

Galaxy Demographics and the SkyPy project

Lucia F. de la Bella USM 13/07/2022

Who I am...



<https://orcid.org/0000-0002-1064-3400>

<https://howtoreachthecosmos.jimdofree.com>



B.Sc

Quantum
Cosmology

2008-2013



M.Sc

EFTDE

2013-2014



University of Sussex

PhD

EFTLSS
RSDs

2014-2018

Teaching

2018-2019



The University of Manchester

PDRA



Weak Lensing
Unequal-time
correlators
2019-2021

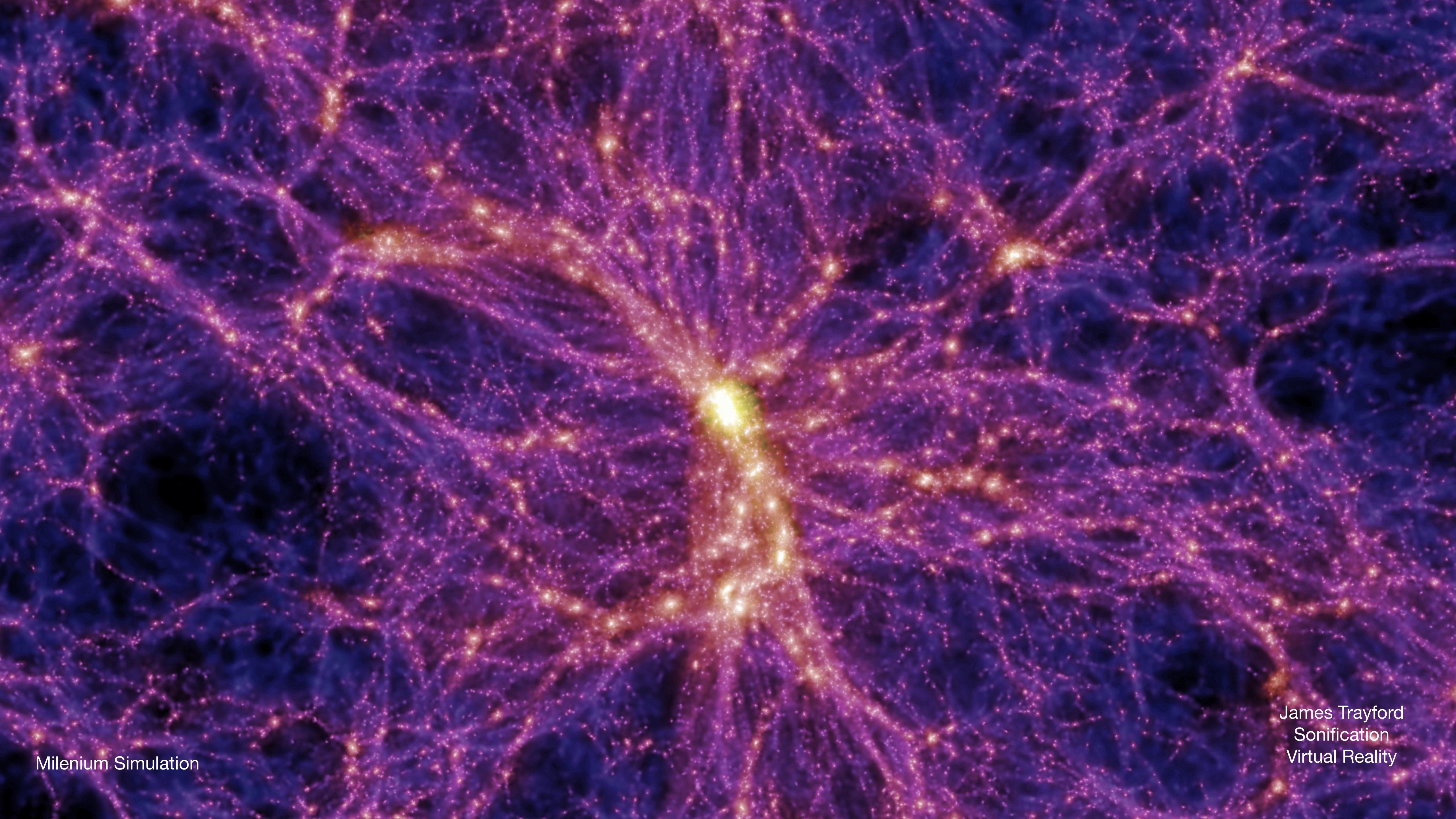
2021-present



I. Galaxy Demographics

L. F. de la Bella, A. Amara, S. Birrer, W. Hartley & P. Sudek, [arXiv 2112.11110](https://arxiv.org/abs/2112.11110)

1. Context and goals
2. Method
3. Validation



Milenum Simulation

James Trayford
Sonification
Virtual Reality



The challenge

Open-data revolution in Astronomy

Sophisticated analysis methods that heavily rely on realistic simulations

Goal: produce realistic simulations faster and more accurately

This work: galaxy populations

Active & Passive



The challenge

Open-data revolution in Astronomy

Sophisticated analysis methods that heavily rely on realistic simulations

Goal: produce realistic simulations faster and more accurately

This work: galaxy populations

Active & Passive

Blue & Red

Galaxy populations

Tools

Schechter mass function

$$\phi(m, t) dm = \phi_*(t) \left(\frac{m}{m_*} \right)^\alpha e^{-m/m_*} dm$$

- Star-forming galaxies $\{\phi_{*b}, \alpha_b, m_{*b}\}$ Star formation rate: $SFR = \frac{dm}{dt} = m sSFR$
- Satellite quenched $\{\phi_{*\rho}, \alpha_\rho, m_{*\rho}\}$ Quenching rate: $\eta_\rho \simeq 50\%$
- Mass quenched $\{\phi_{*m}, \alpha_{sm}, m_{*m}\}$ Quenching rate: $\eta_m = \frac{SFR}{m_*}$

$\eta \rightarrow$ Probability for isolated active galaxies being quenched per unit time

(Inverse: time scale an isolated active galaxy waits to be quenched)

$f \rightarrow$ Fraction of quenched galaxies



The goal

Main goal

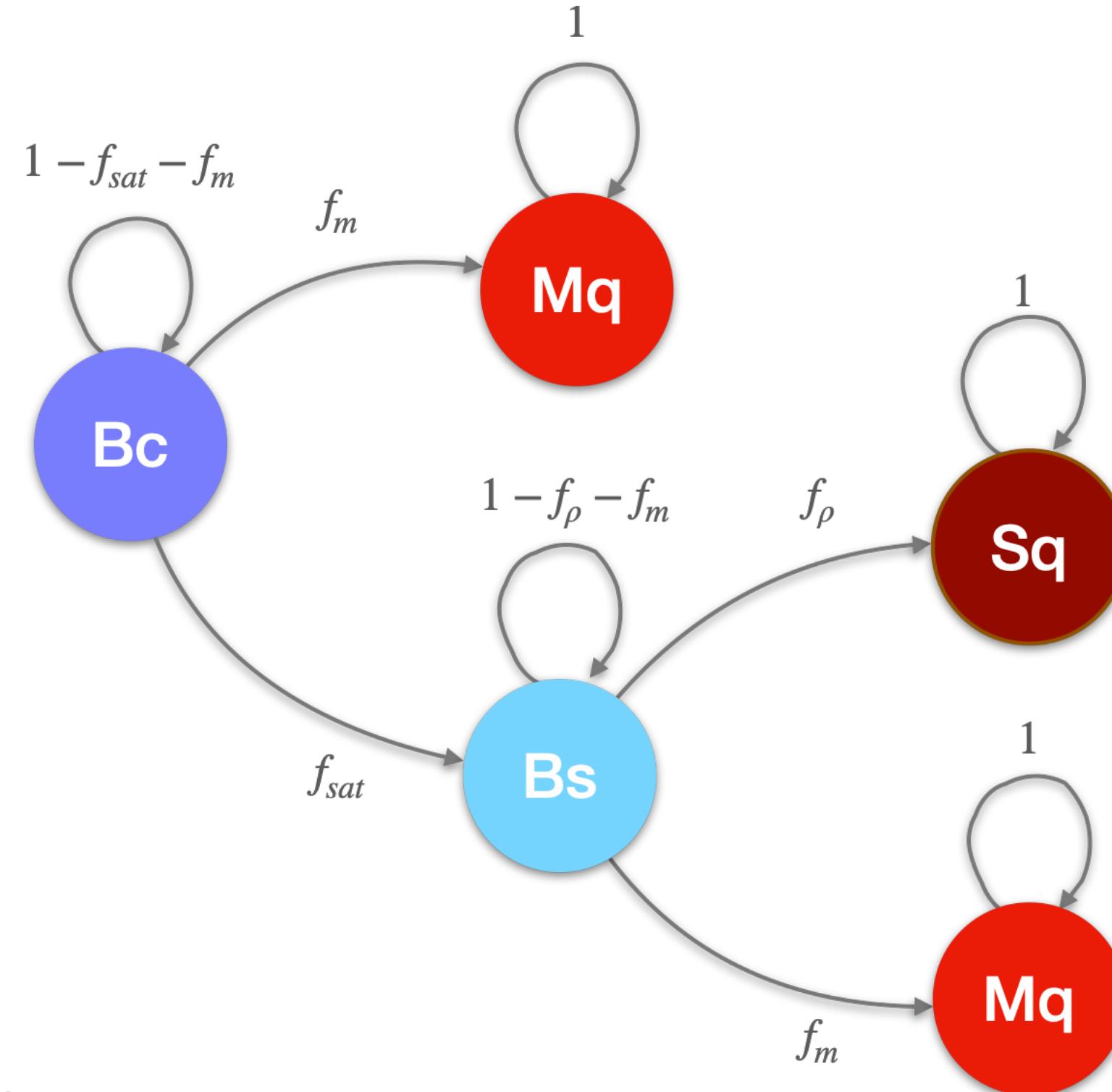
Produce realistic simulations faster and more accurately

Objectives

1. Write the “passive” **Schechter parameters** in terms of the active population.
2. Derive the **time evolution** of $\phi_*(z)$.

The method

Describe galaxy demographics with a set of continuity equations
invoking two quenching mechanisms



B_c Active central galaxy

B_s Active satellite galaxy

M_q Mass-quenched galaxy

S_q Satellite-quenched galaxy

$$\frac{dB_c}{dt} = \alpha sSFR B_c - \eta_m B_c - \eta_{sat} B_c$$

$$\frac{dB_s}{dt} = \alpha sSFR B_s - \eta_m B_s - \eta_\rho B_s + \eta_{sat} B_c$$

$$\frac{dM_q}{dt} = \eta_m B_c + \eta_m B_s$$

$$\frac{dS_q}{dt} = \eta_\rho B_s$$

The results

1. Write the “passive” **Schechter parameters** in terms of the active population.
2. Derive the **time evolution** of $\phi_*(z)$.

Reduce parameter space for simulations

Passive galaxies described by double Schechter function

Satellite-quenched galaxies subset of the active population

Exact analytical time dependence of the amplitude
of the active galaxies

	ϕ_*	α	M_*
Star-f	ϕ_b^*	α_b	m_b^*
Sat-q	$F_\rho \phi_{bs}^*$	α_b	m_b^*
Mass-q	$\simeq \phi_b^*$	$\alpha_b + 1$	m_b^*

$$\phi_b(m, t) = B_c(m, t) + B_s(m, t)$$

$$\phi_{*b}(z) = A(\eta_{sat}) e^{f(z; \eta_{sat})}$$

$$f_\rho = 1 - e^{-\int \eta_\rho dt} \quad F_\rho = \frac{1}{\ln(1 - f_\rho)}$$

$$\begin{Bmatrix} \phi_{*b} & \alpha_b & m_{*b} \\ \phi_{*m} & \alpha_m & m_{*m} \\ \phi_{*\rho} & \alpha_\rho & m_{*\rho} \end{Bmatrix} \rightarrow \begin{Bmatrix} \phi_{*b} & \alpha_b & m_{*b} \\ f_\rho & f_{sat} & \end{Bmatrix}$$

Note:
perform a polynomial expansion and
retrieve the Herbel et al. model 2017

$$\phi_*(z) = b e^{az}$$

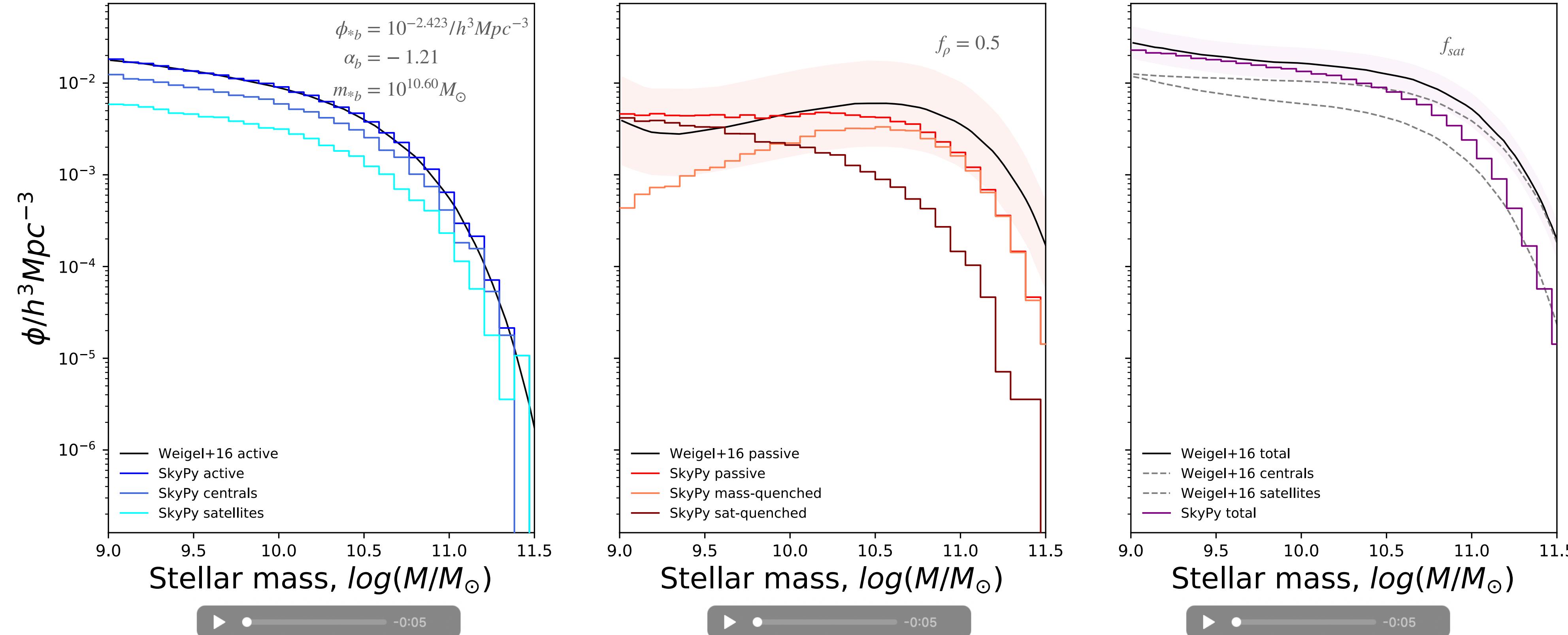


Implemented in the **SkyPy** galaxy module

Validation

Weigel et al. 2016 best-fit for SDSS DR7

Galaxy Demographics Simulation



Example + Sonification with Strauss

[Trayford J., 2021, james-trayford/strauss: v0.1.0](#)

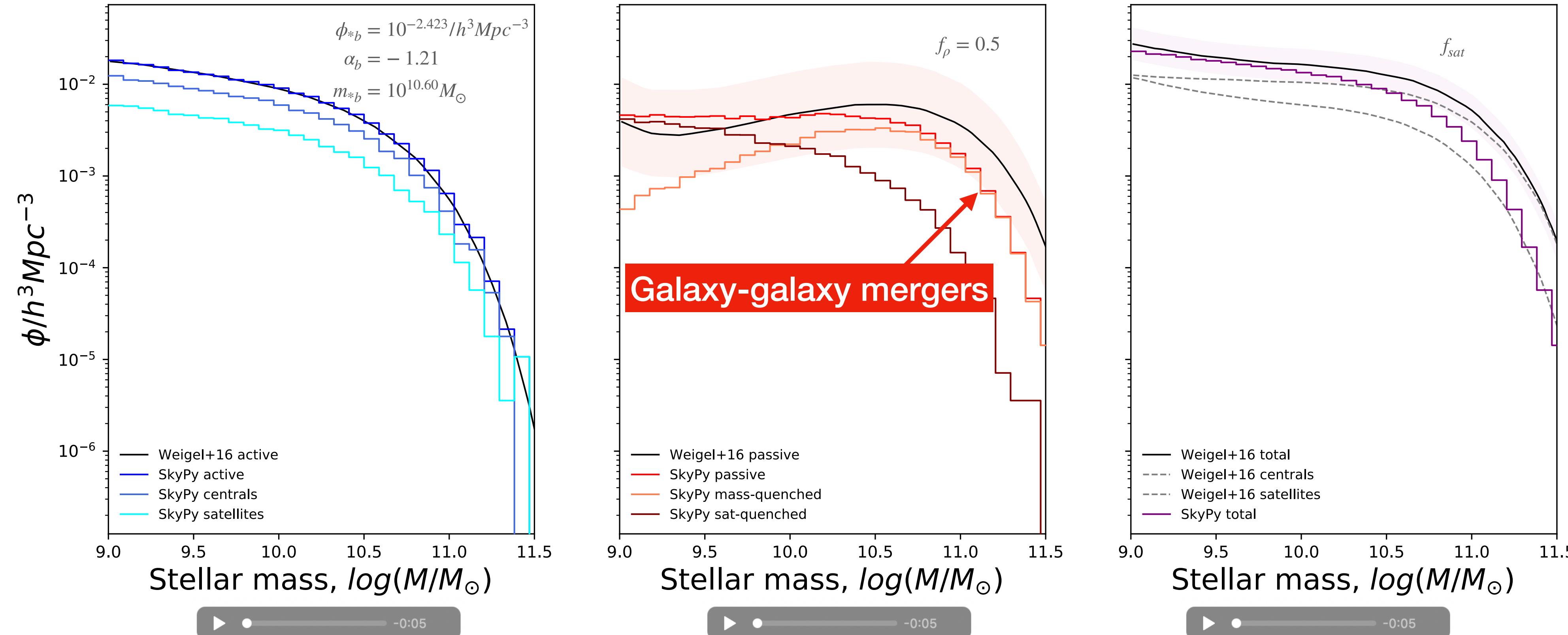


Implemented in the **SkyPy** example page

Validation

Weigel et al. 2016 best-fit for SDSS DR7

Galaxy Demographics Simulation



Example + Sonification with Strauss

[Trayford J., 2021, james-trayford/strauss: v0.1.0](#)



Implemented in the **SkyPy** example page

Remember

Main goal

Produce realistic simulations faster and more accurately

Method

Describe galaxy demographics with a set of continuity equations
invoking two **quenching mechanisms**

Results

1. Write the “passive” **Schechter parameters** in terms of the active population.
2. Derive the **time evolution** of $\phi_*(z)$.

Conclusions

Reduce parameter space for simulations

Passive galaxies described by double Schechter function

Satellite-quenched galaxies subset of the active population

Exact analytical time dependence of the amplitude
of the active galaxies

Validation against Weigel et al. 2016 best-fit for SDSS DR7
 $f_\rho \sim 50\%$ in agreement with literature values



Implemented in the **SkyPy** galaxy module

Implemented in the **SkyPy** example page

Sonification with Strauss



Questions



II. and the SkyPy project



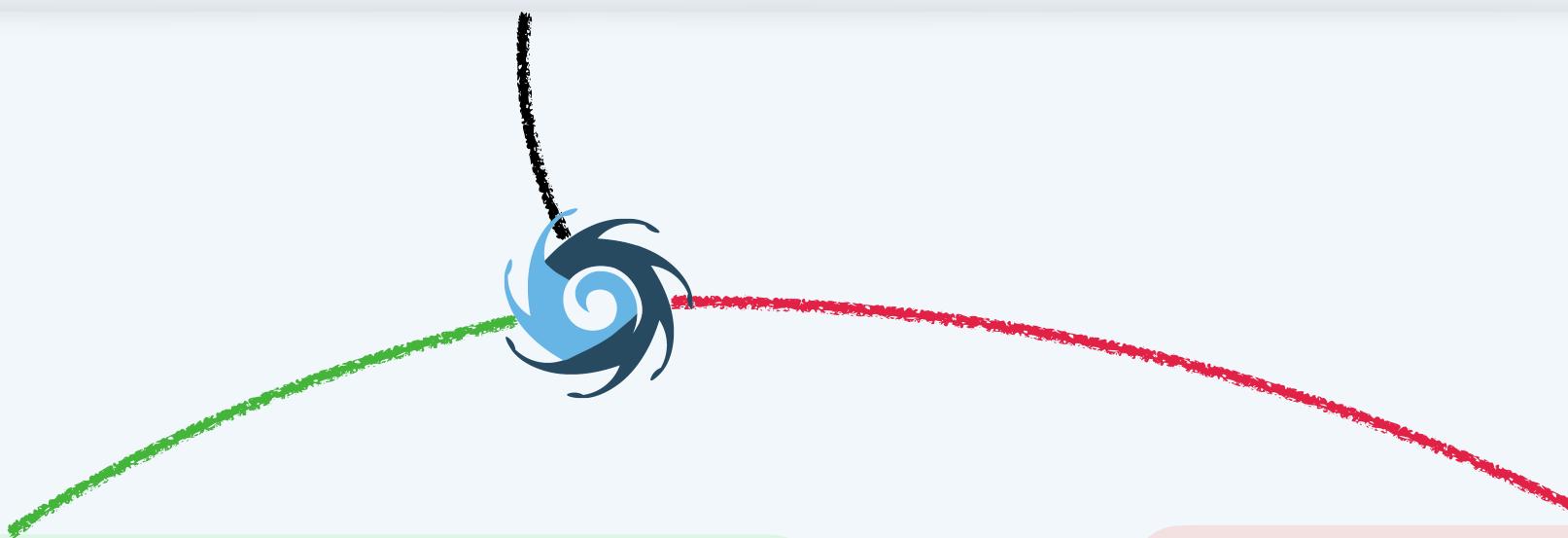
- 7. CONFIGURATION FILES**
- 6. SIMULATION PIPELINES**
- 5. RESEARCH & DEVELOPMENT**
- 4. THE LIBRARY**

- 1. VISION**
- 2. MEMBERS**
- 3. COMMUNITY PACKAGE**



1. The vision

- Observational cosmology limited by data access
- Open-data revolution in astronomy
- Challenge: access to sophisticated analysis **methods**.
- Emerging methods: forward modelling & machine learning.



- III generation of catalog production (**user-generated outputs**)
- **Open-source** off-project high-quality **Python** package
- **End-to-end simulations** of the astrophysical sky
- Interface with external software
- Enable **Forward Modelling** and **Machine Learning**

- Not a single pipeline simulation
 - Do not replicate existing code
- **Reuse**
 - Astropy-affiliated packages
 - High-quality codes
 - **Ecosystem** of compatible software



2. Members

Sarah Bridle

Juan Pablo Cordero

Ian Harrison

Laura Wolz



THE UNIVERSITY
of EDINBURGH



Brian Nord

Simon Birrer



Stanford
University

Richard Rollins

Nicolas Tessore

Adam Amara

Lucia F. de la Bella

Philipp Sudek

Ginevra Favole

Arthur Tolley



Coleman Krawczyk

Ian Harry

Laura Nutall

Andrew Lundgren

Andrew Williamson



Keiichi Umetsu

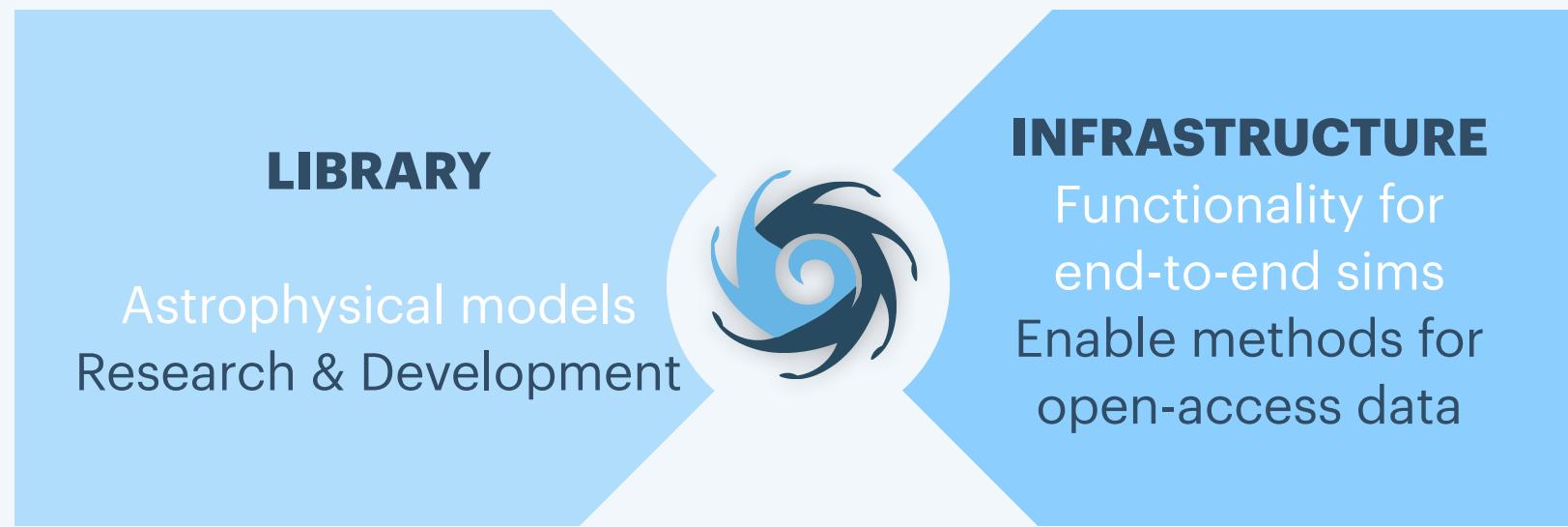
Sut-ieng Tam

William Hartley



3. Community Package

<https://github.com/skypyproject/skypy.git>



- GitHub organisation
- Unit tests & high-quality documentation
- Code review & Infrastructure team

```
my-pc: -$ pip install skypy or
my-pc: -$ conda install -c conda-forge skypy or
my-pc: -$ git clone https://github.com/skypyproject/skypy.git

my-pc: -$ ipython
...
[1]: import skypy
```

skypy:docs
skypy v0.5.dev24+gb377ea0 »

Page Contents

SkyPy Documentation

- [Getting Started](#)
- [User Documentation](#)
 - [Packages](#)
 - [Pipeline](#)
- [Developer Documentation](#)
- [Project details](#)
- [Index](#)
- [Acknowledgements](#)

SkyPy Documentation

This package contains methods for modelling the Universe, galaxies and the Milky Way. Also included are methods for generating observed data.

Getting Started

- [Installation](#)
- [Feature List](#)
- [Configuration Files](#)
- [Examples](#)

User Documentation

Packages

- [Galaxies \(`skypy.galaxies`\)](#)
- [Utils \(`skypy.utils`\)](#)

Pipeline

- [Pipeline \(`skypy.pipeline`\)](#)

Developer Documentation

- [Contributor Guidelines](#)

Project details

- [Code of Conduct](#)

<https://skypy.readthedocs.io/en/latest>



4. The Library

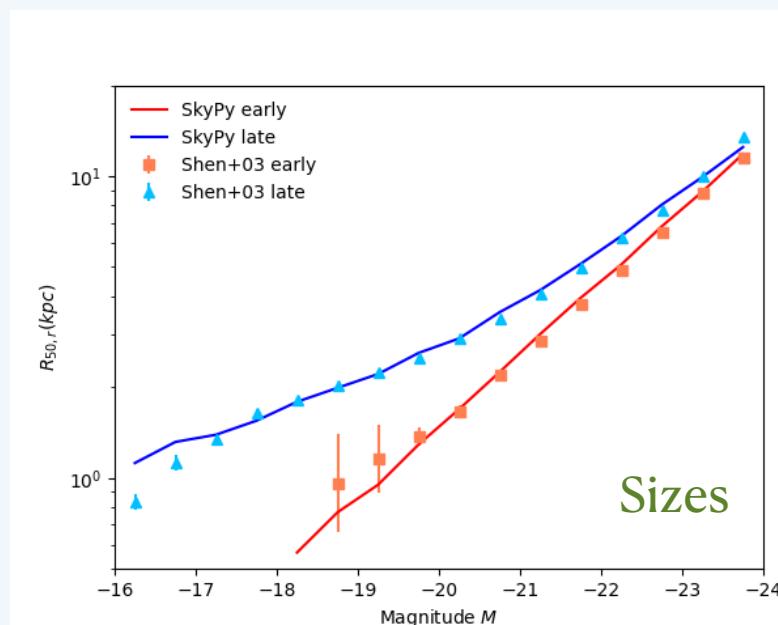
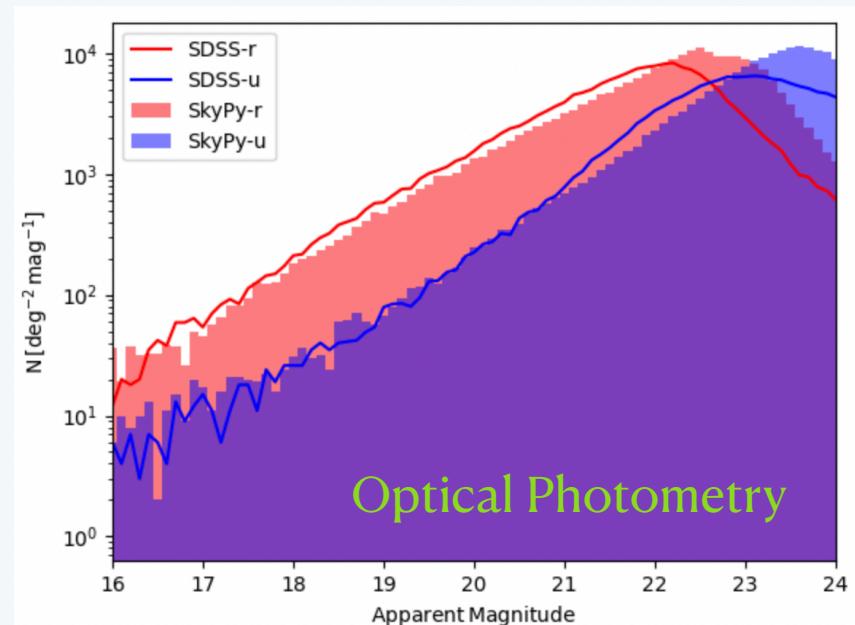
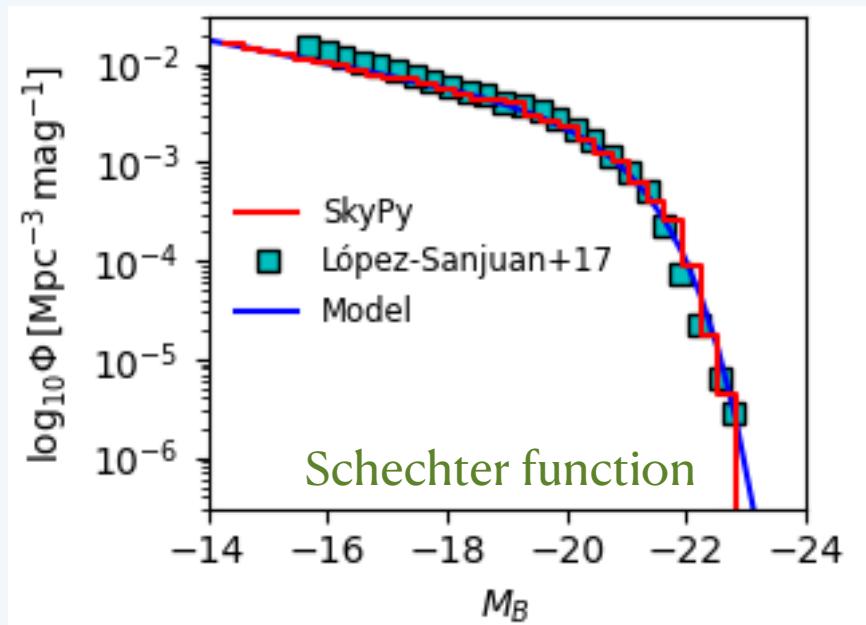




5. Research & Development

SkyPy v0.5

- **Luminosity Distributions** — Schechter Luminosity function
- **Morphological Distributions** — angular size, (early- and late-type) linear lognormal size distribution, beta ellipticity and Ryden 2004 ellipticity distributions.
- **Redshift Distributions** — redshifts from co-moving density, Schechter (luminosity and stellar mass) redshift distribution, Smail+94 redshift distribution.
- **Spectral Energy Distribution Modelling** — Dirichlet coefficients, Correct templates.
- **Stellar Mass Distribution** — Schechter stellar mass function.



SkyPy v0.6

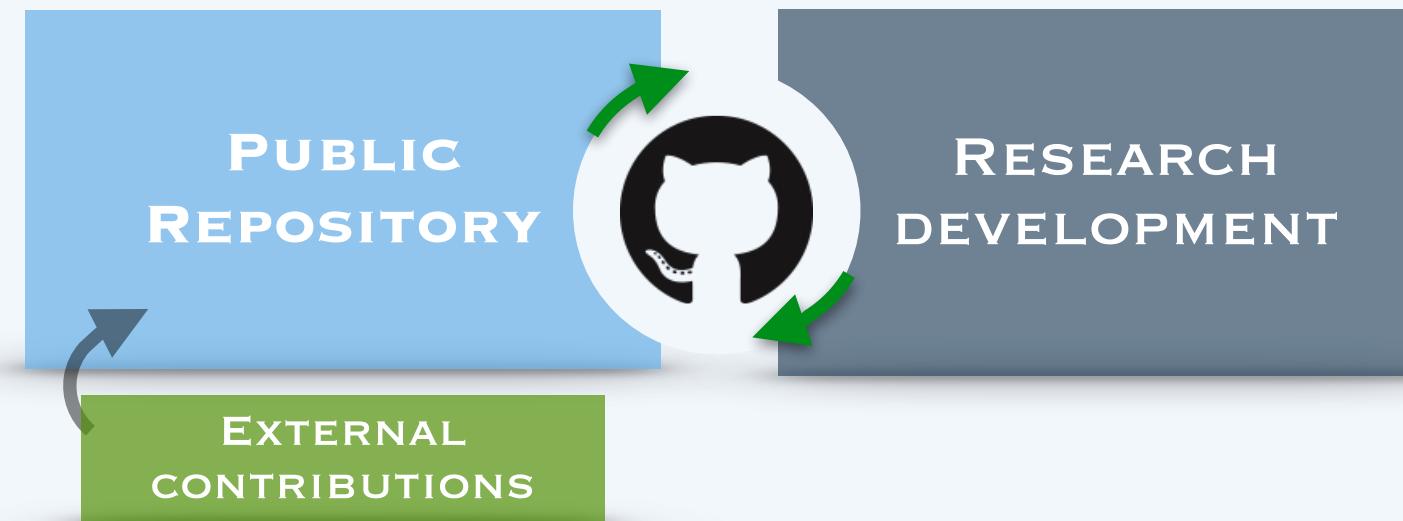
- **Power Spectrum** — CAMB, Halofit, CLASS, Eisenstein & Hu, growth functions
- **Dark Matter Halos** — Colossus, halo and sub-halo mass sampler, ellipsoidal and spherical collapse functions (Press-Schechter, Sheth-Tormen), abundance matching, quenching models



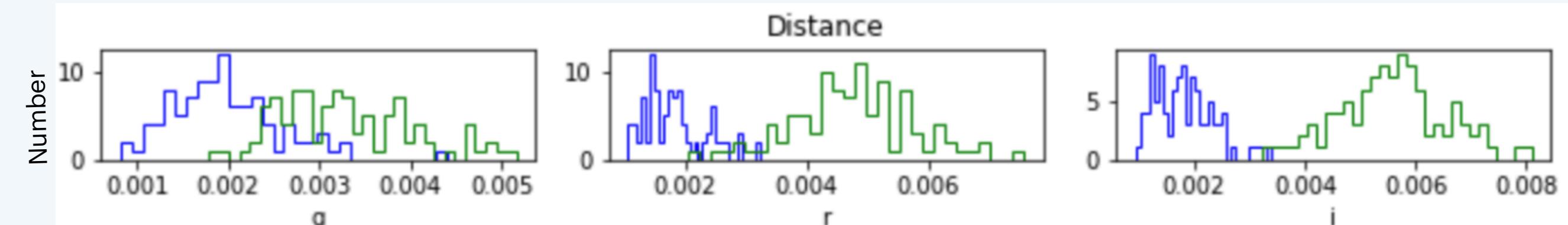
5. Research & Development

Key

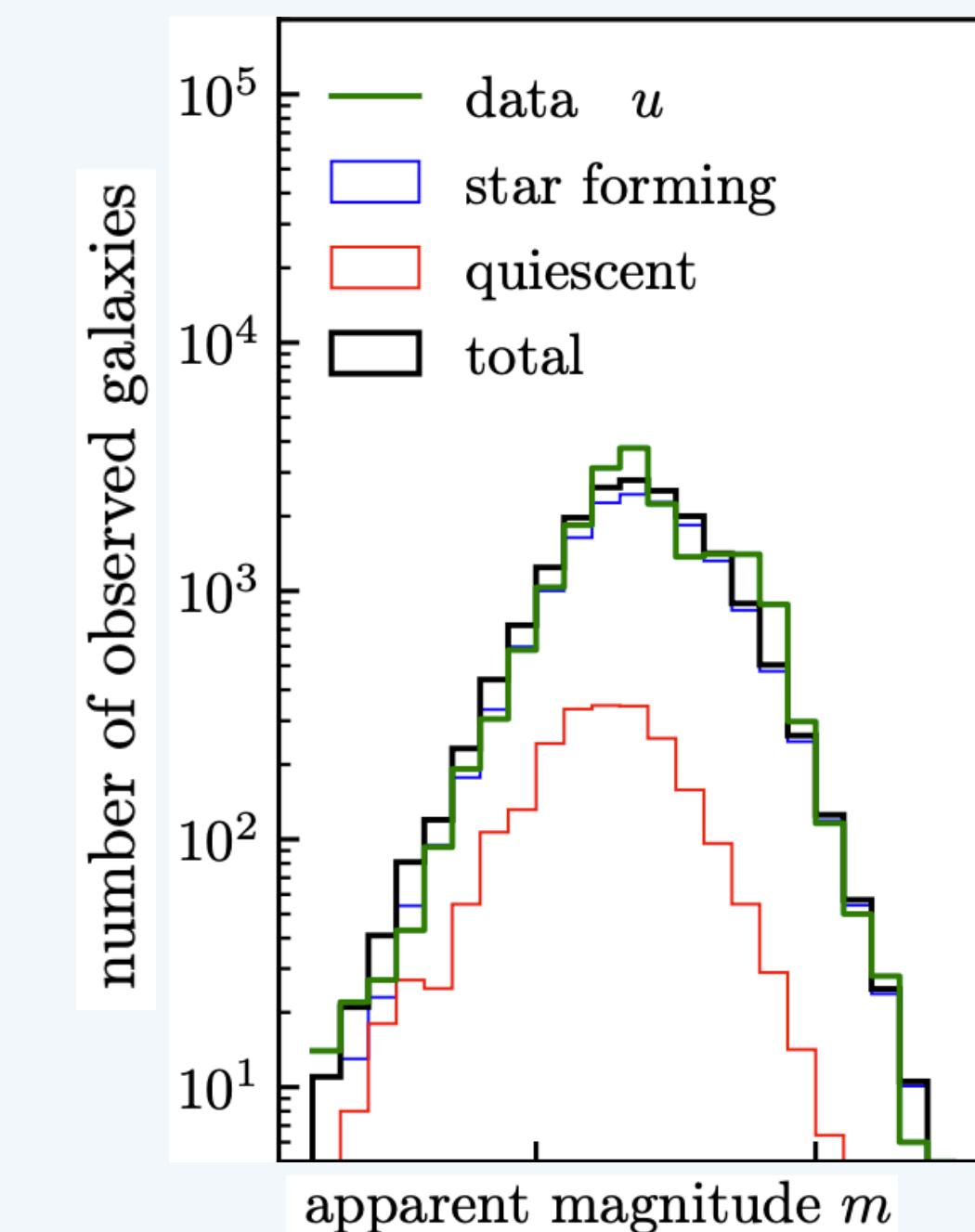
SkyPy is driven by science projects



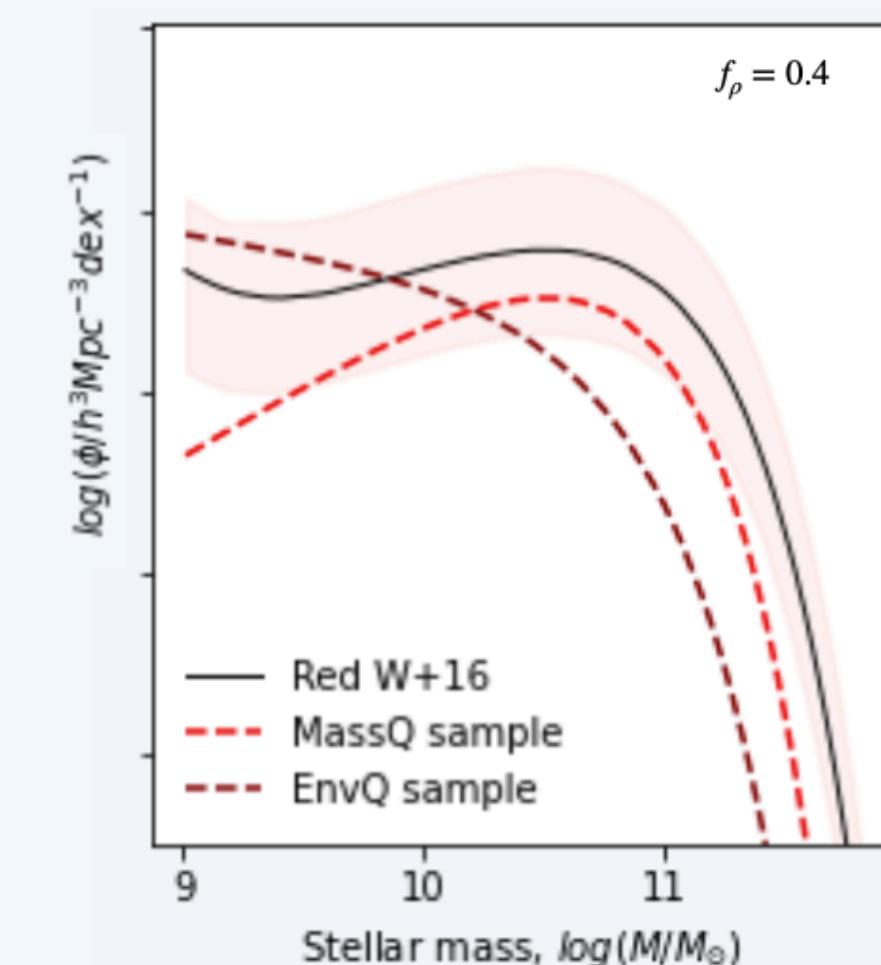
- **Likelihood-Free Inference for Cluster Weak Lensing** - Sut-leng Tam, ASIAA
- **Projected Galaxy Ellipticities** - Juan Pablo Cordero, Univ. of Manchester
- **Galaxy Demographics** - Lucia F. de la Bella, Univ. of Portsmouth
- **Forecasting Optical Galaxy Surveys** - Philipp Sudek, Univ. of Portsmouth
- **Galaxies** - William Hartley
- **Gravitational Wave Binary Merger Populations** - Arthur Tolley, Univ. of Portsmouth



Investigation of Schechter parameter sensitivity of a DES-like survey (Sudek+ in prep).
Big difference of the green and blue histogram indicates high constraining power using the corresponding observable



Apparent magnitude distribution in the SDSS u filter simulated with SkyPy (blue, red, black) compared to SDSS data (Tessore+ in prep.)



Schechter function in the quenching model vs SDSS best fit (de la Bella + in prep.)



6. Simulation pipelines

Key

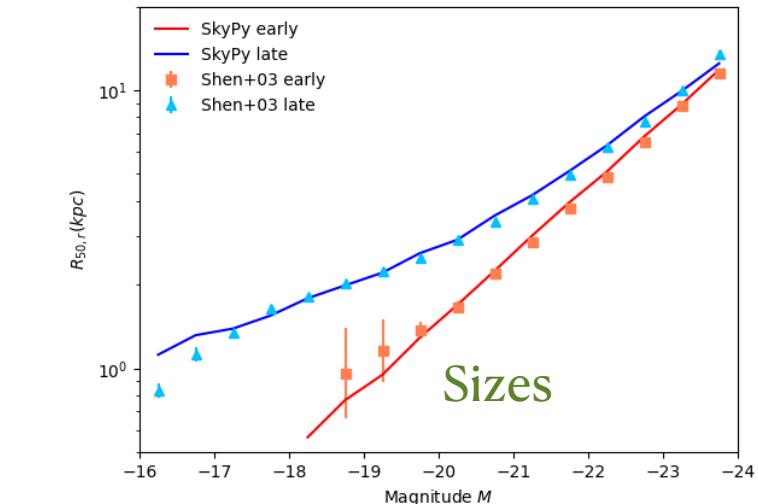
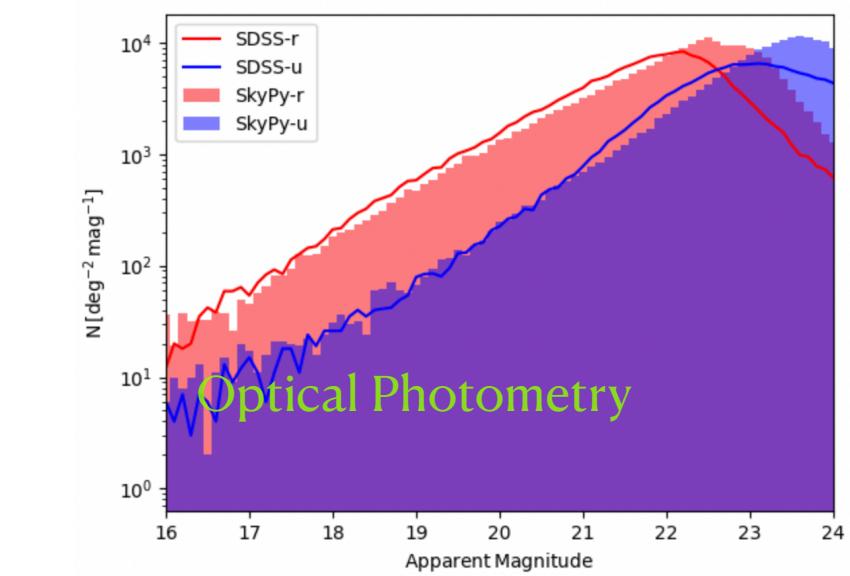
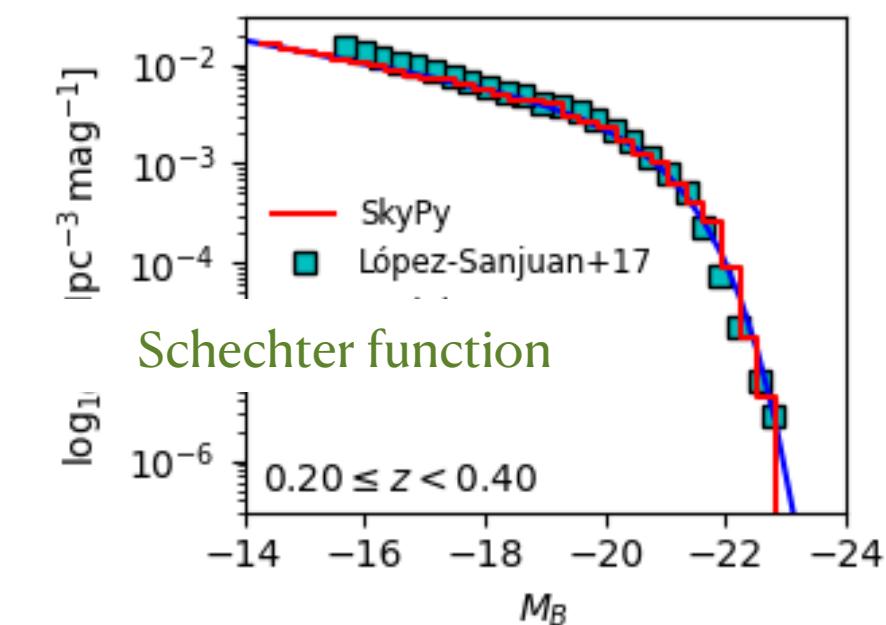
- YAML-based config files
- The **SkyPy Driver** runs end-to-end **pipelines**
- **Total flexibility!**

- **SkyPy Pipeline**
- **KEY:** you can write your own **pipelines!**

Combine SkyPy with your favourite software!



SkyPy PIPELINES





7. Configuration files

SkyPy Syntax

- **Variables** — Astropy quantities, import objects
- **Parameters** — variables modified at execution
- **Functions** — cosmology, job completion
- **Tables** — multicolumn assignment, table reference

Example: luminosity.yml

```
cosmology: !astropy.cosmology.default_cosmology.get []
z_range: !numpy.linspace [0, 2, 21]
M_star: !astropy.modeling.models.Linear1D [-0.9, -20.4]
phi_star: !astropy.modeling.models.Exponential1D [3e-3, -9.7]
magnitude_limit: 23
sky_area: 0.1 deg2
tables:
    blue_galaxies:
        redshift, magnitude: !skypy.galaxies.schechter_lf
        redshift: $z_range
        M_star: $M_star
        phi_star: $phi_star
        alpha: -1.3
        m_lim: $magnitude_limit
        sky_area: $sky_area
```

```
import matplotlib.pyplot as plt
from skypy.pipeline import Pipeline

# Execute SkyPy luminosity pipeline
pipeline = Pipeline.read("luminosity.yml")
pipeline.execute()

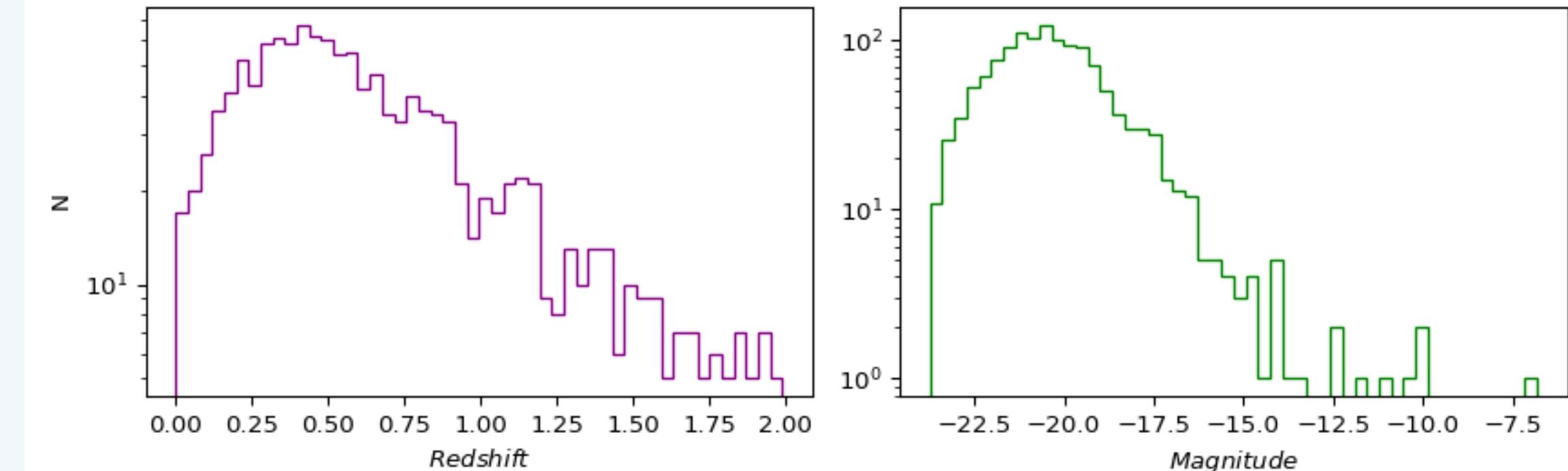
# Blue population
skypy_galaxies = pipeline['blue_galaxies']

# Plot histograms
fig, axs = plt.subplots(1, 2, figsize=(9, 3))

axs[0].hist(skypy_galaxies['redshift'], bins=50, histtype='step', color='purple')
axs[0].set_xlabel(r'$Redshift$')
axs[0].set_ylabel(r'$\mathcal{N}$')
axs[0].set_yscale('log')

axs[1].hist(skypy_galaxies['magnitude'], bins=50, histtype='step', color='green')
axs[1].set_xlabel(r'$Magnitude$')
axs[1].set_yscale('log')

plt.tight_layout()
plt.show()
```



You can also run the pipeline directly from the command line and write the outputs to a fits file:

```
$ skypy luminosity.yml luminosity.fits
```



7. Configuration files

SkyPy Syntax

- **Variables** — Astropy quantities, import objects
- **Parameters** — variables modified at execution
- **Functions** — cosmology, job completion
- **Tables** — multicolumn assignment, table reference

https://skypy.readthedocs.io/en/latest/configuration_files.html

Example: luminosity.yml

```
cosmology: !astropy.cosmology.default_cosmology.get []
z_range: !numpy.linspace [0, 2, 21]
M_star: !astropy.modeling.models.Linear1D [-0.9, -20.4]
phi_star: !astropy.modeling.models.Exponential1D [3e-3, -9.7]
magnitude_limit: 23
sky_area: 0.1 deg2
tables:
  blue_galaxies:
    redshift, magnitude: !skypy.galaxies.schechter_lf
    redshift: $z_range
    M_star: $M_star
    phi_star: $phi_star
    alpha: -1.3
    m_lim: $magnitude_limit
    sky_area: $sky_area
```

```
import matplotlib.pyplot as plt
from skypy.pipeline import Pipeline

# Execute SkyPy luminosity pipeline
pipeline = Pipeline.read("luminosity.yml")
pipeline.execute()

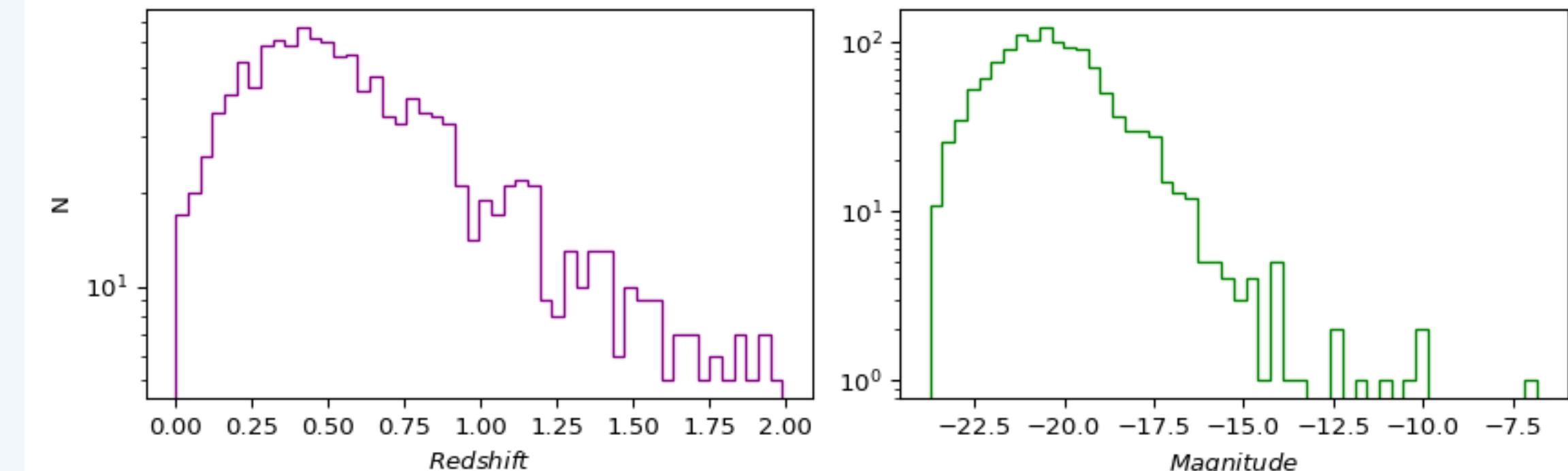
# Blue population
skypy_galaxies = pipeline['blue_galaxies']

# Plot histograms
fig, axs = plt.subplots(1, 2, figsize=(9, 3))

axs[0].hist(skypy_galaxies['redshift'], bins=50, histtype='step', color='purple')
axs[0].set_xlabel(r'$Redshift$')
axs[0].set_ylabel(r'$\mathcal{N}$')
axs[0].set_yscale('log')

axs[1].hist(skypy_galaxies['magnitude'], bins=50, histtype='step', color='green')
axs[1].set_xlabel(r'$Magnitude$')
axs[1].set_ylabel(r'$\mathcal{N}$')
```

```
plt.tight_layout()
plt.show()
```



You can also run the pipeline directly from the command line and write the outputs to a fits file:

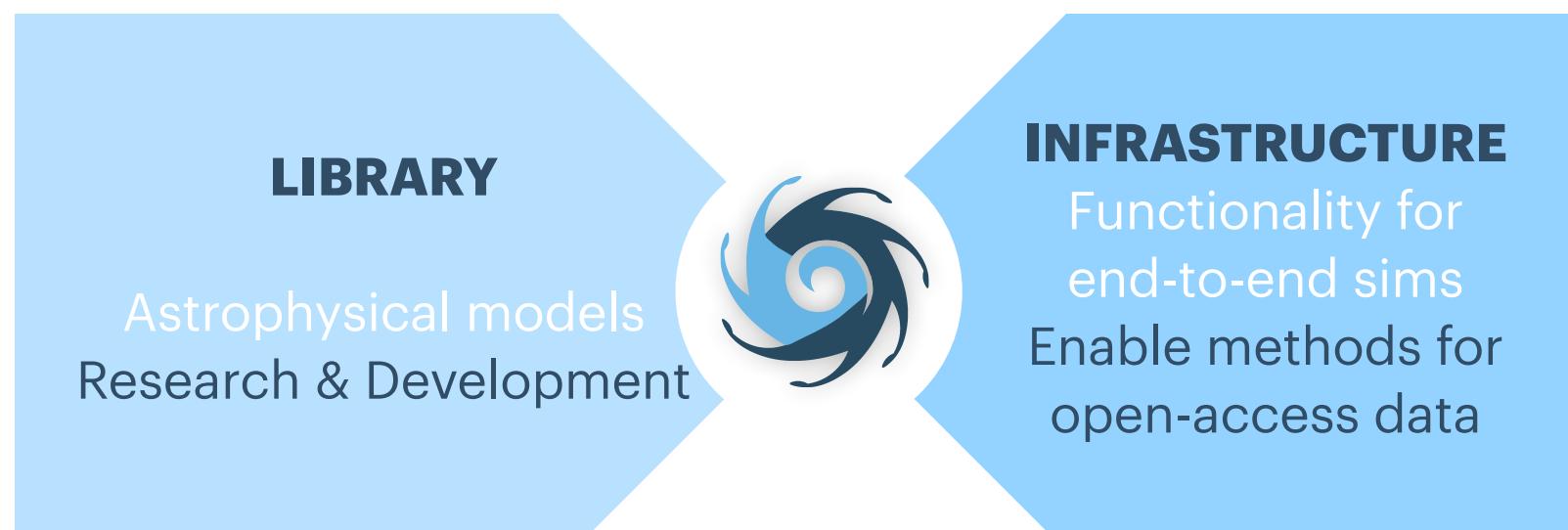
```
$ skypy luminosity.yml luminosity.fits
```



Summary

COMMUNITY PACKAGE

- Open-source off-project
- High-quality **Python** package

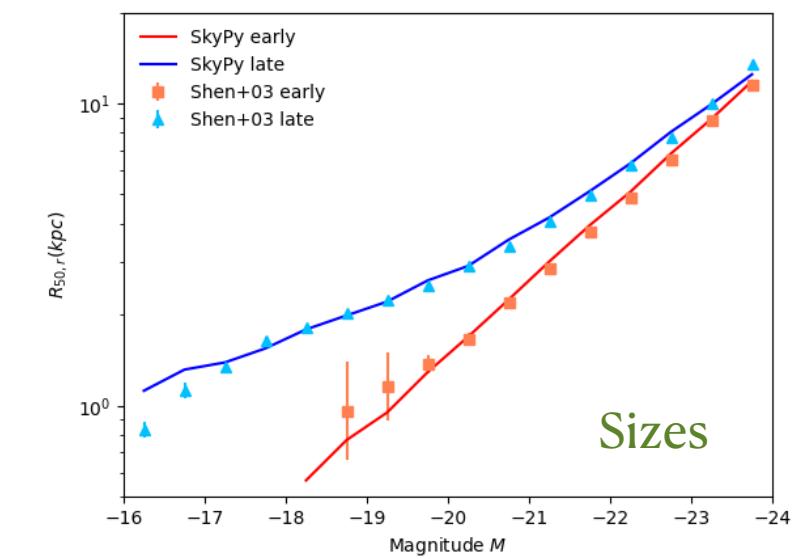
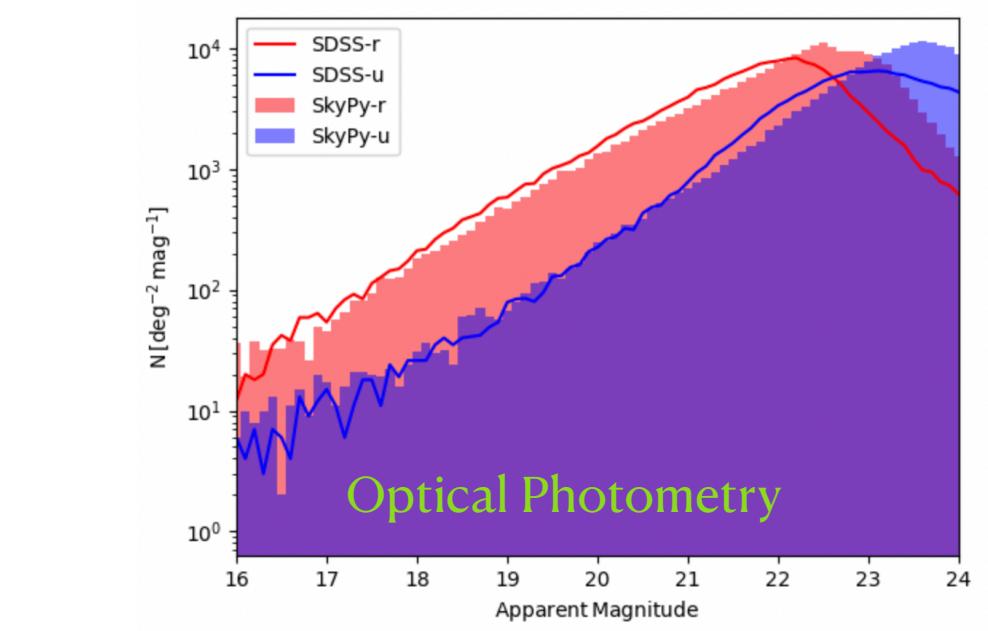
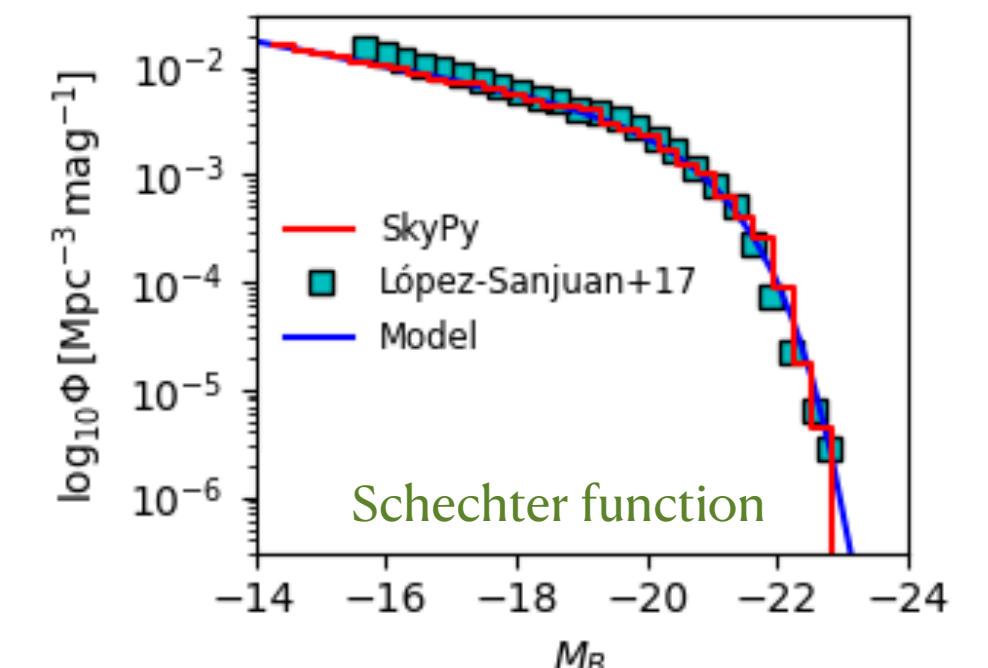
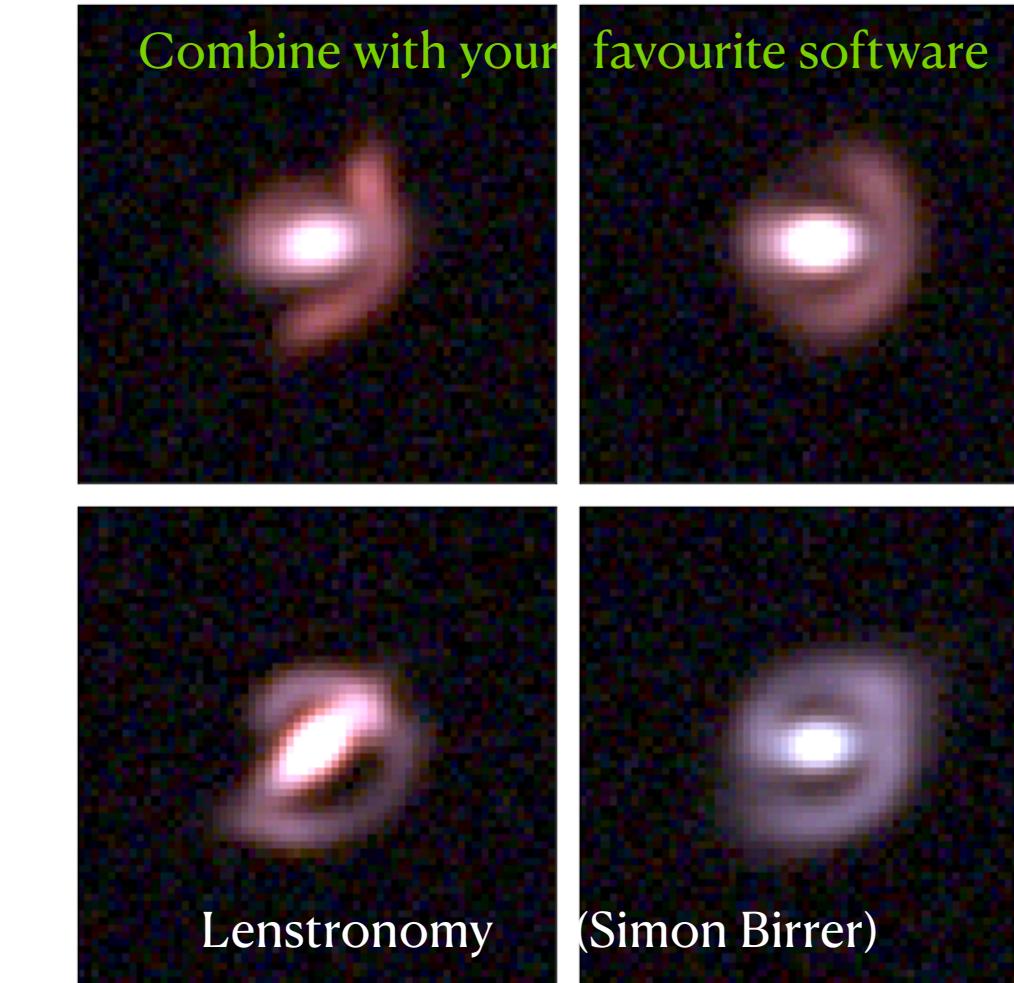


- GitHub organisation
 - Unit tests & high-quality documentation
 - Code review & Infrastructure team
- News**
- **vo.6** release: halo & power spectrum.
 - Journal of Open-Source Software

SIMULATION PIPELINES

- YAML-based config files
- The **SkyPy Driver** runs end-to-end **pipelines**
- **Total flexibility!**

- SkyPy Pipeline
- **KEY:** you can write your own **pipelines!**



<https://github.com/skypyproject/skypy.git>

<https://skypyproject.org>

<https://skypy.readthedocs.io/en/latest/examples/index.html>

Open your terminal...

```
my-pc: -$ pip install skypy or  
my-pc: -$ conda install -c conda-forge skypy or  
my-pc: -$ git clone https://github.com/skypyproject/skypy.git  
  
my-pc: -$ ipython  
...  
[1]: import skypy
```



Questions?