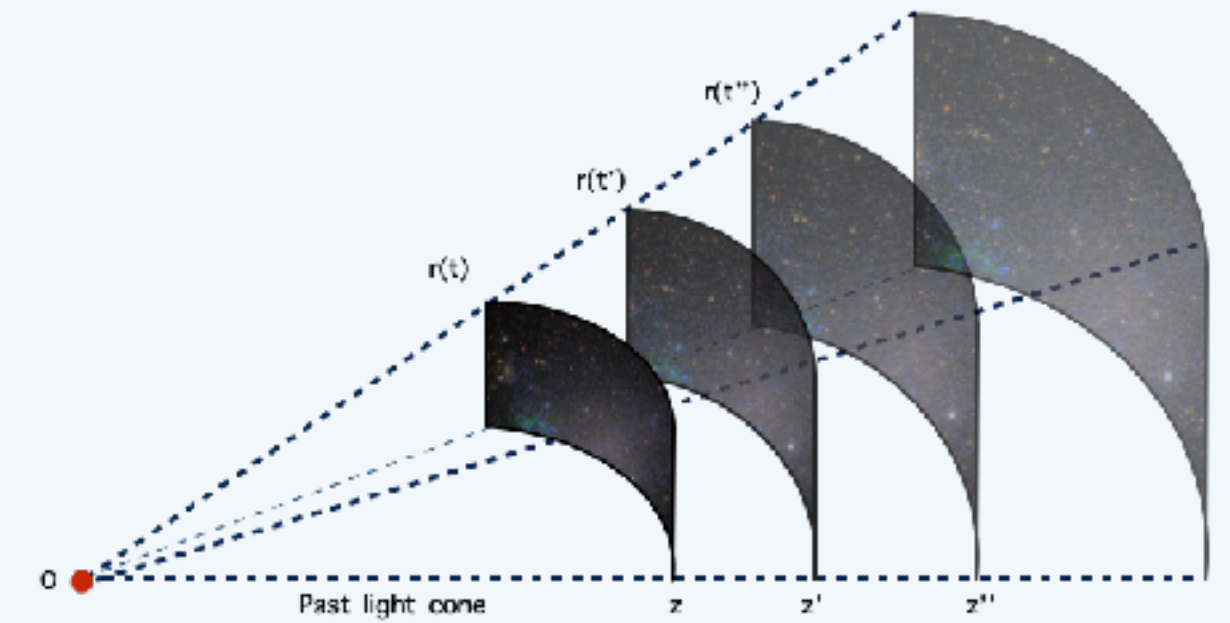
- Crucial to truly compare with observations



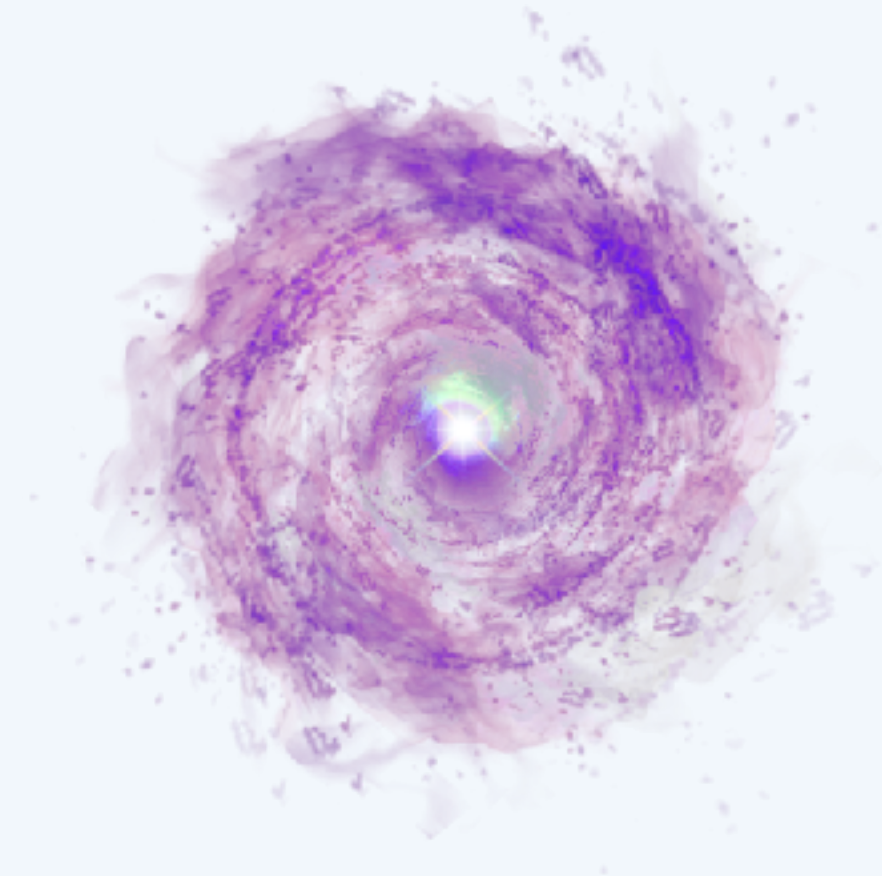
- Star-forming galaxies $\{\phi_b^*, \alpha_b, m_b^*\}$
- Satellite-quenched g. $\{\phi_\rho^*, \alpha_\rho, m_\rho^*\}$
- Mass-quenched g. $\{\phi_m^*, \alpha_m, m_m^*\}$

Quenched galaxies





**Interested in these topics?
Happy to chat!**

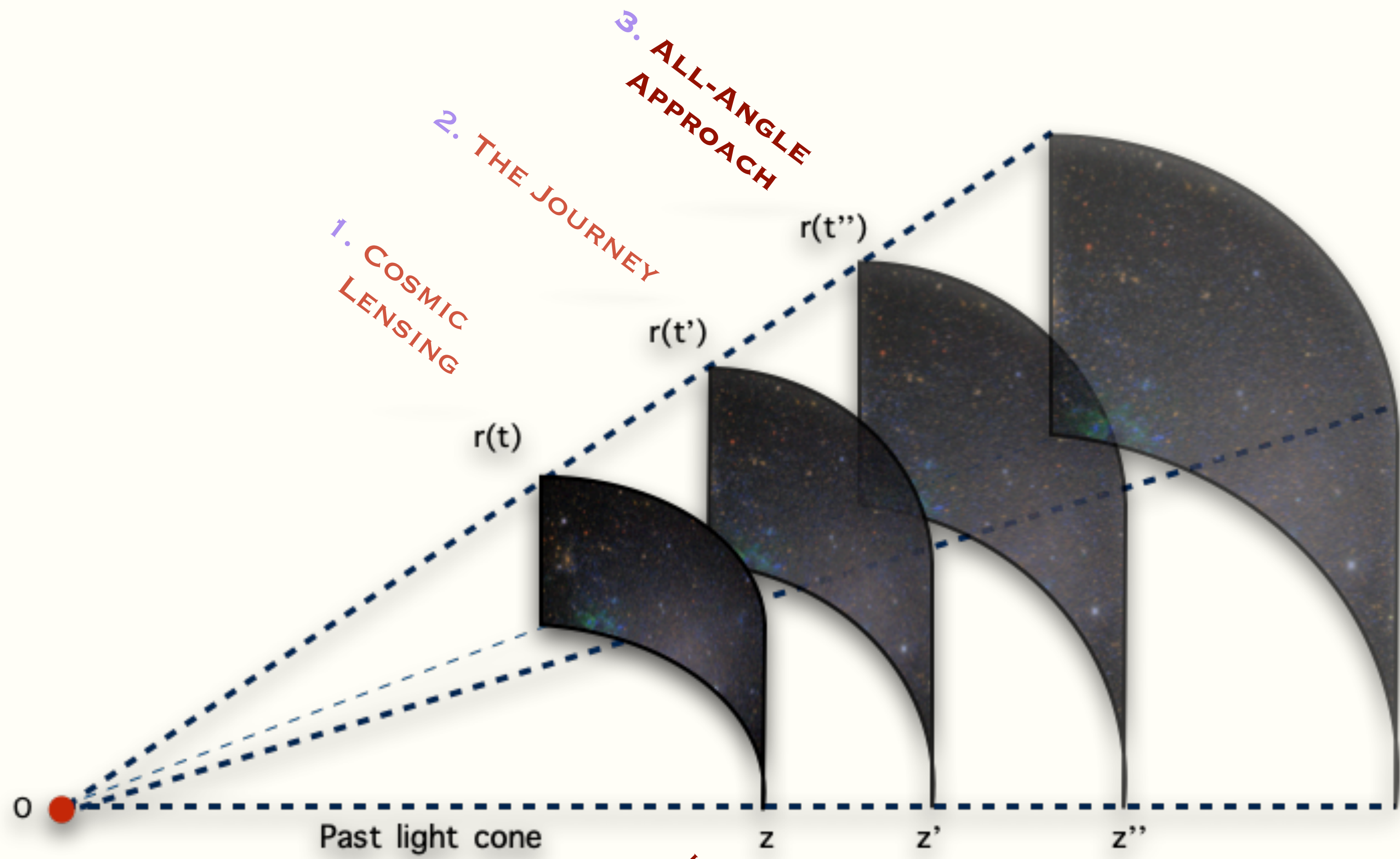


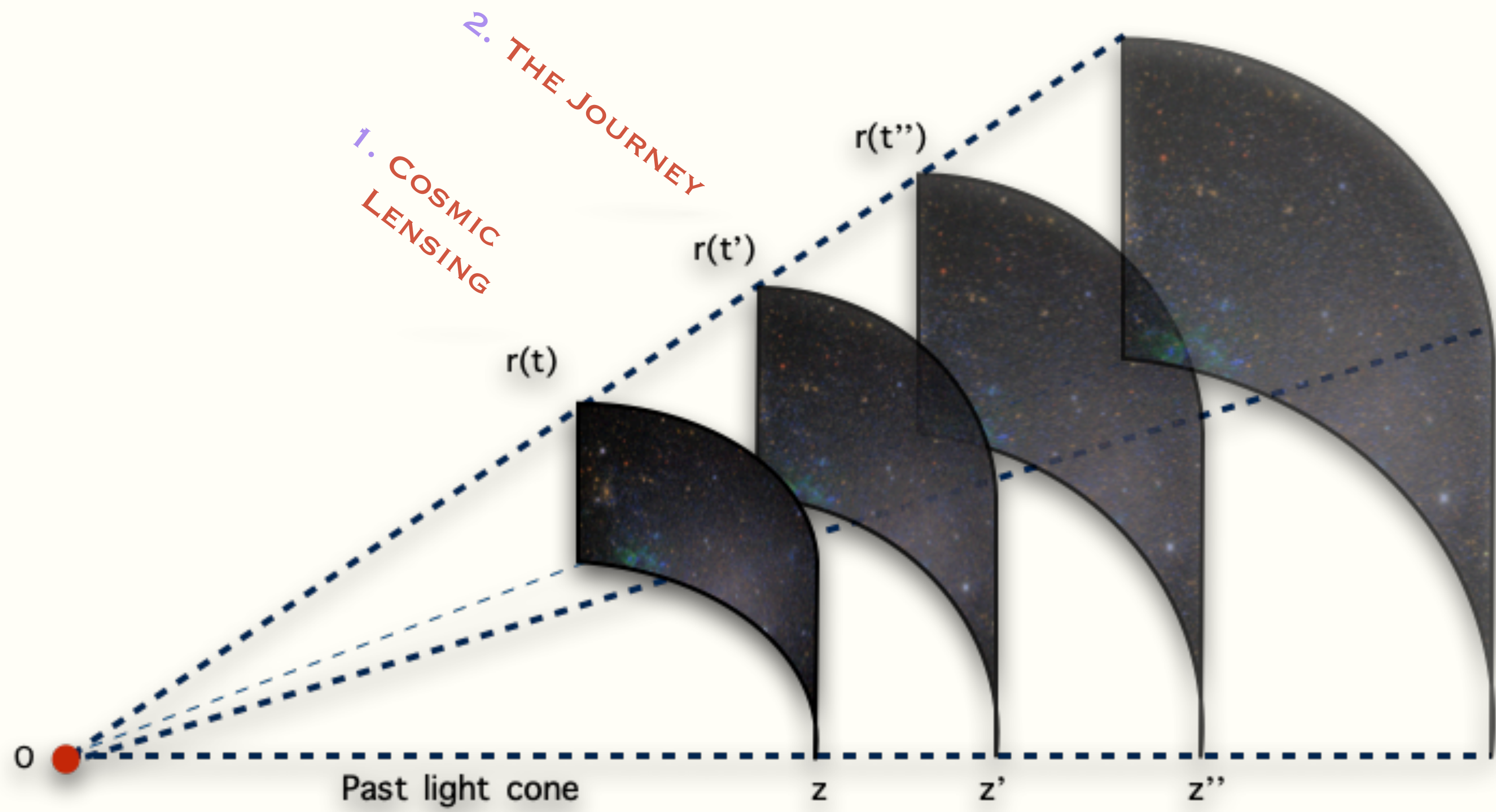


THE JOURNEY OF LIMBER

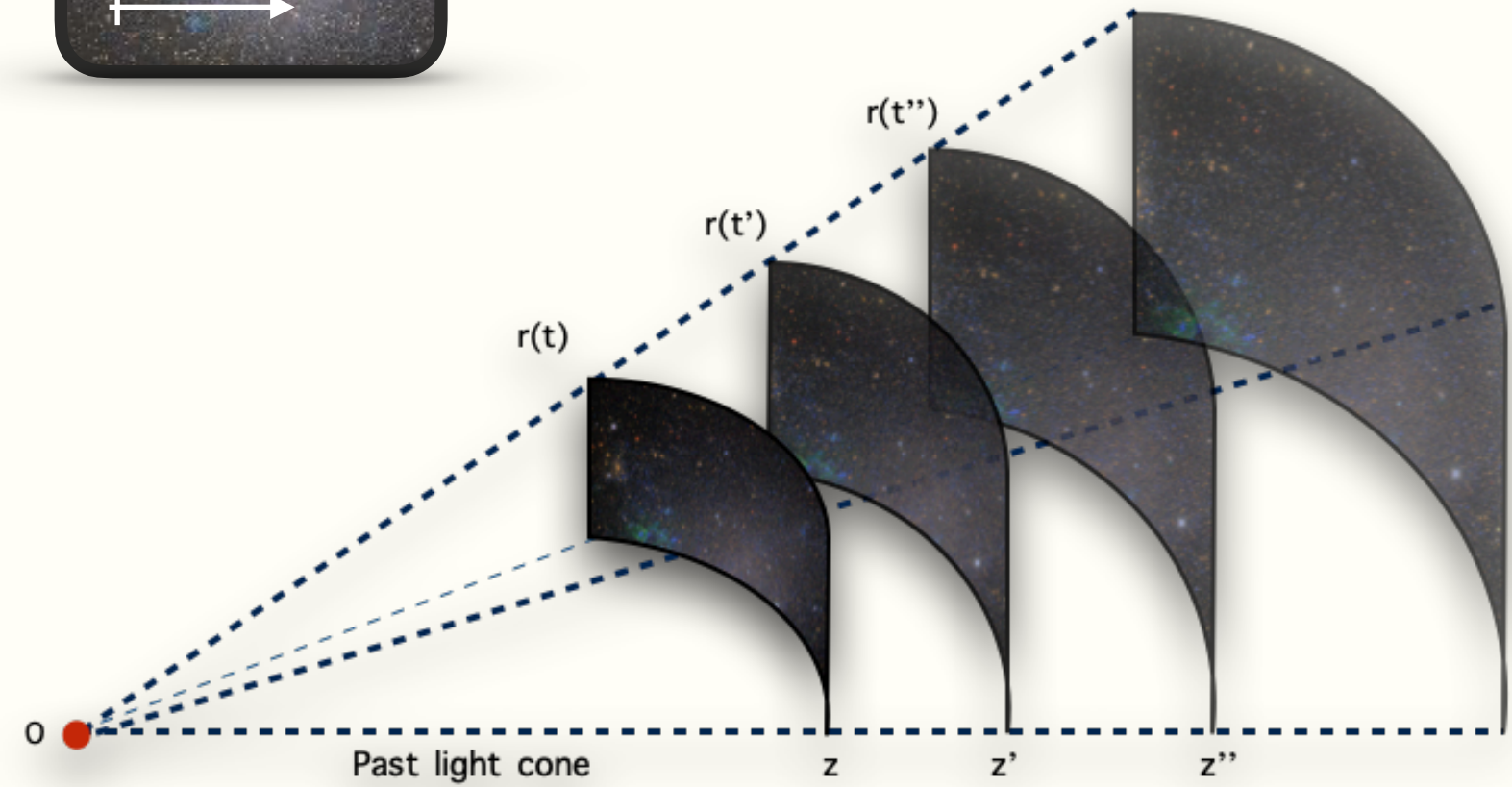
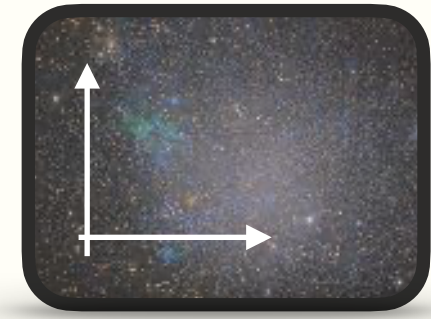
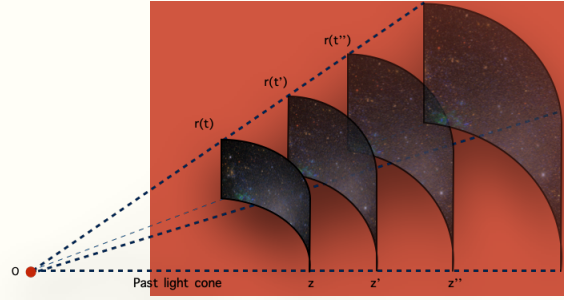
L.F. de la Bella, N. Tessore and S. Bridle (arXiv 2011.06185)
N. Tessore and L. F. de la Bella (in prep)

Past Python packages **UNEQUALPY**
CORFU





1. Cosmic Lensing



Correlations between fields

- Same time slice: *equal-time correlators*
- Different time slices: *unequal-time correlators*

Examples of fields:

Matter, convergence, cosmic shear

Angular correlation function

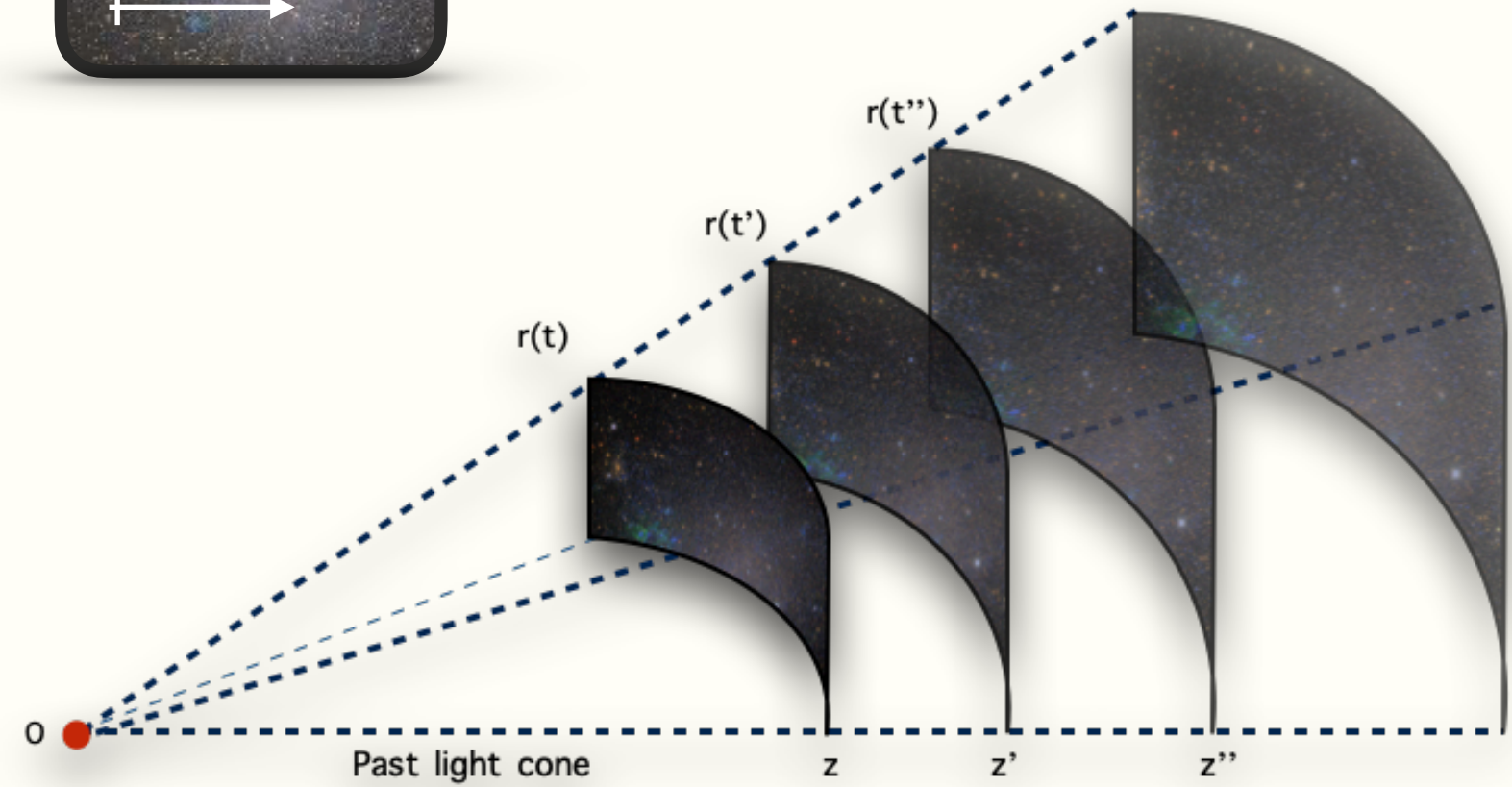
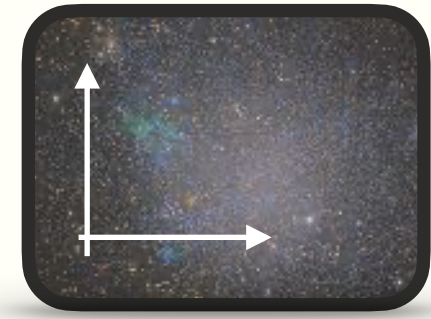
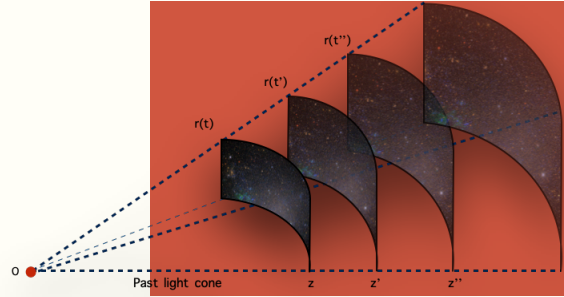
$$w(\theta) = \int_0^\infty \int_0^\infty dx_1 dx_2 \underbrace{f_1(x_1) f_2(x_2)}_{\text{FILTERS}} \underbrace{\xi(r_{12}; t_1, t_2)}_{\text{UNEQUAL-TIME CORRELATION FUNCTION}}$$

Angular power spectrum

$$C(\ell) = \int_0^\infty \frac{dk}{k^2} \int_0^\infty dx_1 dx_2 \underbrace{f_1(x_1) f_2(x_2) j_\ell(kx_1) j_\ell(kx_2)}_{\text{BESSEL FUNCTIONS}} \underbrace{P(k; t_1, t_2)}_{\text{UNEQUAL-TIME POWER SPECTRUM}}$$

Note: lookback time = redshift = co-moving distance

1. Cosmic Lensing



Correlations between fields

- Same time slice: *equal-time correlators*
- Different time slices: *unequal-time correlators*

Examples of fields:

Matter, convergence, cosmic shear

Angular correlation function

$$w(\theta) = \iint_0^\infty dx_1 dx_2 f_1(x_1) f_2(x_2) \xi(r_{12}; t_1, t_2)$$

HARD TO COMPUTE

FILTERS

UNEQUAL-TIME
CORRELATION FUNCTION

Angular power spectrum

$$C(\ell) = \int_0^\infty \frac{dk}{k^2} \iint_0^\infty dx_1 dx_2 f_1(x_1) f_2(x_2) j_\ell(kx_1) j_\ell(kx_2) P(k; t_1, t_2)$$

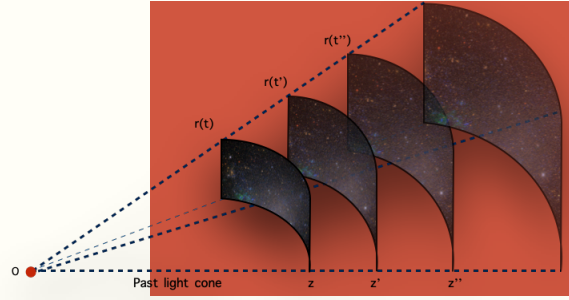
EVEN HARDER TO COMPUTE

BESSEL FUNCTIONS

UNEQUAL-TIME
POWER SPECTRUM

This work

Note: lookback time = redshift = co-moving distance

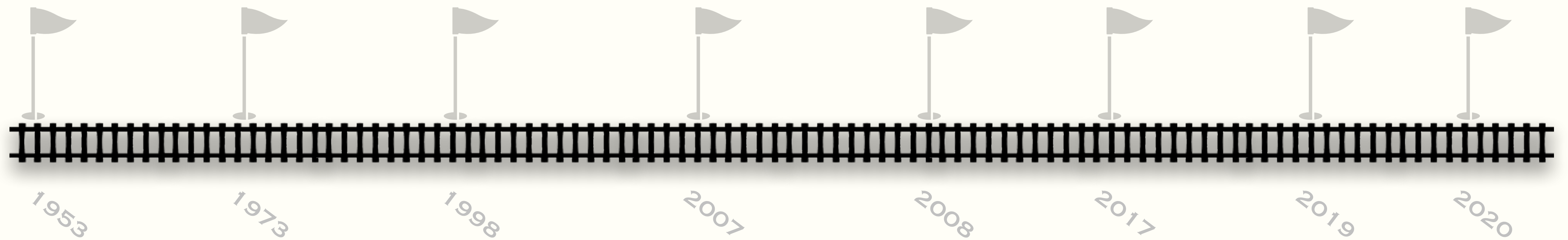


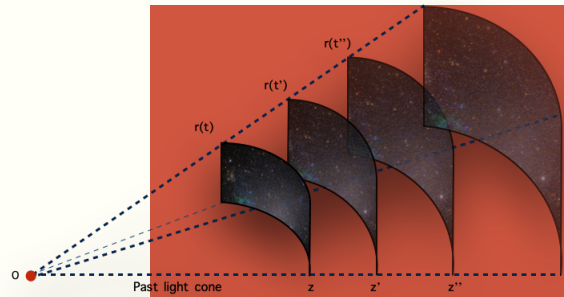
2. The Journey

$$w(\theta) = \iint_0^\infty dx_1 dx_2 f_1(x_1) f_2(x_2) \xi(r_{12}; t_1, t_2)$$

$$C(\ell) = \int_0^\infty \frac{dk}{k^2} \iint_0^\infty dx_1 dx_2 f_1(x_1) f_2(x_2) j_\ell(kx_1) j_\ell(kx_2) P(k; t_1, t_2)$$

- Filters
- Bessel functions
- Unequal-time correlators





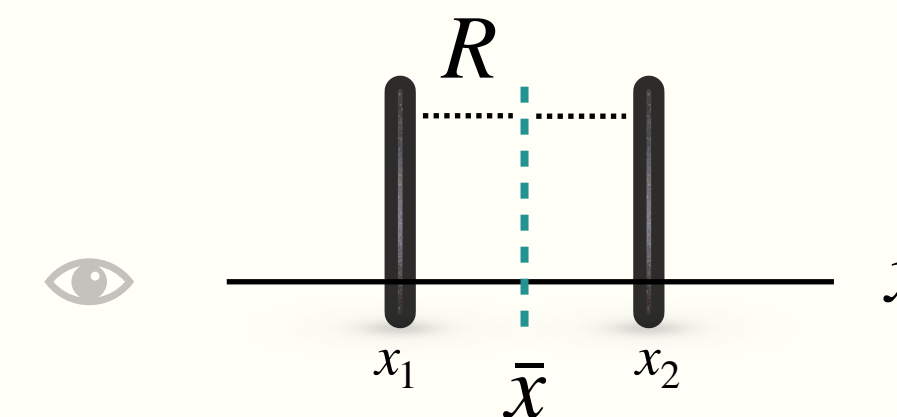
2. The Journey

$$w(\theta) = \iint_0^\infty dx_1 dx_2 f_1(x_1) f_2(x_2) \xi(r_{12}; t_1, t_2)$$

$$C(\ell) = \int_0^\infty \frac{dk}{k^2} \iint_0^\infty dx_1 dx_2 f_1(x_1) f_2(x_2) j_\ell(kx_1) j_\ell(kx_2) P(k; t_1, t_2)$$

Assumptions:

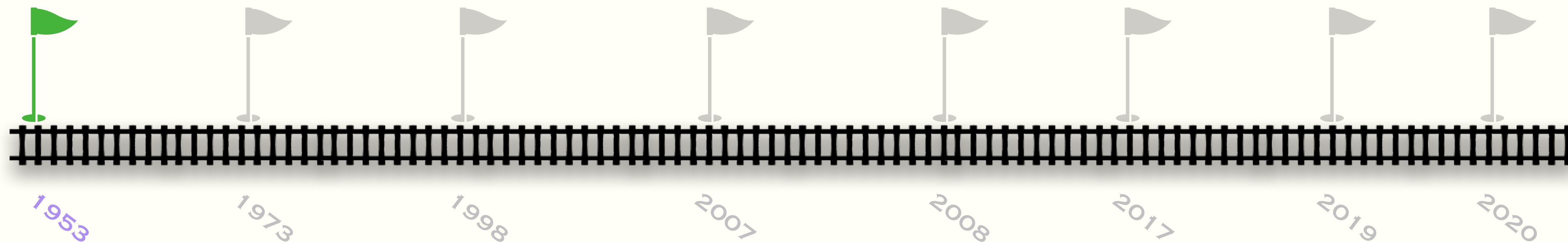
- i) Smooth filters
- ii) Correlation falls off fast
- iii) Small angle separation

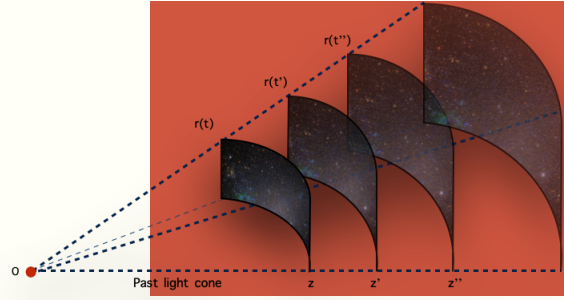


$$w(\theta) \approx \int_0^\infty d\bar{x} \int_{-2\bar{x}}^{2\bar{x}} dR f_1(\bar{x}) f_2(\bar{x}) \xi(r_{12}; \bar{t})$$

- i) Filters
- ii) Mid point

LIMBER





2. The Journey

$$w(\theta) = \iint_0^\infty dx_1 dx_2 f_1(x_1) f_2(x_2) \xi(r_{12}; t_1, t_2)$$

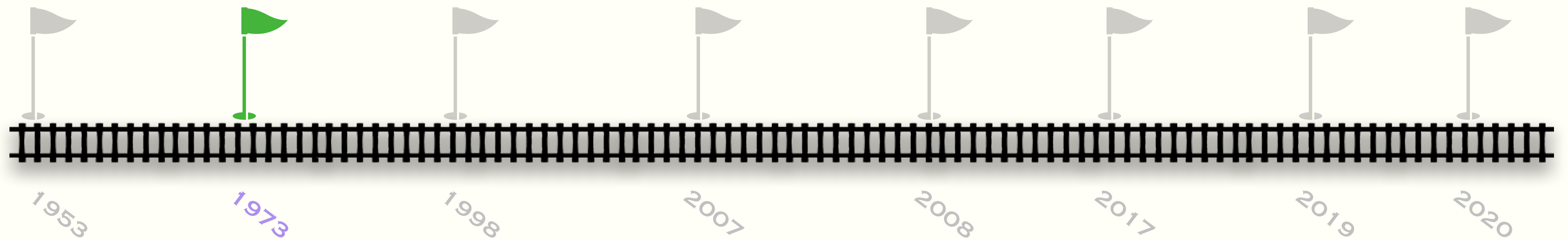
$$C(\ell) = \int_0^\infty \frac{dk}{k^2} \iint_0^\infty dx_1 dx_2 f_1(x_1) f_2(x_2) j_\ell(kx_1) j_\ell(kx_2) P(k; t_1, t_2)$$

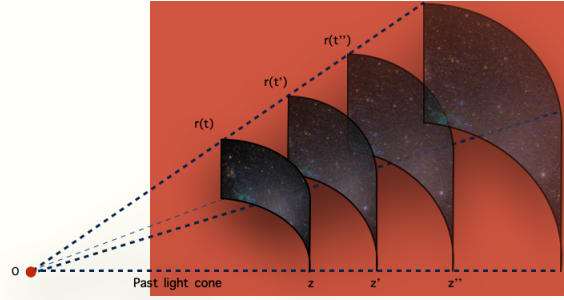
$$w(\theta) \approx \int_0^\infty d\bar{x} \int_{-2\bar{x}}^{2\bar{x}} dR f_1(\bar{x}) f_2(\bar{x}) \xi(r_{12}; \bar{t})$$

- i) Filters
- ii) Mid point
- i) Sphere
- ii) Discrete case

LIMBER

PEEBLES





2. The Journey

$$w(\theta) = \iint_0^\infty dx_1 dx_2 f_1(x_1) f_2(x_2) \xi(r_{12}; t_1, t_2)$$

$$C(\ell) = \int_0^\infty \frac{dk}{k^2} \iint_0^\infty dx_1 dx_2 f_1(x_1) f_2(x_2) j_\ell(kx_1) j_\ell(kx_2) P(k; t_1, t_2)$$

$$C(\ell) = \int d^2 \vec{\theta} w(\theta) e^{-i \vec{\ell} \cdot \vec{\theta}}$$

$$\approx \int dx \frac{f_1(x) f_2(x)}{x^2} P(\ell/x; t)$$

Fourier space
Delta functions

$$w(\theta) \approx \int_0^\infty d\bar{x} \int_{-2\bar{x}}^{2\bar{x}} dR f_1(\bar{x}) f_2(\bar{x}) \xi(r_{12}; \bar{t})$$

- i) Filters
- ii) Mid point

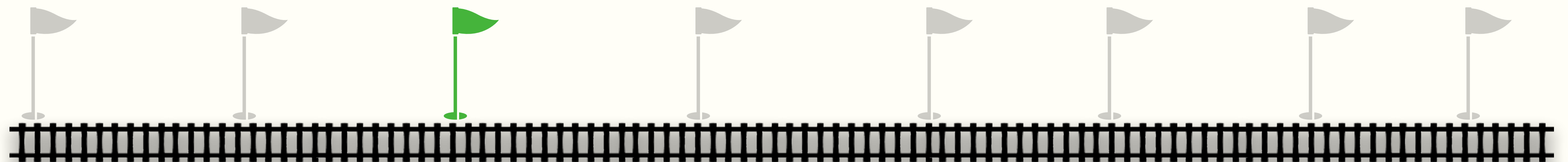
- i) Sphere
- ii) Discrete case

- i) Flat sky
- ii) Curved universe
- iii) Dirac delta

LIMBER

PEEBLES

KAISER



1953

1973

1998

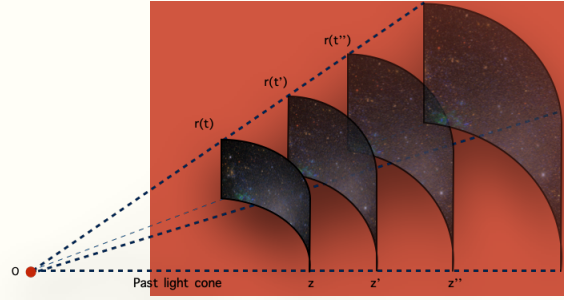
2007

2008

2017

2019

2020



2. The Journey

$$w(\theta) = \iint_0^\infty dx_1 dx_2 f_1(x_1) f_2(x_2) \xi(r_{12}; t_1, t_2)$$

$$C(\ell) = \int_0^\infty \frac{dk}{k^2} \iint_0^\infty dx_1 dx_2 f_1(x_1) f_2(x_2) j_\ell(kx_1) j_\ell(kx_2) P(k; t_1, t_2)$$

Accuracy of Limber:

- Small angles: Limber
- Large angles: Thin-layer approximation
- Compares Limber's vs "exact"
- **Applies small-angle to "exact"**

$$w(\theta) \approx \int_0^\infty d\bar{x} \int_{-2\bar{x}}^{2\bar{x}} dR f_1(\bar{x}) f_2(\bar{x}) \xi(r_{12}; \bar{t})$$

$$C(\ell) \approx \int dx \frac{f_1(x) f_2(x)}{x^2} P(\ell/x; t)$$

- i) Filters
- ii) Mid point

- i) Sphere
- ii) Discrete case

- i) Flat sky
- ii) Curved universe
- iii) Dirac delta

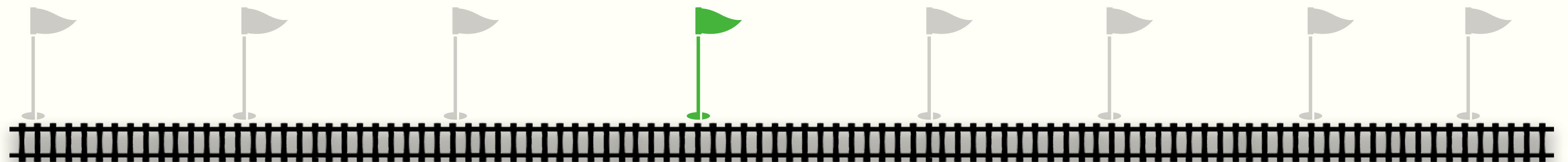
- i) Small and large angles
- ii) Thin-layer approx.
- iii) **Limber's inaccurate**

LIMBER

PEEBLES

KAISER

SIMON



1953

1973

1998

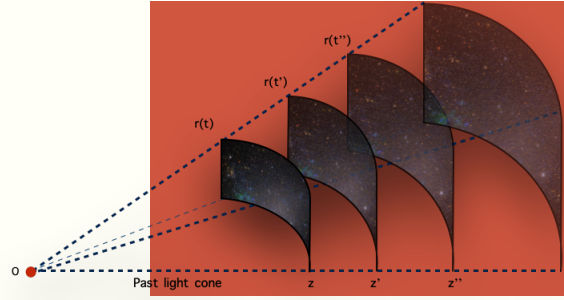
2007

2008

2017

2019

2020



2. The Journey

$$w(\theta) = \iint_0^\infty dx_1 dx_2 f_1(x_1) f_2(x_2) \xi(r_{12}; t_1, t_2)$$

$$C(\ell) = \int_0^\infty \frac{dk}{k^2} \iint_0^\infty dx_1 dx_2 f_1(x_1) f_2(x_2) j_\ell(kx_1) j_\ell(kx_2) P(k; t_1, t_2)$$

Dirac delta version:

$$j_\ell(kx) \approx \sqrt{\frac{\pi}{2\nu}} \delta_D(kx - \nu)$$

Series expansion in $1/\nu$

$$\nu = \ell + 1/2$$

$$w(\theta) \approx \int_0^\infty d\bar{x} \int_{-2\bar{x}}^{2\bar{x}} dR f_1(\bar{x}) f_2(\bar{x}) \xi(r_{12}; \bar{t})$$

$$C(\ell) \approx \int dx \frac{f_1(x) f_2(x)}{x^2} P(\ell/x; t)$$

- i) Filters
- ii) Mid point

- i) Sphere
- ii) Discrete case

- i) Flat sky
- ii) Curved universe
- iii) Dirac delta

- i) Small and large angles
- ii) Thin-layer approx.
- iii) **Limber's inaccurate**

- i) Post-Limber
- ii) **Divergence small ℓ**

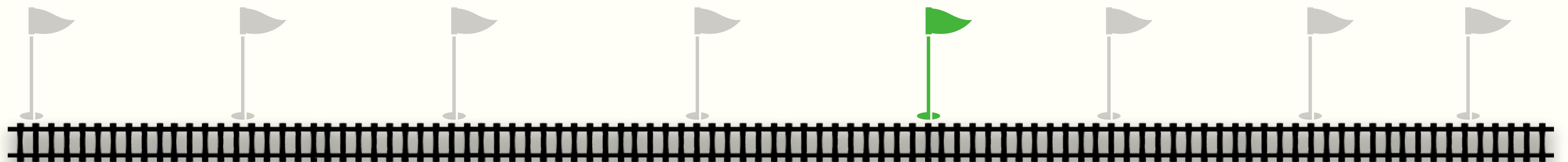
LIMBER

PEEBLES

KAISER

SIMON

LOVERDE
ET AL



1953

1973

1998

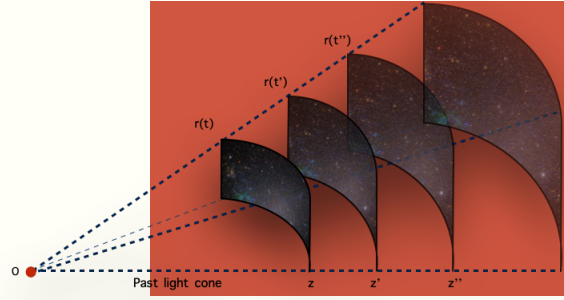
2007

2008

2017

2019

2020



2. The Journey

$$w(\theta) = \iint_0^\infty dx_1 dx_2 f_1(x_1) f_2(x_2) \xi(r_{12}; t_1, t_2)$$

$$C(\ell) = \int_0^\infty \frac{dk}{k^2} \iint_0^\infty dx_1 dx_2 f_1(x_1) f_2(x_2) j_\ell(kx_1) j_\ell(kx_2) P(k; t_1, t_2)$$

Geometric approximation (large scales)

$$P(k; t_1, t_2) \approx \sqrt{P(k, t_1)P(k, t_2)}$$

$$j_\ell(kx) \approx \sqrt{\frac{\pi}{2\nu}} \delta_D(kx - \nu)$$

$$C(\ell) \approx \frac{\pi}{2\nu} \int \frac{dk}{k^2} f_1(\nu/k) f_2(\nu/k) P(k, \nu/k)$$

$$w(\theta) \approx \int_0^\infty d\bar{x} \int_{-2\bar{x}}^{2\bar{x}} dR f_1(\bar{x}) f_2(\bar{x}) \xi(r_{12}; \bar{t})$$

$$C(\ell) \approx \int dx \frac{f_1(x) f_2(x)}{x^2} P(\ell/x; t)$$

Expansion $1/\nu$

- i) Filters
- ii) Mid point

- i) Sphere
- ii) Discrete case

- i) Flat sky
- ii) Curved universe
- iii) Dirac delta

- i) Small and large angles
- ii) Thin-layer approx.
- iii) **Limber's inaccurate**

- i) Post-Limber
- ii) **Divergence small ℓ**

- i) Geometric
- ii) Dirac delta
- iii) **Valid large ℓ**
- iv) **Valid small k**

LIMBER

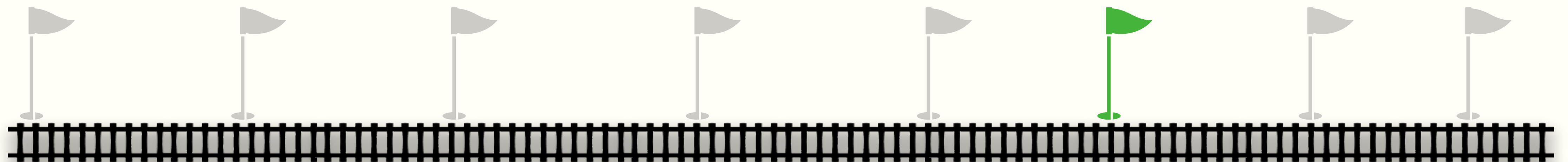
PEEBLES

KAISER

SIMON

LOVERDE
ET AL

LE MOS
ET AL



1953

1973

1998

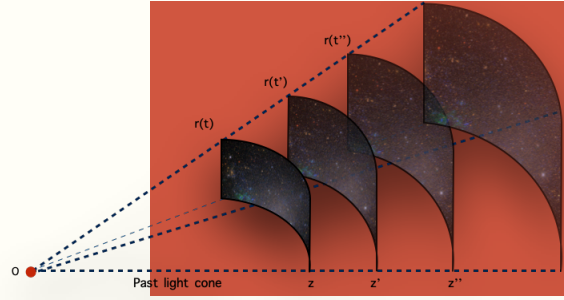
2007

2008

2017

2019

2020



2. The Journey

$$w(\theta) = \iint_0^\infty dx_1 dx_2 f_1(x_1) f_2(x_2) \xi(r_{12}; t_1, t_2)$$

$$C(\ell) = \int_0^\infty \frac{dk}{k^2} \iint_0^\infty dx_1 dx_2 f_1(x_1) f_2(x_2) j_\ell(kx_1) j_\ell(kx_2) P(k; t_1, t_2)$$

$$C(\ell) \approx \int_0^\infty \frac{dk}{k^2} (\dots) P_{lin}(k; t_1, t_2) + \int_0^\infty \frac{dk}{k^2} (\dots) P_{NL}(k; t_1, t_2)$$

Linear
Exact
 $P(k; t_1, t_2) \equiv \sqrt{P(k, t_1)P(k, t_2)}$

Non-linear
No unequal-time
Dirac delta $\ell \gg 1$
Geometric $k \ll 1$?
NL Physics?

$$w(\theta) \approx \int_0^\infty d\bar{x} \int_{-2\bar{x}}^{2\bar{x}} dR f_1(\bar{x}) f_2(\bar{x}) \xi(r_{12}; \bar{t})$$

$$C(\ell) \approx \int dx \frac{f_1(x) f_2(x)}{x^2} P(\ell/x; t)$$

Expansion $1/\nu$

$$C(\ell) \approx \frac{\pi}{2\nu} \int \frac{dk}{k^2} f_1(\nu/k) f_2(\nu/k) P(k, \nu/k)$$

- i) Filters
- ii) Mid point

- i) Sphere
- ii) Discrete case

- i) Flat sky
- ii) Curved universe
- iii) Dirac delta

- i) Small and large angles
- ii) Thin-layer approx.
- iii) **Limber's inaccurate**

- i) Post-Limber
- ii) **Divergence small ℓ**

- i) Geometric
- ii) Dirac delta
- iii) **Valid large ℓ**
- iv) **Valid small k**

- i) FFTLog
- ii) Split regimes
- iii) **Limber Inaccurate**
- iv) **Accuracy large k ?**

LIMBER

PEEBLES

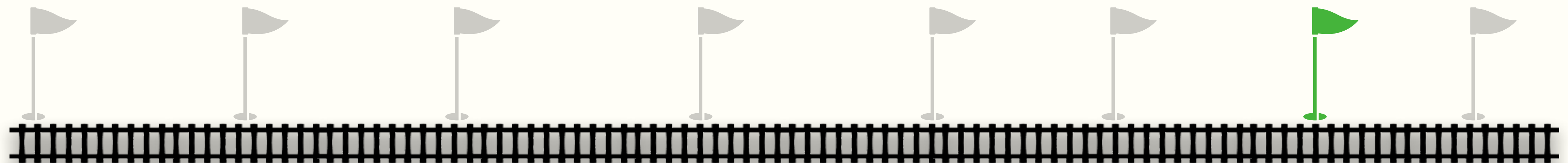
KAISER

SIMON

LOVERDE
ET AL

LEMONS
ET AL

FANG
ET AL



1953

1973

1998

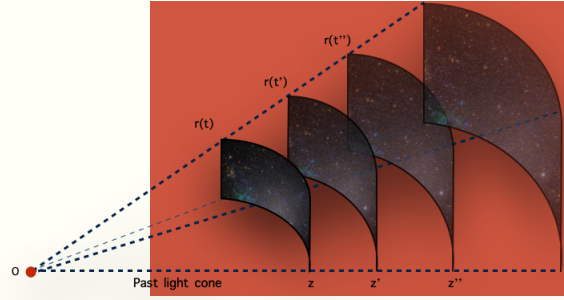
2007

2008

2017

2019

2020



2. The Journey

$$w(\theta) = \iint_0^\infty dx_1 dx_2 f_1(x_1) f_2(x_2) \xi(r_{12}; t_1, t_2)$$

$$C(\ell) = \int_0^\infty \frac{dk}{k^2} \iint_0^\infty dx_1 dx_2 f_1(x_1) f_2(x_2) j_\ell(kx_1) j_\ell(kx_2) P(k; t_1, t_2)$$

Issues

- Limber's accuracy
- Divergence on large angles
- Geometric app good on large scales
- Accuracy of non-linear physics

Need

- ★ Unequal-time correlators
- ★ All-angle calculations

$$w(\theta) \approx \int_0^\infty d\bar{x} \int_{-2\bar{x}}^{2\bar{x}} dR f_1(\bar{x}) f_2(\bar{x}) \xi(r_{12}; \bar{t})$$

$$C(\ell) \approx \int dx \frac{f_1(x) f_2(x)}{x^2} P(\ell/x; t)$$

Expansion $1/\nu$

$$C(\ell) \approx \frac{\pi}{2\nu} \int \frac{dk}{k^2} f_1(\nu/k) f_2(\nu/k) P(k, \nu/k)$$

- i) Filters
- ii) Mid point

- i) Sphere
- ii) Discrete case

- i) Flat sky
- ii) Curved universe
- iii) Dirac delta

- i) Small and large angles
- ii) Thin-layer approx.
- iii) **Limber's inaccurate**

- i) Post-Limber
- ii) **Divergence small ℓ**

- i) Geometric
- ii) Dirac delta
- iii) **Valid large ℓ**
- iv) **Valid small k**

- i) FFTLog
- ii) Split regimes
- iii) **Limber Inaccurate**
- iv) **Accuracy large k ?**

LIMBER

PEEBLES

KAISER

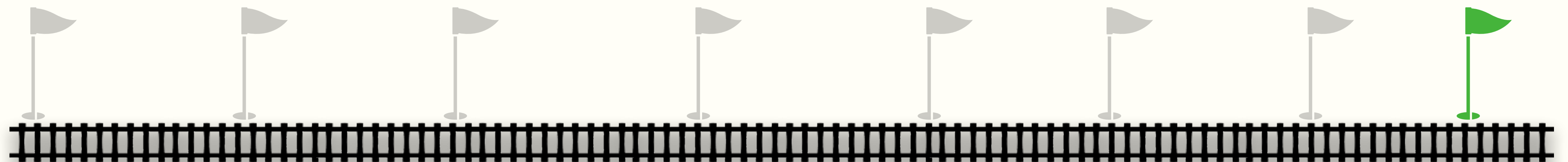
SIMON

LOVERDE
ET AL

LEMONS
ET AL

FANG
ET AL

This work



1953

1973

1998

2007

2008

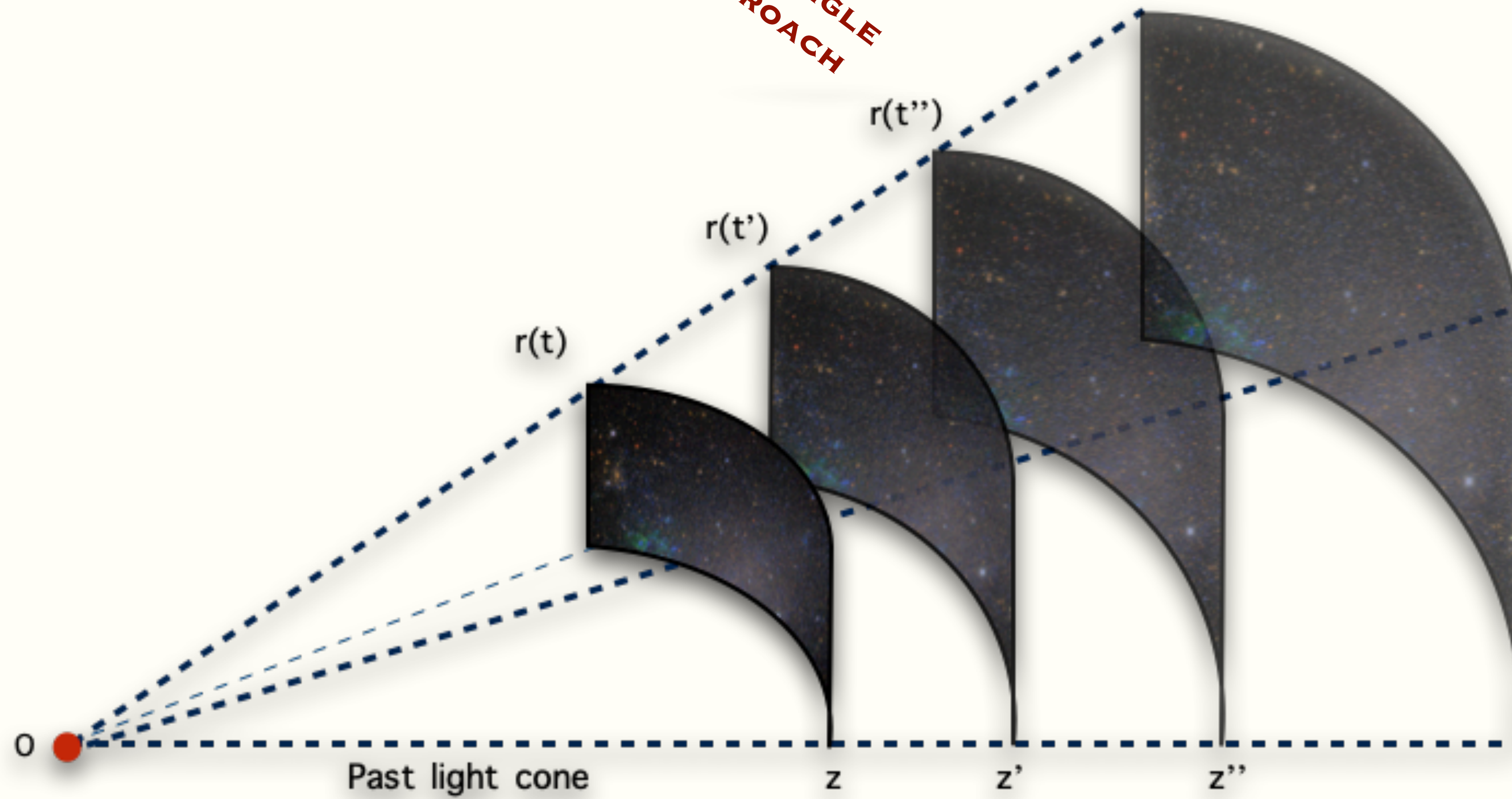
2017

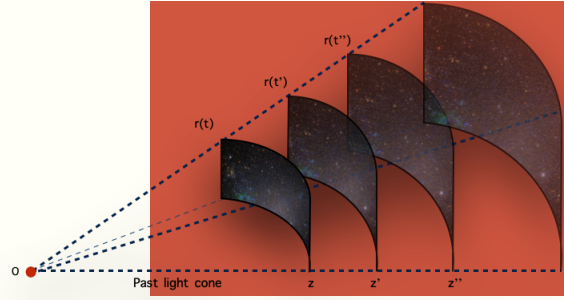
2019

2020

N. Tessore, L. F. de la Bella (in prep)
Python package **CORFU**

**3. ALL-ANGLE
APPROACH**



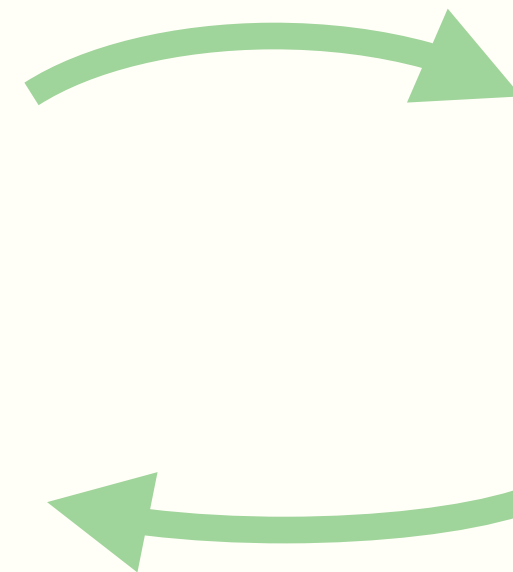


3. All-angle approach

$$w(\theta) = \iint_0^\infty dx_1 dx_2 f_1(x_1) f_2(x_2) \xi(r_{12}; t_1, t_2)$$

$$C(\ell) = \int_0^\infty \frac{dk}{k^2} \iint_0^\infty dx_1 dx_2 f_1(x_1) f_2(x_2) j_\ell(kx_1) j_\ell(kx_2) \boxed{P(k; t_1, t_2)}$$

- Need accurate unequal-time power spectrum
- Deal with non-linear physics (one-loop, EFT)
- Impact on weak lensing. How?
 - Midpoint approximation ★
 - Compare with geometric approx.
- Analyse regimes of validity
- Drop approximations
- Python package **UNEQUALPY**

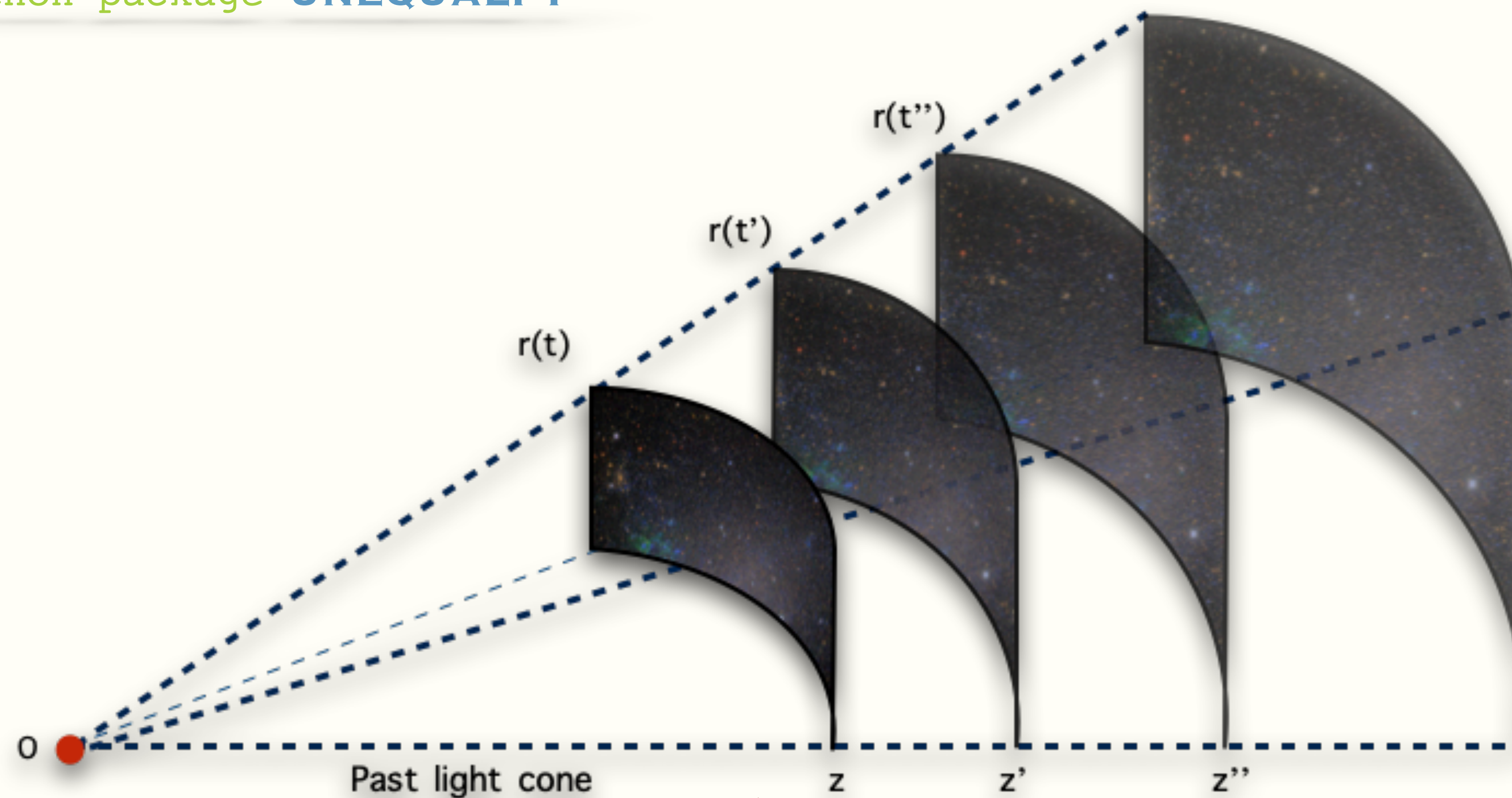


- Need exact calculation at all angular scales
- Python package **CORFU**
 - FFTLog methods
 - Inverse Fourier transform
- Legendre polynomials

$$P(k; t_1, t_2) \longrightarrow \xi(r; t_1, t_2)$$

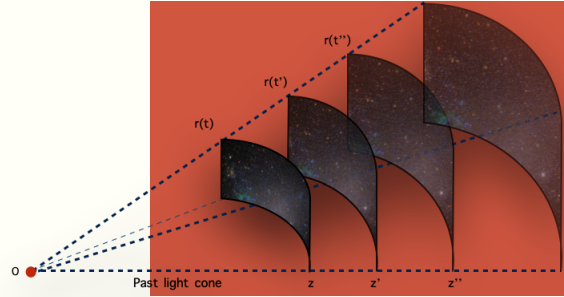
$$w(\theta) \longrightarrow C(\ell)$$

L. F. de la Bella, N. Tessore and
S. L. Bridle (arXiv 2011.06185)
Python package **UNEQUALPY**



**4. UNEQUAL-TIME
CORRELATORS**

4. Unequal-time Correlators



2point functions

Homogeneous
Isotropic

- EQUAL-TIME POWER SPECTRUM

Same time slice

- UNEQUAL-TIME POWER SPECTRUM

Different time slices

$$(2\pi)^3 \delta_D(\mathbf{k}_1 + \mathbf{k}_2) P(k, z) := \langle \delta(\mathbf{k}_1, z) \delta^*(\mathbf{k}_2, z) \rangle$$

$$(2\pi)^3 \delta_D(\mathbf{k}_1 + \mathbf{k}_2) P(k, z_1, z_2) := \langle \delta(\mathbf{k}_1, z_1) \delta^*(\mathbf{k}_2, z_2) \rangle$$

LINEAR THEORY

$$P(k, z) = D(z)^2 P_{11}(k)$$

$$P(k, z_1, z_2) = D(z_1) D(z_2) P_{11}(k)$$

@ $z_1 = z_2$

$$P(k, z, z) = P(k, z)$$

Non-linear

ONE-LOOP THEORY

$$P(k, z_1, z_2) = P_{11}(k, z_1, z_2) + P_{22}(k, z_1, z_2) + P_{13}(k, z_1, z_2)$$

- STANDARD PERTURBATION THEORY
- EFFECTIVE FIELD THEORY

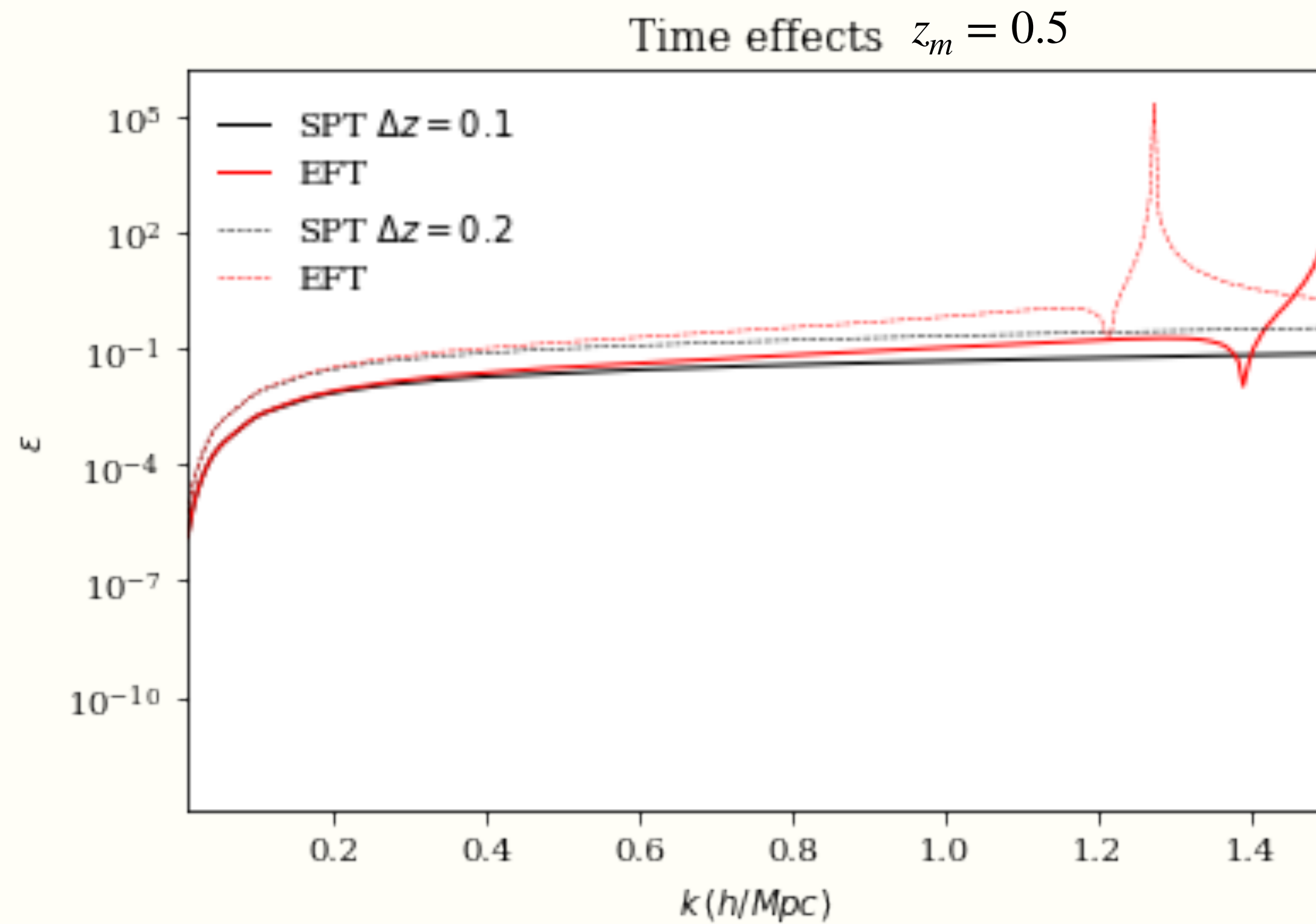
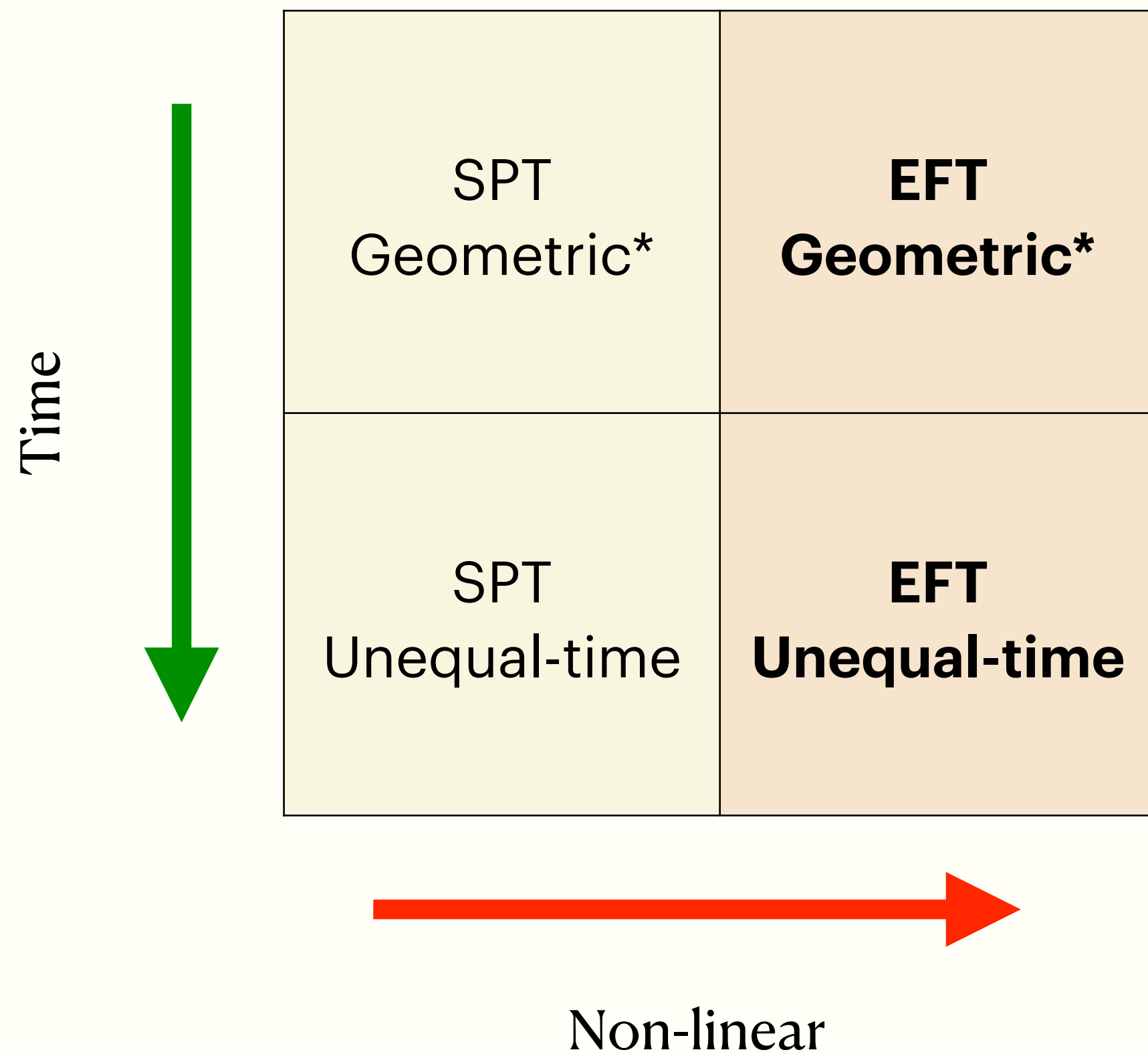
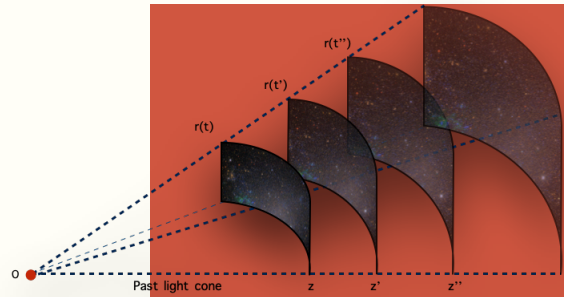
NL physics \rightarrow counterterms

$$P_{EFT} = P_{SPT} - c^2 k^2 P_{11}$$

- This work shows unequal-time EFT breaks
- New idea: the midpoint approximation!



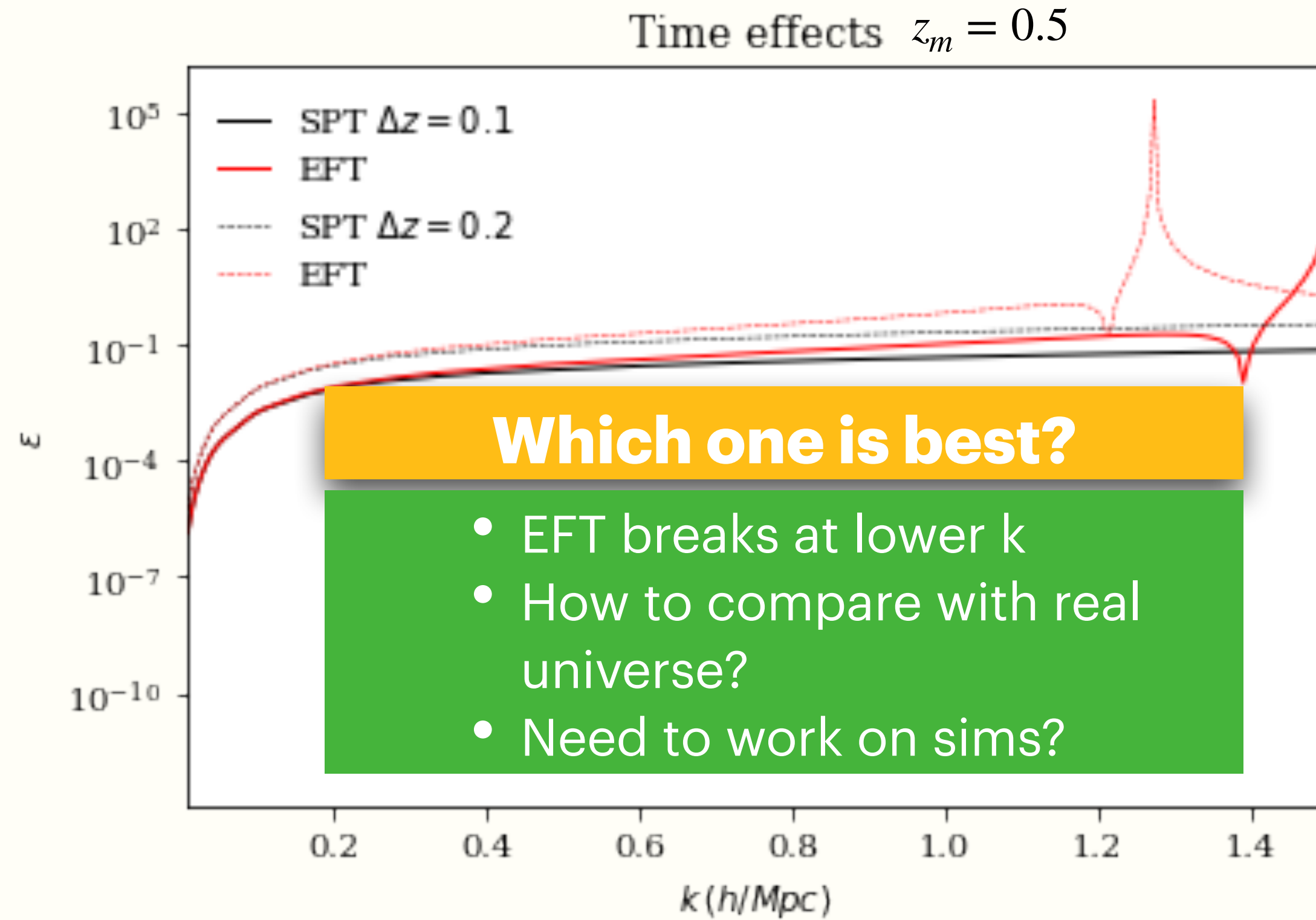
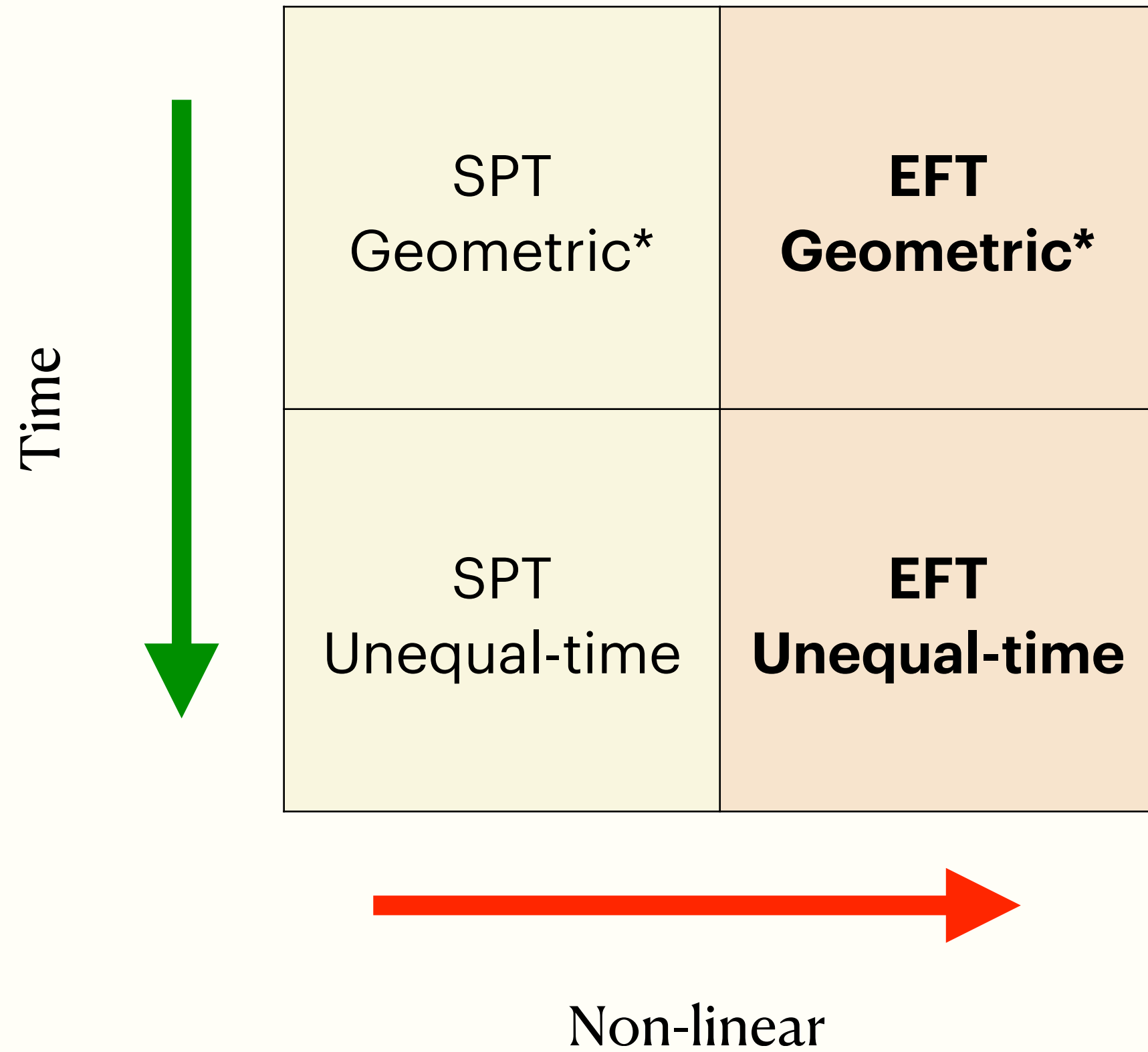
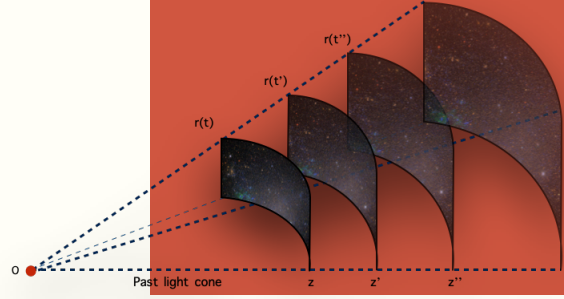
4.1. Time and non-linear effects



* $P(k; t_1, t_2) \approx \sqrt{P(k, t_1)P(k, t_2)}$

$$\frac{\Delta P}{P^2} = \frac{|P_{theory}(k, z_1)P_{theory}(k, z_2) - P_{theory}(k; z_1, z_2)^2|}{P_{theory}(k; z_1, z_2)^2}$$

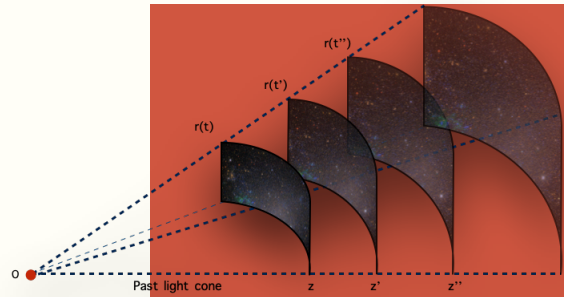
4.1. Time and non-linear effects



* $P(k; t_1, t_2) \approx \sqrt{P(k, t_1)P(k, t_2)}$

$$\frac{\Delta P}{P^2} = \frac{|P_{theory}(k, z_1)P_{theory}(k, z_2) - P_{theory}(k; z_1, z_2)^2|}{P_{theory}(k; z_1, z_2)^2}$$

4.2. Midpoint approximation



- Unequal-time power spectrum

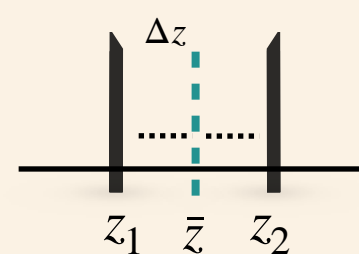
$$P(k; z_1, z_2)$$

- Geometric approximation

$$\sqrt{P(k, z_1)P(k, z_2)}$$

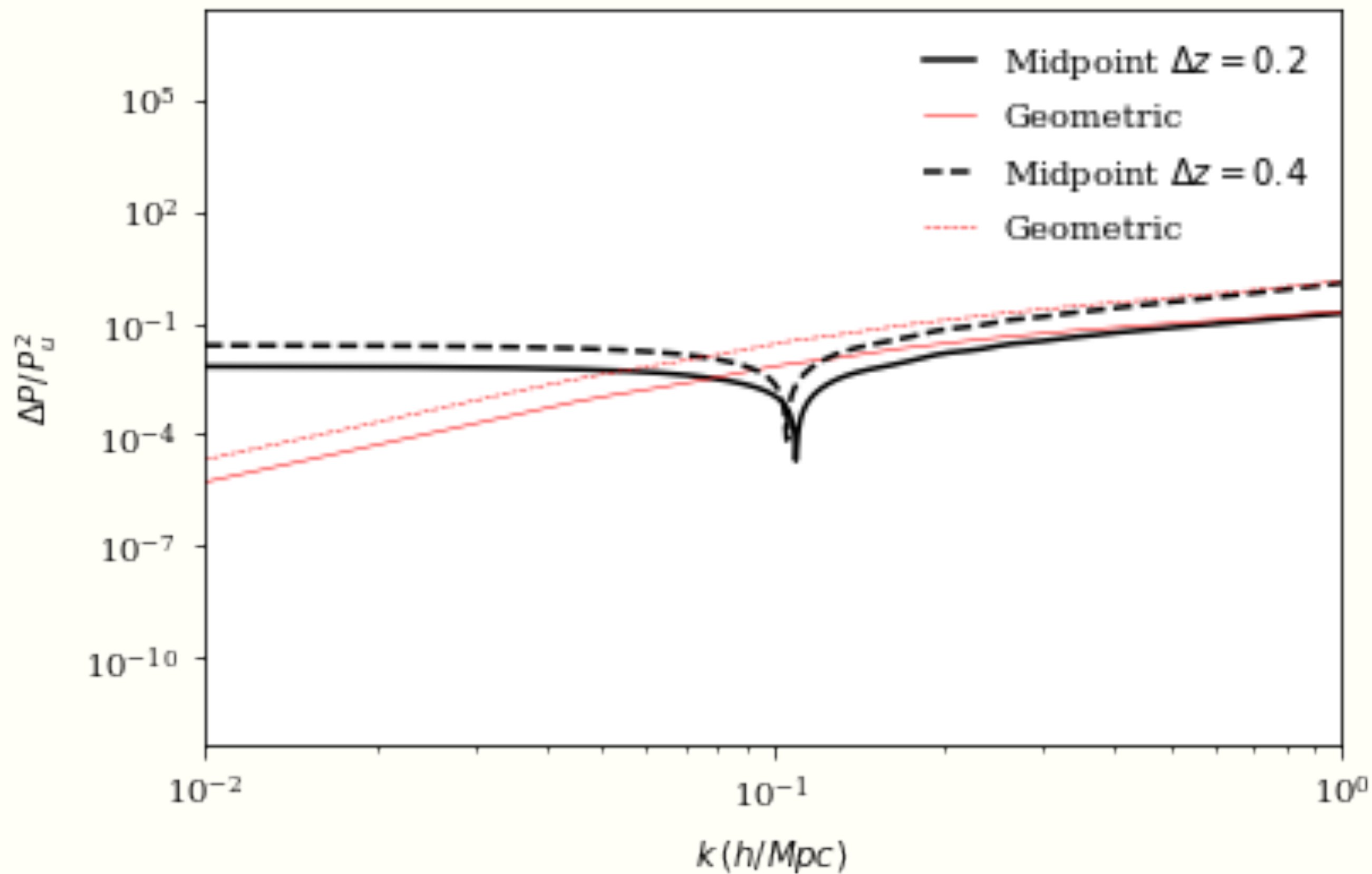
- Midpoint approximation

$$\star P(k; z_m)$$

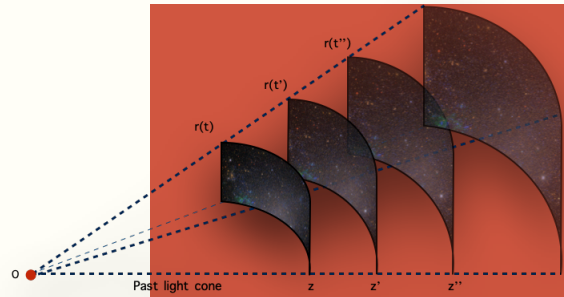


- More accurate for small separation
- Geometric approximation
 - Good on very large scales
 - Error increases on smaller scales
- Midpoint approximation
 - Not perfect on large scales
 - Error much more stable
 - Better prediction on small scales
- EFT does not improve SPT

Mean redshift $z = 0.5$



4.2. Midpoint approximation



- Unequal-time power spectrum

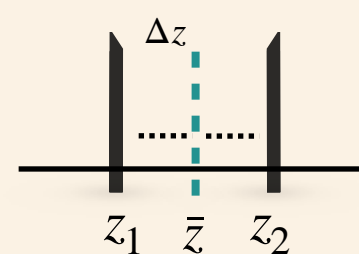
$$P(k; z_1, z_2)$$

- Geometric approximation

$$\sqrt{P(k, z_1)P(k, z_2)}$$

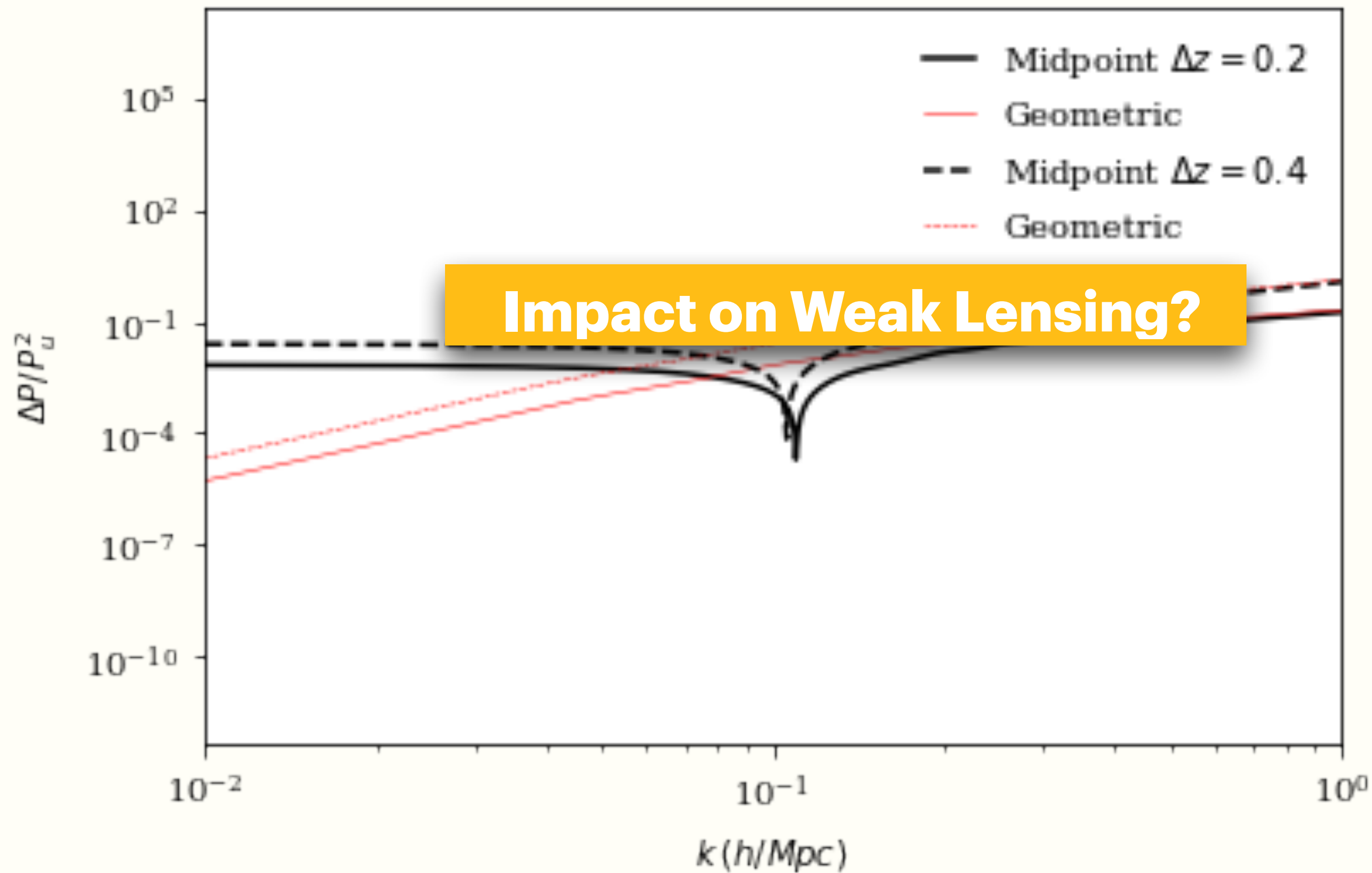
- Midpoint approximation

$$\star P(k; z_m)$$

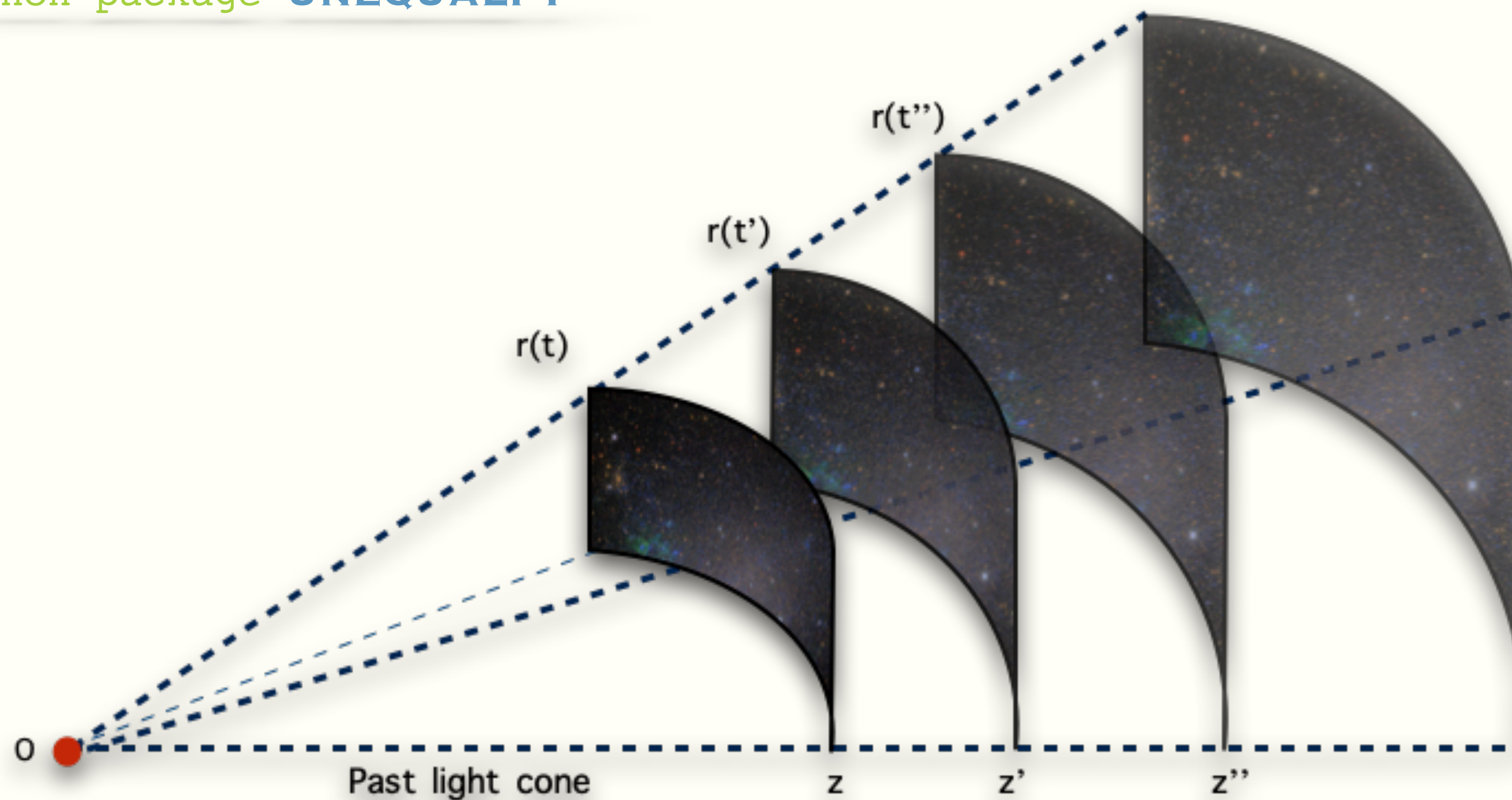


- More accurate for small separation
- Geometric approximation
 - Good on very large scales
 - Error increases on smaller scales
- Midpoint approximation
 - Not perfect on large scales
 - Error much more stable
 - Better prediction on small scales
- EFT does not improve SPT

Mean redshift $z = 0.5$



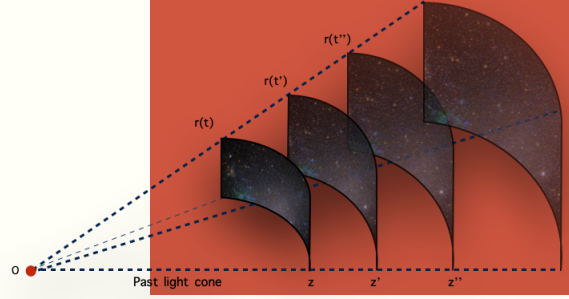
L. F. de la Bella, N. Tessore and
S. L. Bridle (arXiv 2011.06185)
Python package **UNEQUALPY**



5. RESULTS

6. SUMMARY

5. Results



Exact

$$C_{ab}^{(i,j)}(\ell) = \frac{2}{\pi} \int_0^\infty dk k^2 \iint_0^\infty dx_1 dx_2 f_a^i(x_1) f_b^j(x_2) j_\ell(kx_1) j_\ell(kx_2) P(k; t_1, t_2)$$

Limber

$$C_{ab}^{(i,j)}(\ell) \approx \frac{1}{\nu} \int_0^\infty dk k^2 f_a^i(\nu/k) f_b^j(\nu/k) P(k; \nu/k)$$

CORFU

$$P(k; t_1, t_2) \longrightarrow \xi(r; t_1, t_2)$$

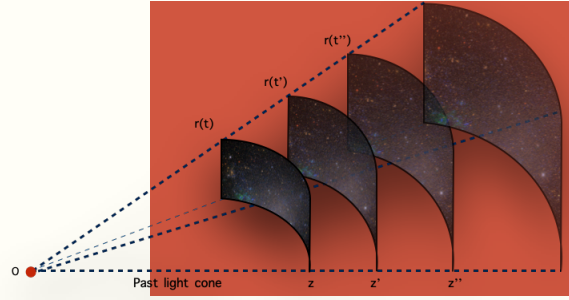
$$w(\theta) \longrightarrow C(\ell)$$

Geometric

$$C_{ab}^{(i,j)}(\ell) \approx \frac{2}{\pi} \int_0^\infty dk k^2 \int_0^\infty dx_1 f_a^i(x_1) j_\ell(kx_1) \sqrt{P(k; x_1)} \int_0^\infty dx_2 f_b^j(x_2) j_\ell(kx_2) \sqrt{P(k; x_2)}$$

Midpoint

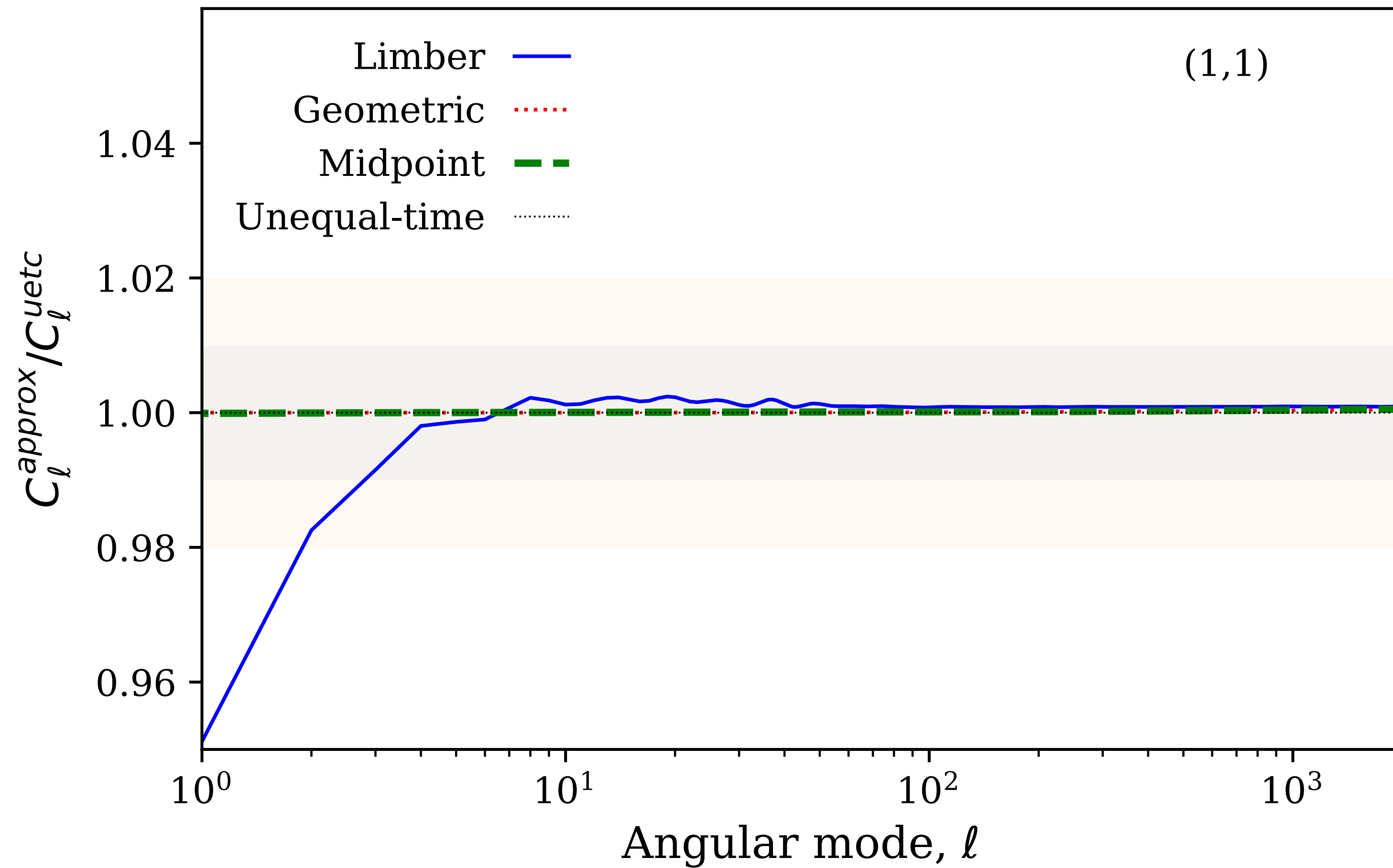
$$C_{ab}^{(i,j)}(\ell) \approx \frac{2}{\pi} \int_0^\infty dk k^2 \iint_0^\infty dx_1 dx_2 f_a^i(x_1) f_b^j(x_2) j_\ell(kx_1) j_\ell(kx_2) P\left(k; \frac{x_1 + x_2}{2}\right)$$



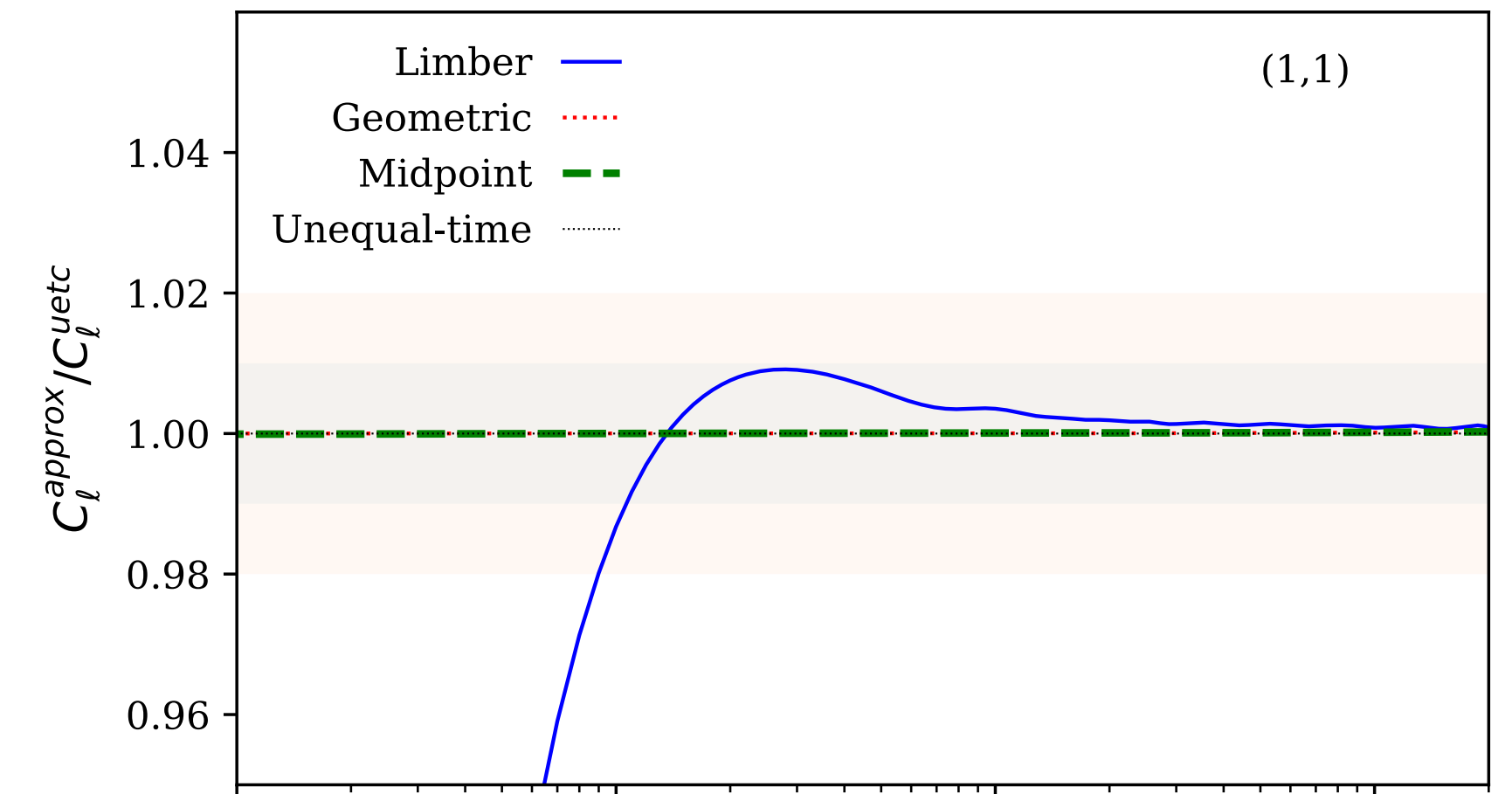
5. Results

DES-Y1 DATA

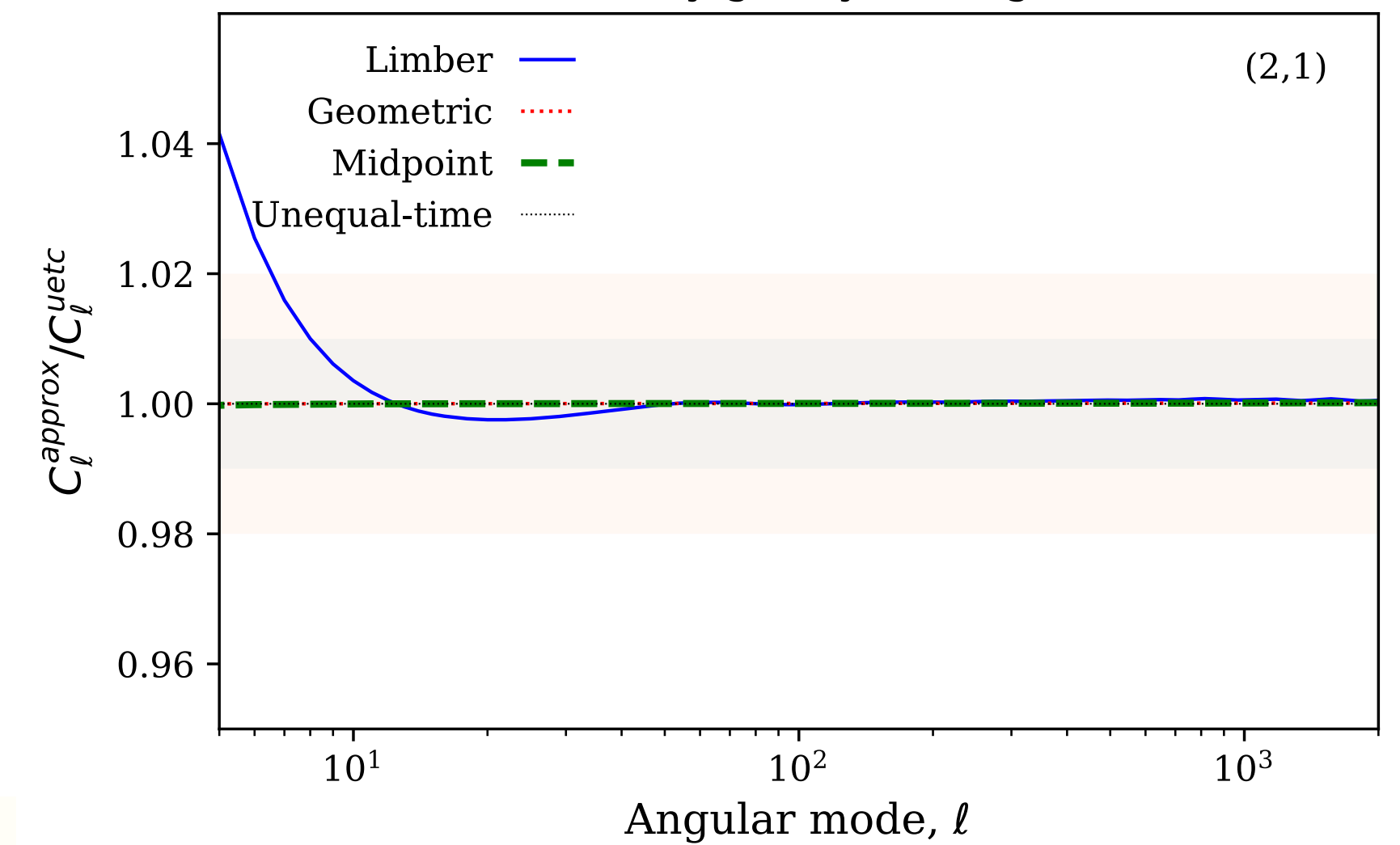
Convergence



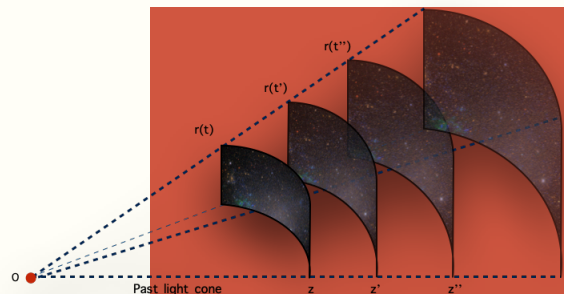
Galaxy clustering



Galaxy-galaxy lensing

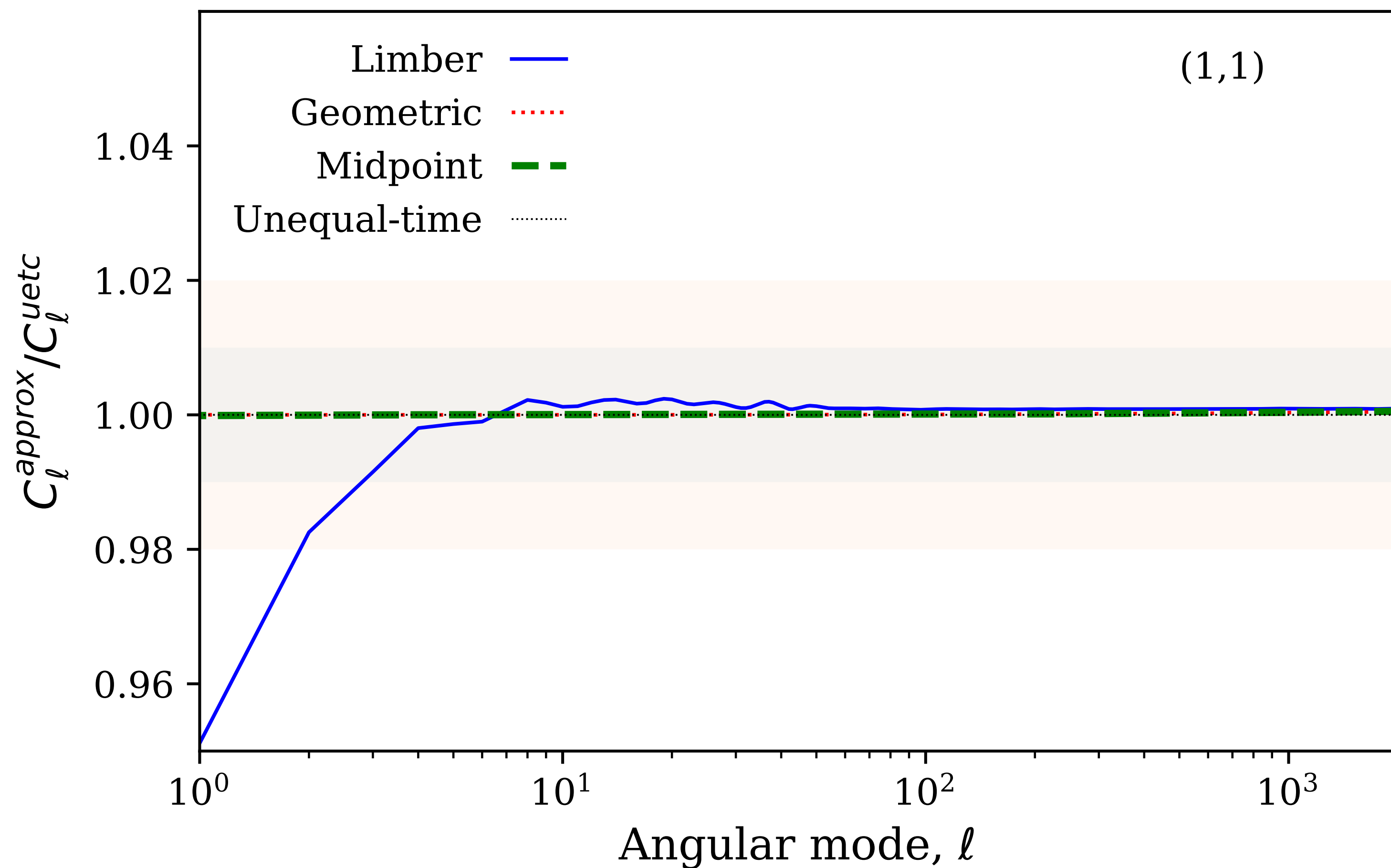


5. Results

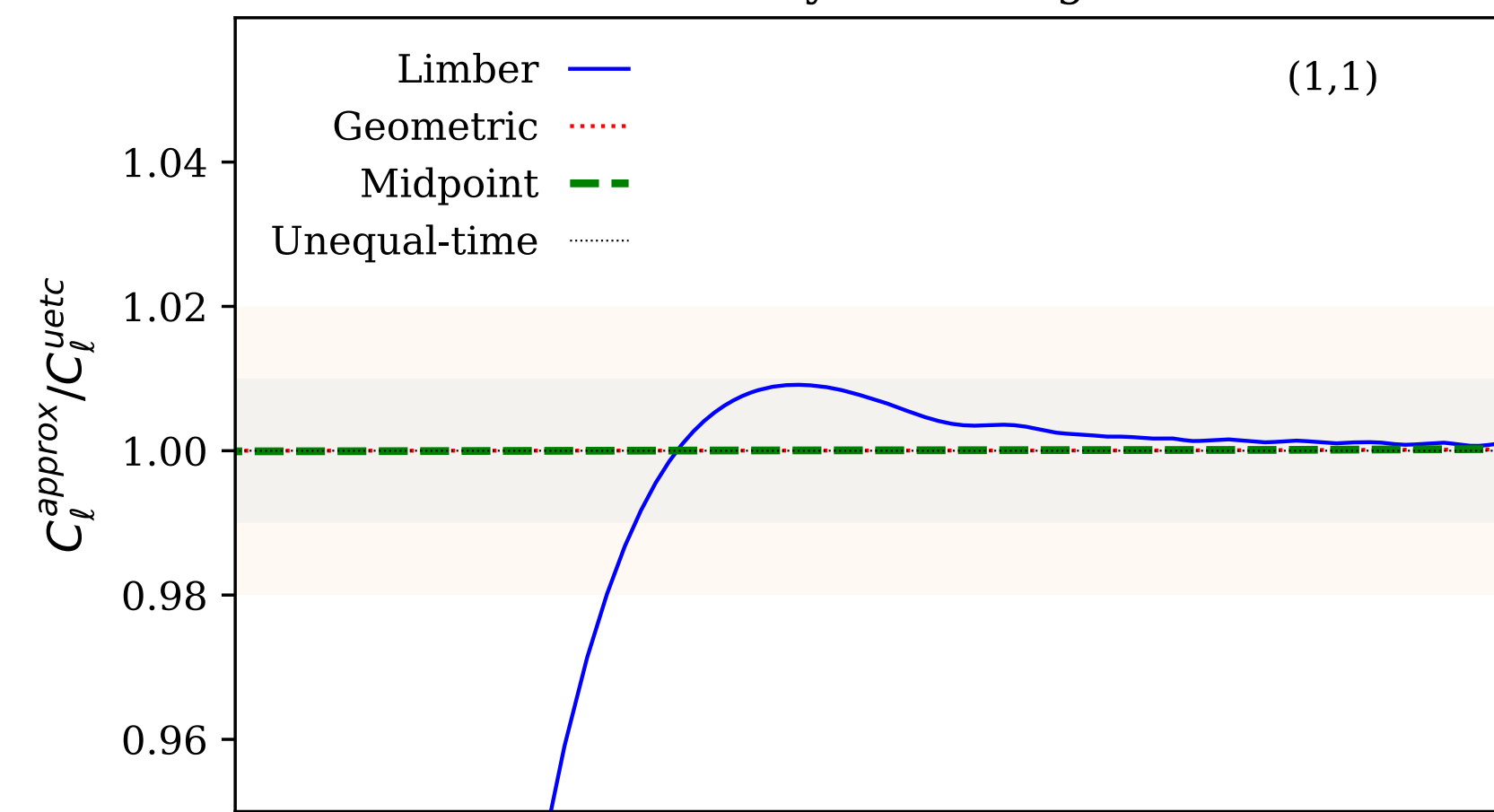


DES-Y1 DATA

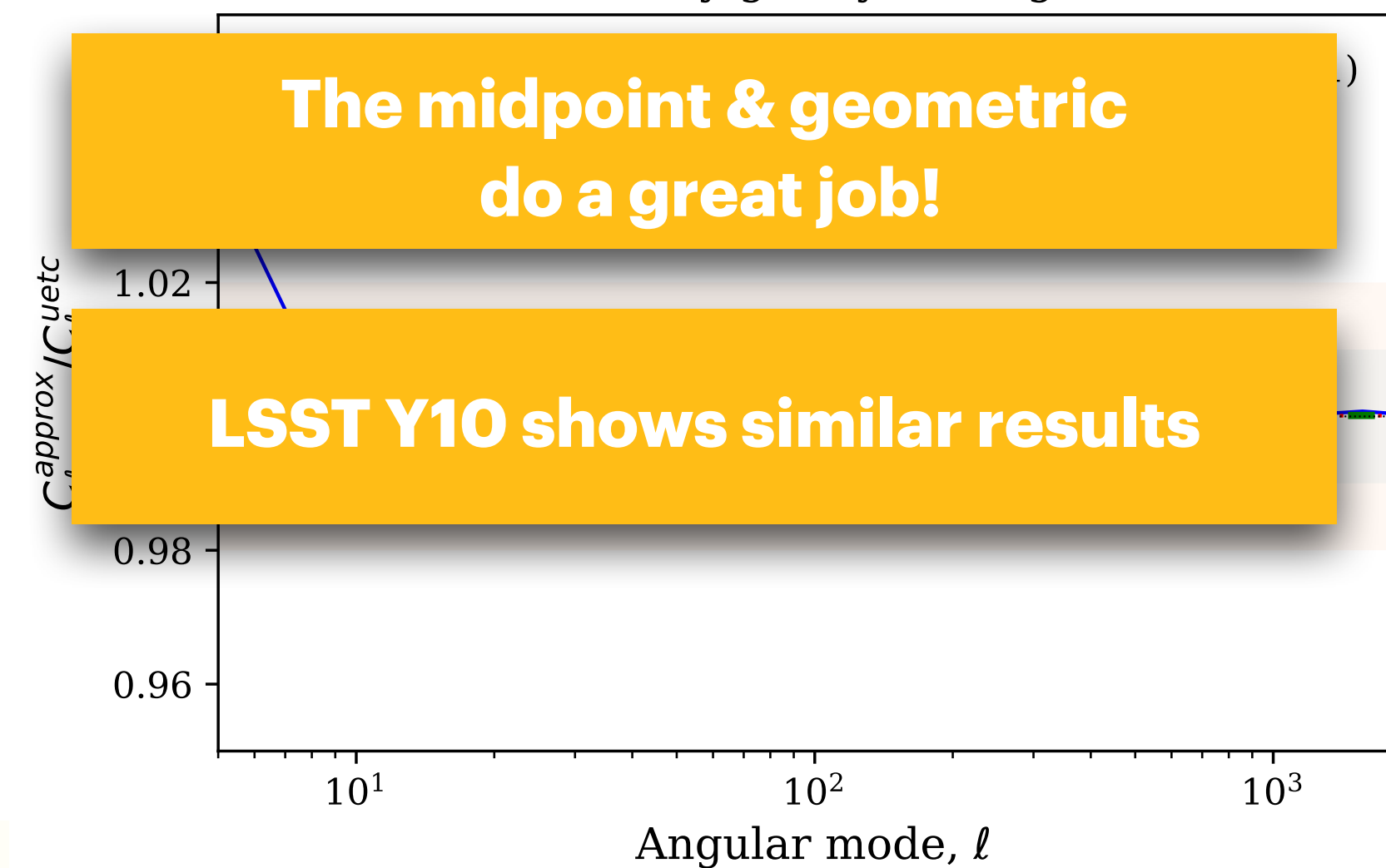
Convergence

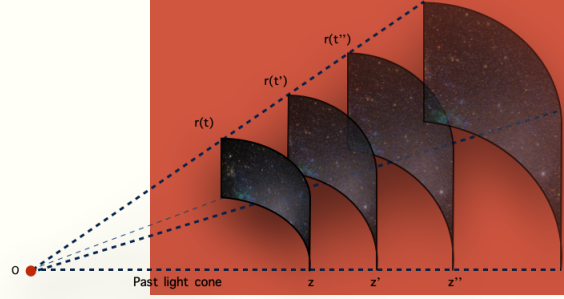


Galaxy clustering



Galaxy-galaxy lensing





6. Summary

Summary

- Angular correlations functions are very hard to compute
- **Limber** is the most widely used approximation
- List of issues: accuracy and validity of approximations
- Need for unequal-time correlators and all-angle computations
- Unequal-time **EFT** does not improve the prediction
- **Midpoint approximation** better on non-linear scales

Coming next!

- **Numeric paper (in prep)**
 - All angular scale computations
 - Equal and unequal-time correlators
 - Python package **CORFU**
- **Science paper (arXiv 2011.06185)**
 - Unequal-time matter power spectrum at one-loop
 - Python package **UNEQUALPY**
 - Analysis of all approximations and validity regimes
 - **Midpoint and Geometric** best to mimic unequal-time features!
 - **Beyond Limber** relevant for **galaxy clustering** and **galaxy-galaxy lensing**.

QUESTIONS?

